

# Deep Generative Models for High Energy Physics

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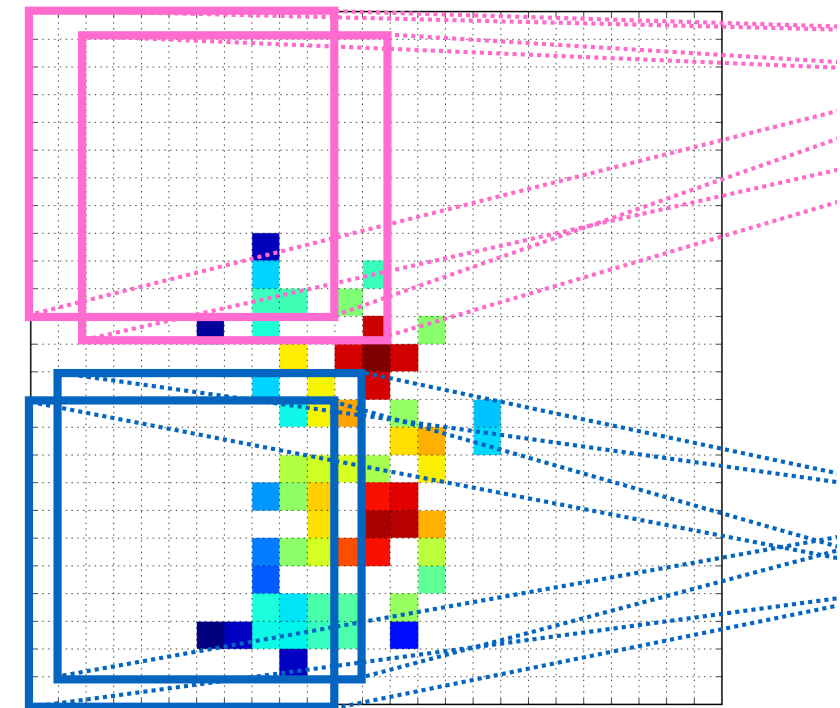
[bpnachman@lbl.gov](mailto:bpnachman@lbl.gov)



@bpnachman



bnachman



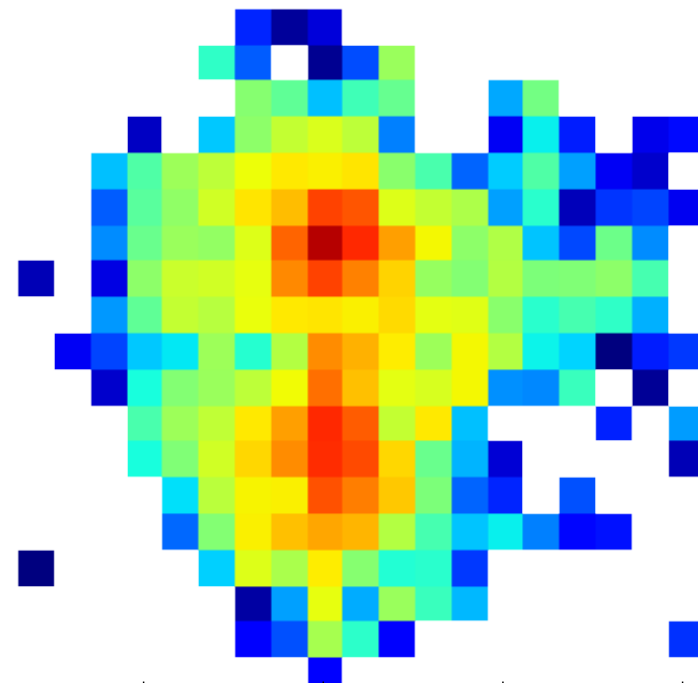
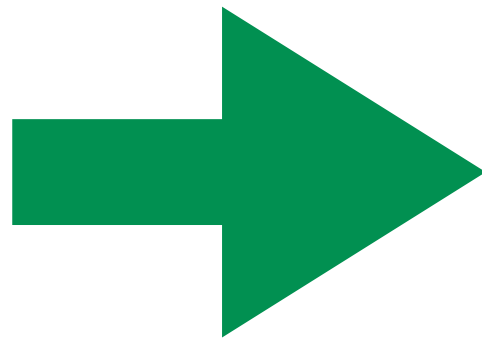
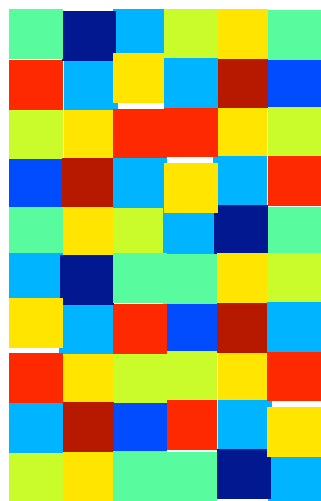
UCSD PHYS 139/239

March 7, 2023

# Introduction: generative models

2

A **generator** is nothing other than a function that maps random numbers to structure.



Deep generative models: the map is a deep neural network.

**GANs**

*Generative  
Adversarial Networks*

**Score-  
based**

**Restricted  
Boltzmann  
Machines**

**Mixture  
Density  
Networks**

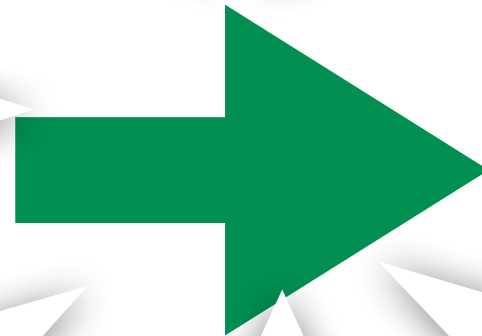
**NFs**

*Normalizing Flows*

**Energy-  
based  
models**

**VAEs**

*Variational Autoencoders*



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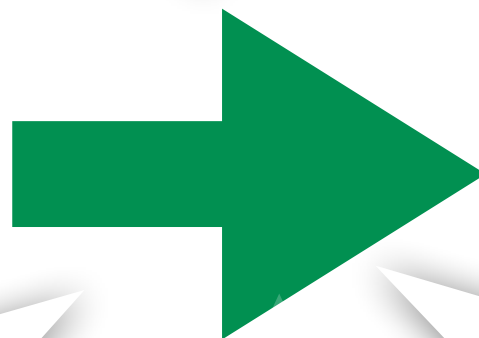
Energy-  
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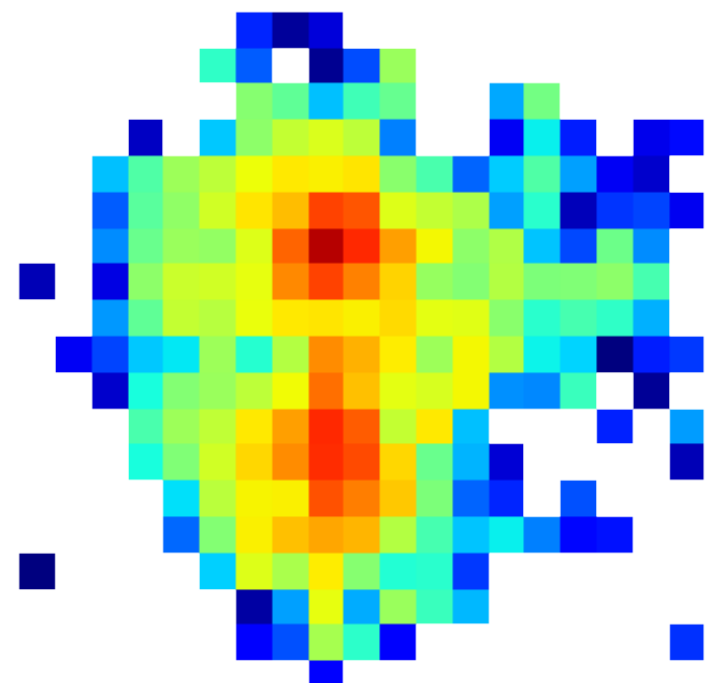
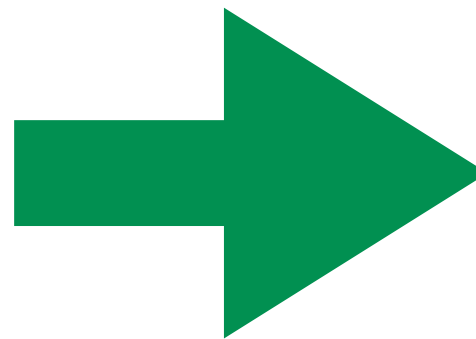
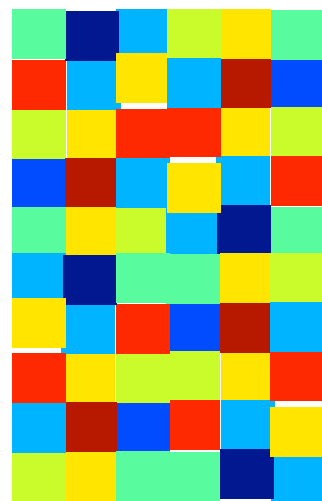
**VAEs**

*Variational Autoencoders*





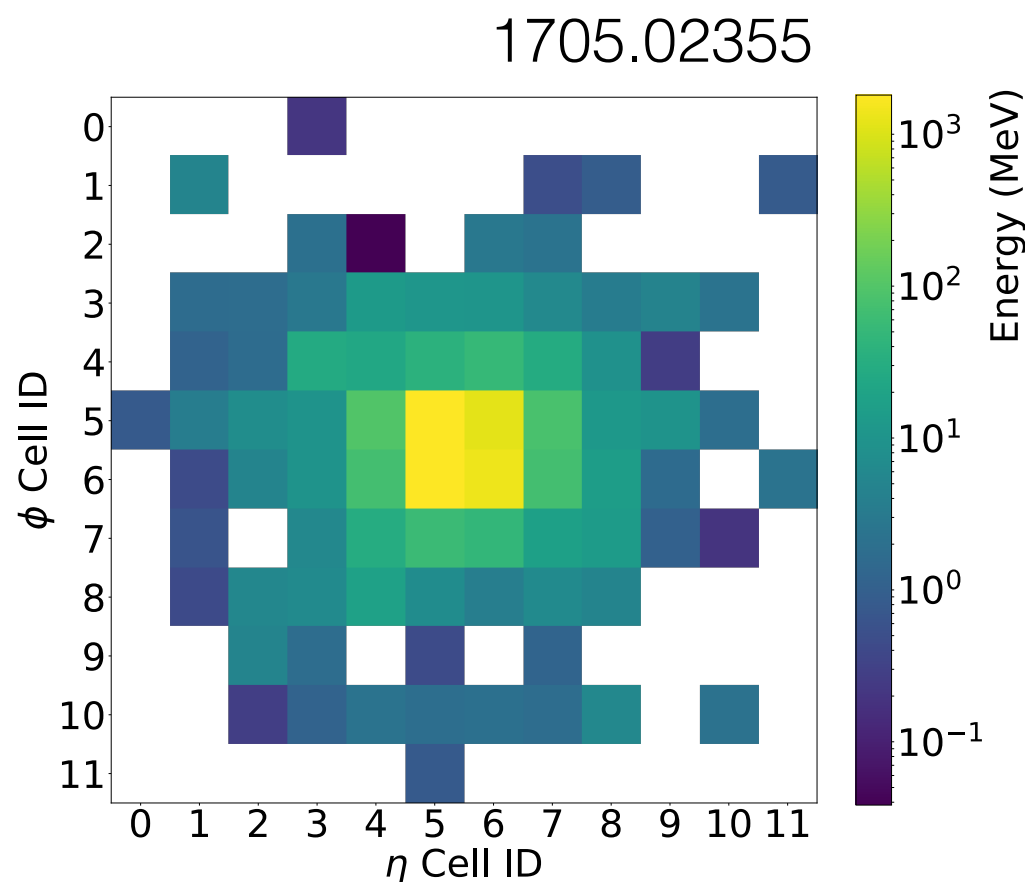
1. Why Generative Models?
2. Warm up: mixture density networks
3. Next: brief overview of GANs, VAEs, NFs
4. Statistical amplification
5. Applications
6. Bonus



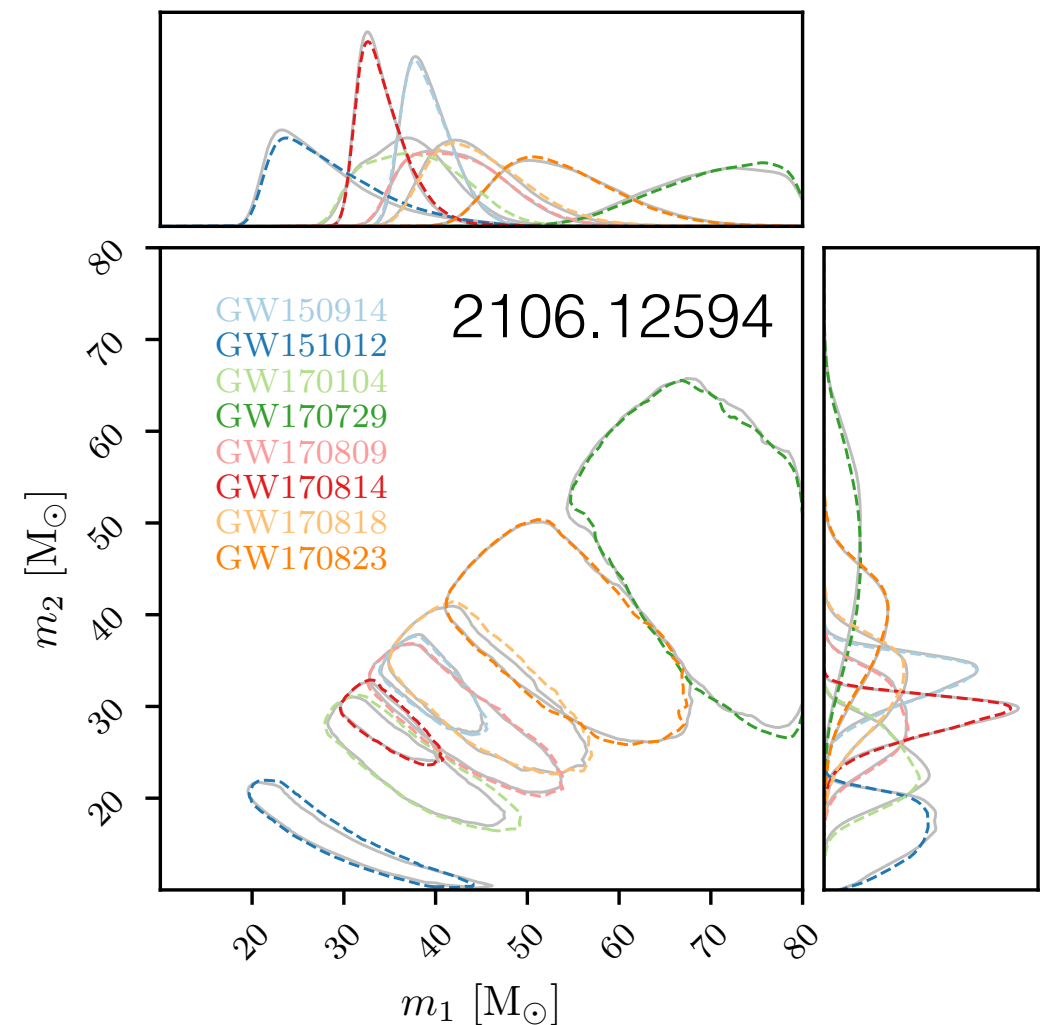
# Why generative models?



1. Augment/replace slow physics-based simulations
2. (Fast) Confidence Intervals / Posterior Analysis
3. Background estimation / anomaly detection
4. ....



Calorimeter sim.  $\times 10^5$  faster



GW inference  $\times 10^4$  faster

# Warm-up: MDNs



If you know the family of functions that generated your data, you can't beat fitting that function.

(I'll be more precise later about "beat")

A lot of physics data are nearly Gaussian, or combinations of a small number of Gaussians, possibly with a complex dependence on energy, time, etc.

Classical mixture model:

$$p(x) \propto \sum_i \alpha_i \exp \left( -\frac{1}{2} (x - \mu)^T \Sigma^{-1} (x - \mu) \right)$$

In case you have not seen this notation...

$$G(x | \mu, \sigma) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{1}{2\sigma^2}(x-\mu)^2}$$

In 1D,  $\Sigma = \sigma^2$ .

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$$p(x) \propto \sum_i \alpha_i \exp\left(-\frac{1}{2}(x - \mu)^T \Sigma^{-1}(x - \mu)\right)$$

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10

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Classical mixture model:

$$p(x) \propto \sum_i \alpha_i \exp \left( -\frac{1}{2} (x - \mu)^T \Sigma^{-1} (x - \mu) \right)$$

Mixture density network:

$$p(x | z) \propto \sum_i \alpha_i(z) \exp \left( -\frac{1}{2} (x - \mu(z))^T \Sigma(z)^{-1} (x - \mu(z)) \right)$$

Mixture density network:

$$p(x | z) \propto \sum_i \alpha_i(z) \exp \left( -\frac{1}{2} (x - \mu(z))^T \Sigma(z)^{-1} (x - \mu(z)) \right)$$

The “parameters”  $\alpha(z)$ ,  $\mu(z)$ ,  $\Sigma(z)$  are deep neural networks.

Since  $\Sigma$  is symmetric and positive semi-definite, need to take care when designing its architecture.

Train an MDN by minimizing  $-\log(p(x | z))$ .

When  $x$  is 1D, there is only one component, and  $\Sigma$  is fixed, then this reduces to mean squared error!

# MDNs: code

12

```
from tensorflow import keras
import tensorflow_probability as tfp

tfd = tfp.distributions
tfpl = tfp.layers
tfk = tf.keras
tfkl = tf.keras.layers

n = len(X[:,0:])
z = np.zeros(n)
x = X[:,0:]

# Model the distribution of y given x with a Mixture Density Network.
event_shape = np.shape(X)
num_components = 5
params_size =
tfpl.MixtureSameFamily.params_size(num_components,component_params_size=tfpl.MultivariateNormalTriL.params_size(event_shape[0]))
model = tfk.Sequential([
    tfkl.Dense(32, activation='relu'),
    tfkl.Dense(64, activation='relu'),
    tfkl.Dense(64, activation='relu'),
    tfkl.Dense(params_size, activation=None),
    tfpl.MixtureSameFamily(num_components, tfpl.MultivariateNormalTriL(event_shape[0]))
])

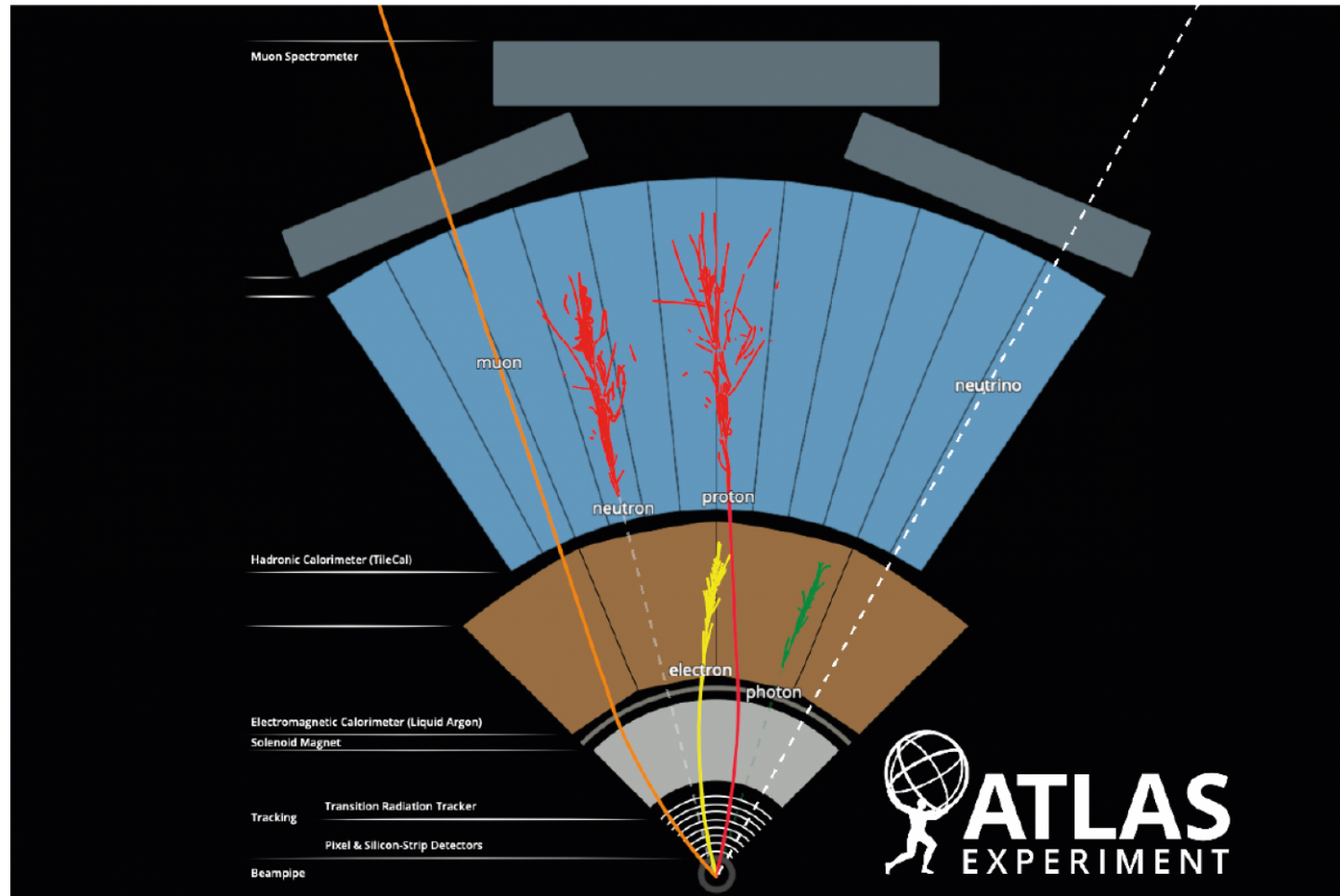
# Fit.
batch_size = 100
model.compile(optimizer='adam',
              loss=lambda x, model: -model.log_prob(x))
myhistory = model.fit(z, x,
                    batch_size=batch_size,
                    epochs=50,
                    steps_per_epoch=n // batch_size)
```

(equivalently simple in PyTorch)



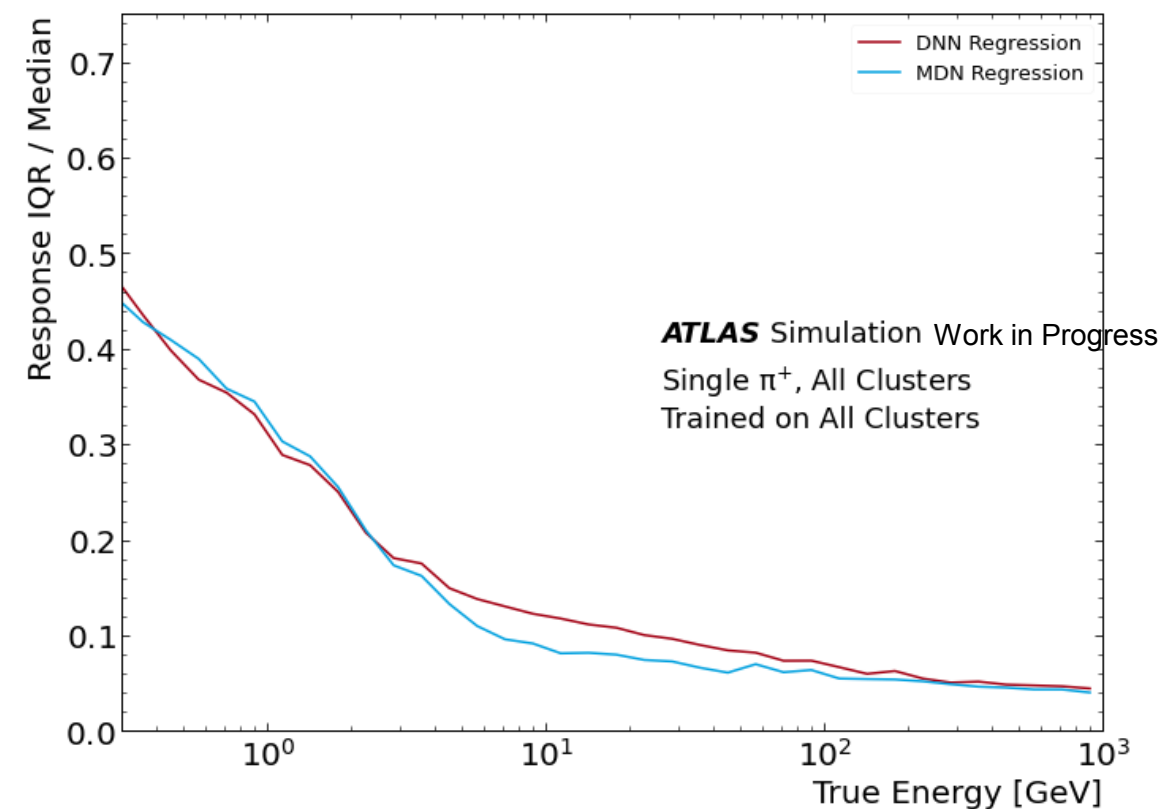
# MDNs in action

13



Example: energy regression for hadrons in the ATLAS calorimeter

(figures without citations are links)



Simultaneously learn confidence interval (“resolution”) and point estimate

## Pros:

- Access to the density
- Fast to sample
- Easy to specify
- Easy to train

## Cons:

- Do not scale well to high dimensions
- Not good when the data are not nearly Gaussian(s)

Up next: deep generative models with more flexibility

**GANs**

*Generative  
Adversarial Networks*

Restricted  
Boltzmann  
Machines

**NFs**

*Normalizing Flows*

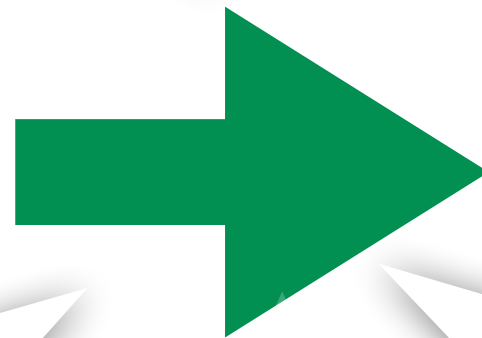
Energy-  
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models

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*Variational Autoencoders*

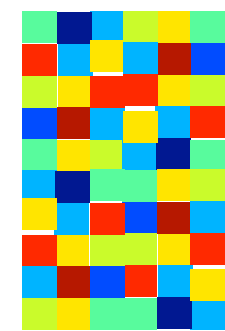


# Introduction: GANs

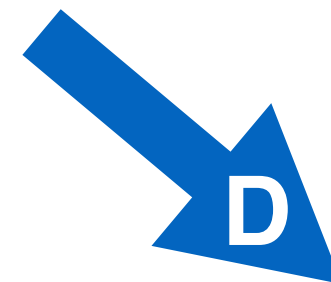
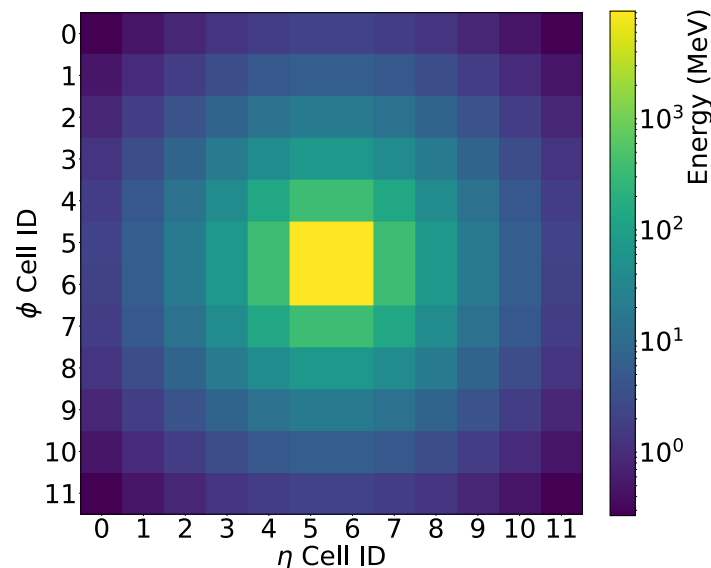
16

Generative Adversarial Networks (GANs):

*A two-network game where one **maps noise to structure** and one **classifies images as fake or real**.*

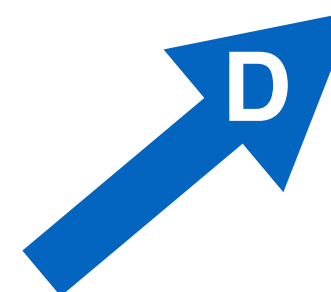
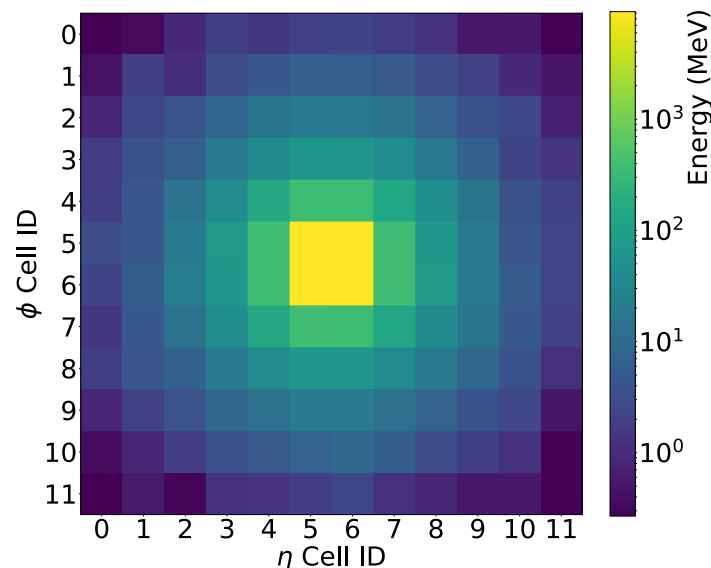


noise



{real, fake}

When **D** is maximally confused, **G** will be a good generator



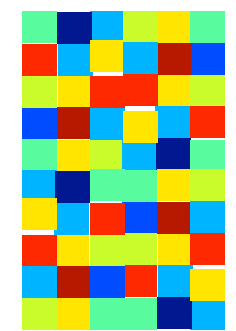
Physics-based simulator or data

# Introduction: GANs

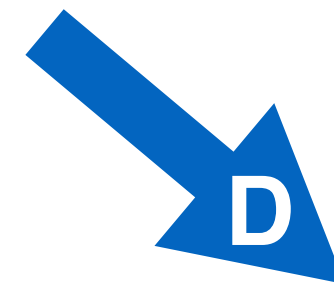
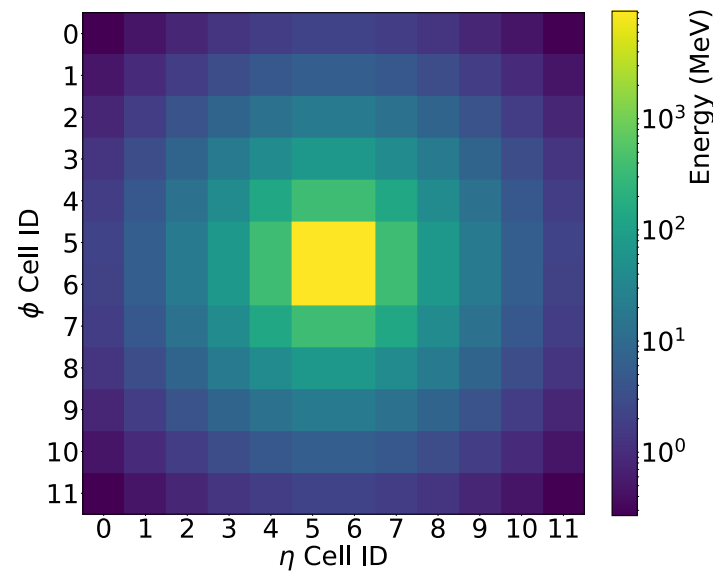
17

Vanilla GAN loss:

$$-\sum_{X \sim \text{data}} \log(D(x)) - \sum_{Z \sim \text{noise}} \log(1 - D(G(z)))$$

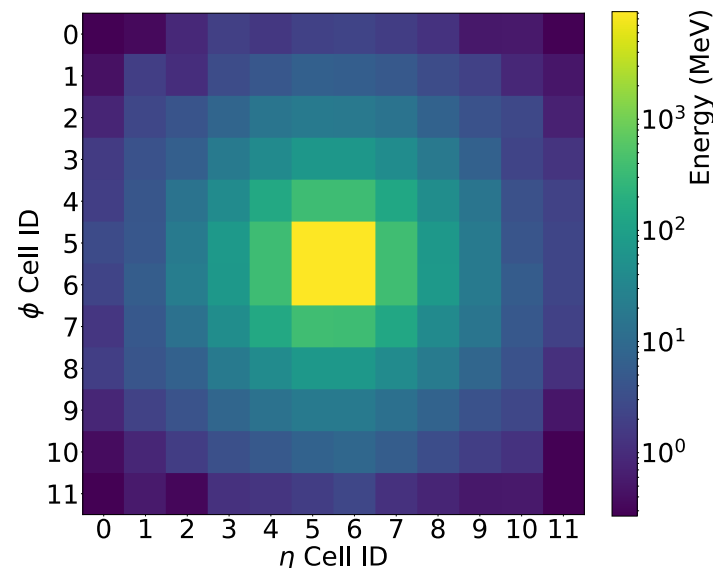


noise



{real, fake}

When **D** is maximally confused, **G** will be a good generator



Physics-based simulator or data

# Introduction: GANs

18

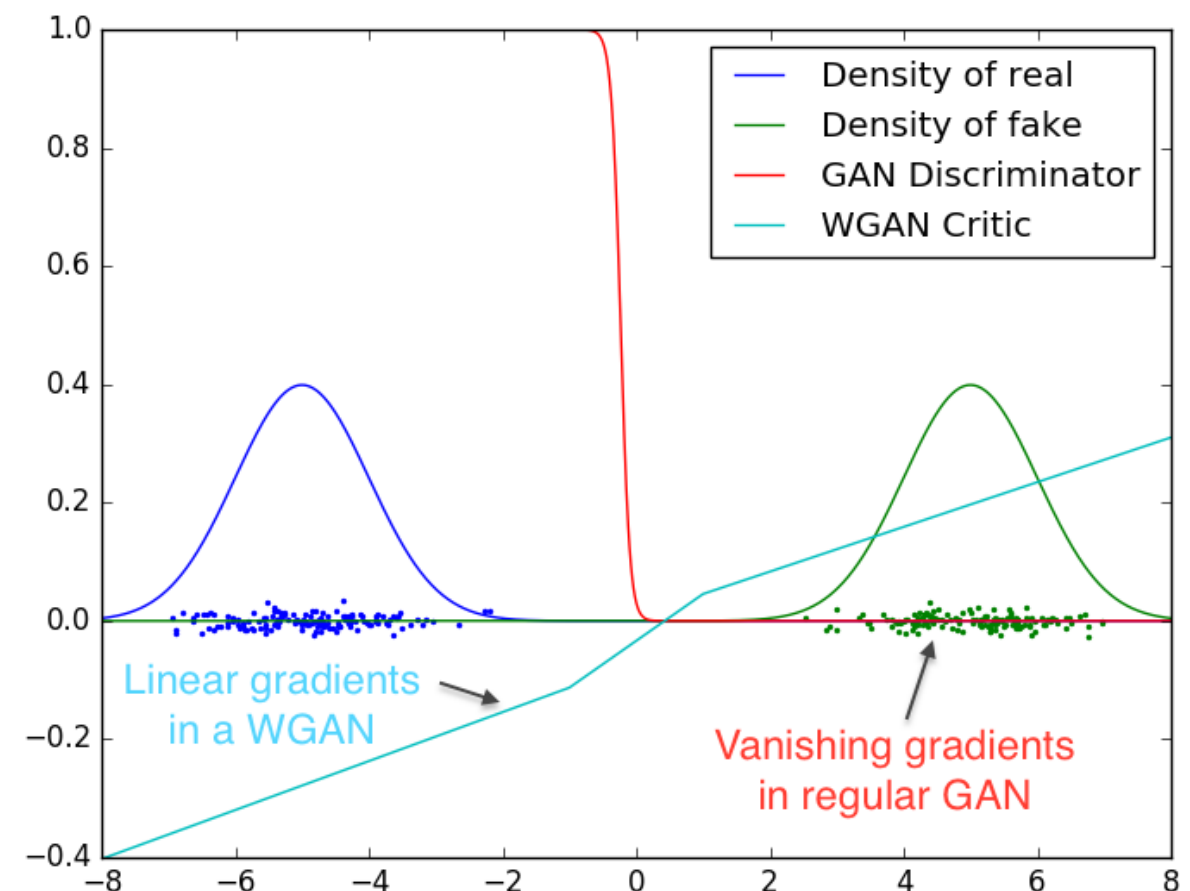
Vanilla GAN loss:

$$-\sum_{X \sim \text{data}} \log(D(x)) - \sum_{Z \sim \text{noise}} \log(1 - D(G(z)))$$

There are many variations on this theme\*.

One important variation is called the Wasserstein or WGAN.

Schematically, the idea of WGAN is to replace the discriminator with a notion of distance ('earth moving') between the real and fake data



\*see e.g. <https://github.com/hindupuravinash/the-gan-zoo>

# GANs: overview

19

## Pros:

- Fast to sample
- Easy to specify (no restrictions on G)
- Easy to make high dimensional

## Cons:

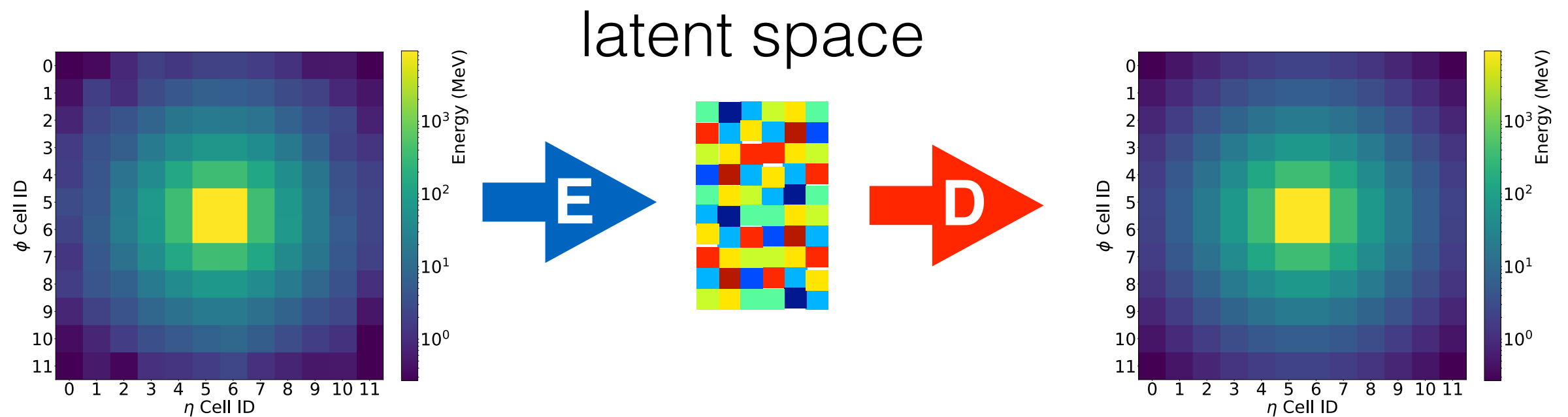
- No access to density
- (Very) Hard to train (minimax)
- Mode collapse

# Introduction: AEs

20

Autoencoders (AEs):

*A pair of networks that embed the data into a latent space and decode back to the data space.*



Physics-based  
simulator or data

**encoder**

**decoder**

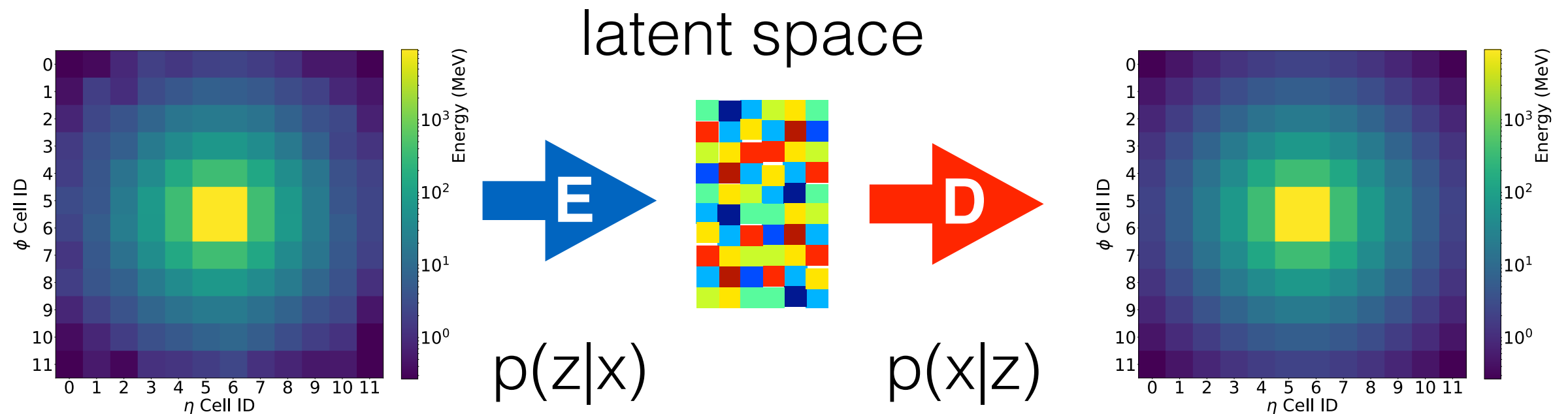


# Introduction: VAEs

21

Variational Autoencoders (VAEs):

*A pair of networks that embed the data into a latent space with a given prior and decode back to the data space.*



Physics-based  
simulator or data

*Probabilistic*  
**encoder**

*Probabilistic*  
**decoder**

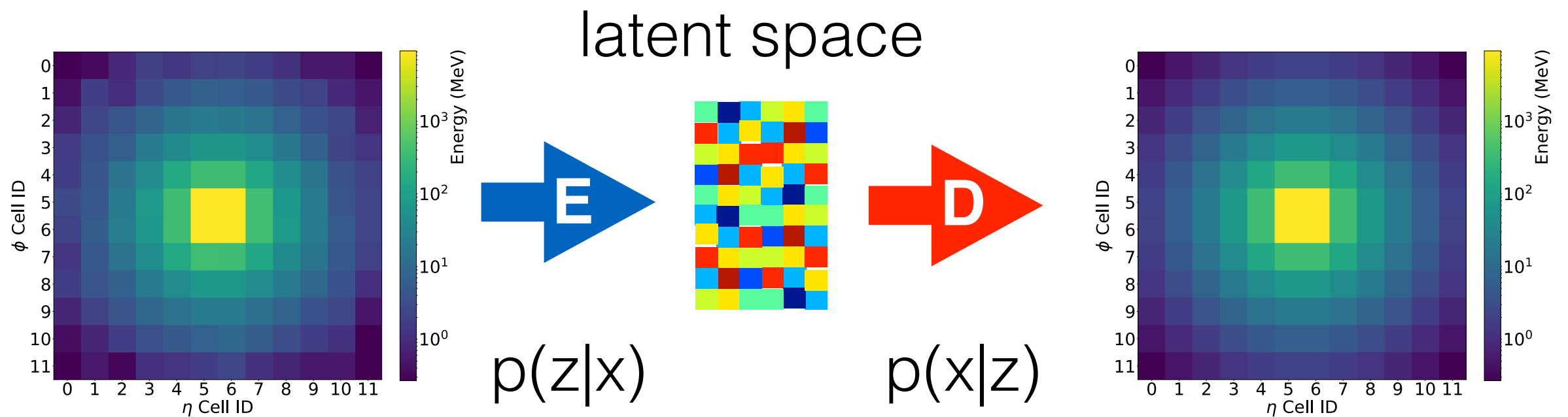
# Introduction: VAEs

Vanilla VAE loss:

$$\sum_{Z \sim p(z|x)} (x - D(z))^2 + \text{KL}(p(z|x) || p_0(z))$$

↑ “reconstruction”      ↑ “regularization”

← prior      ← KL-divergence



Physics-based simulator or data

*Probabilistic encoder*  
**encoder**

*Probabilistic decoder*  
**decoder**

Why “variational”? Another way of looking at this is the Evidence Lower Bound (ELBO) and  $p(z|x)$  is the variational posterior.

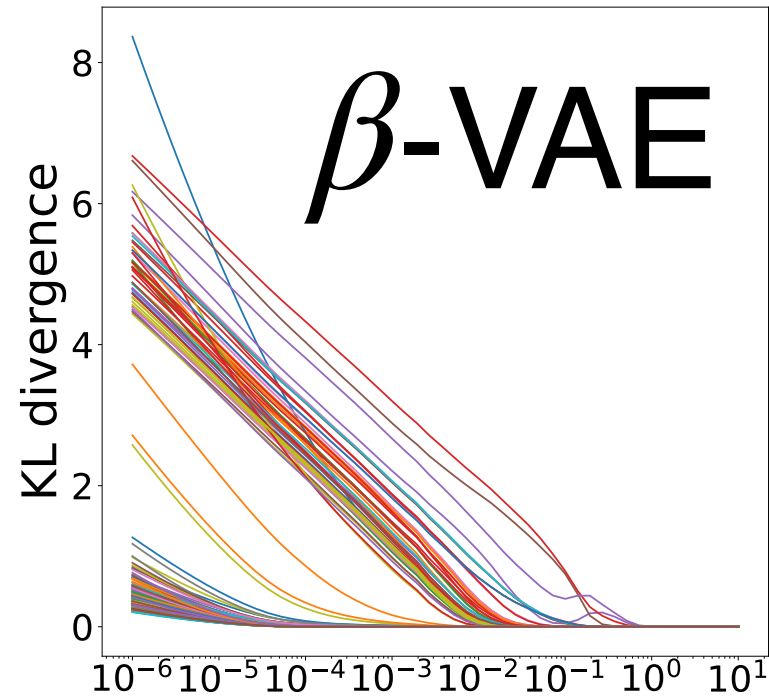
## Pros:

- Fast to sample
- Easy to specify (no restrictions on  $G$ )
- Easy to make high dimensional
- Easy to train

## Cons:

- No access to density
- Tends to over-smooth the density

# ex. AE variations: $\beta$ -VAE, PAE



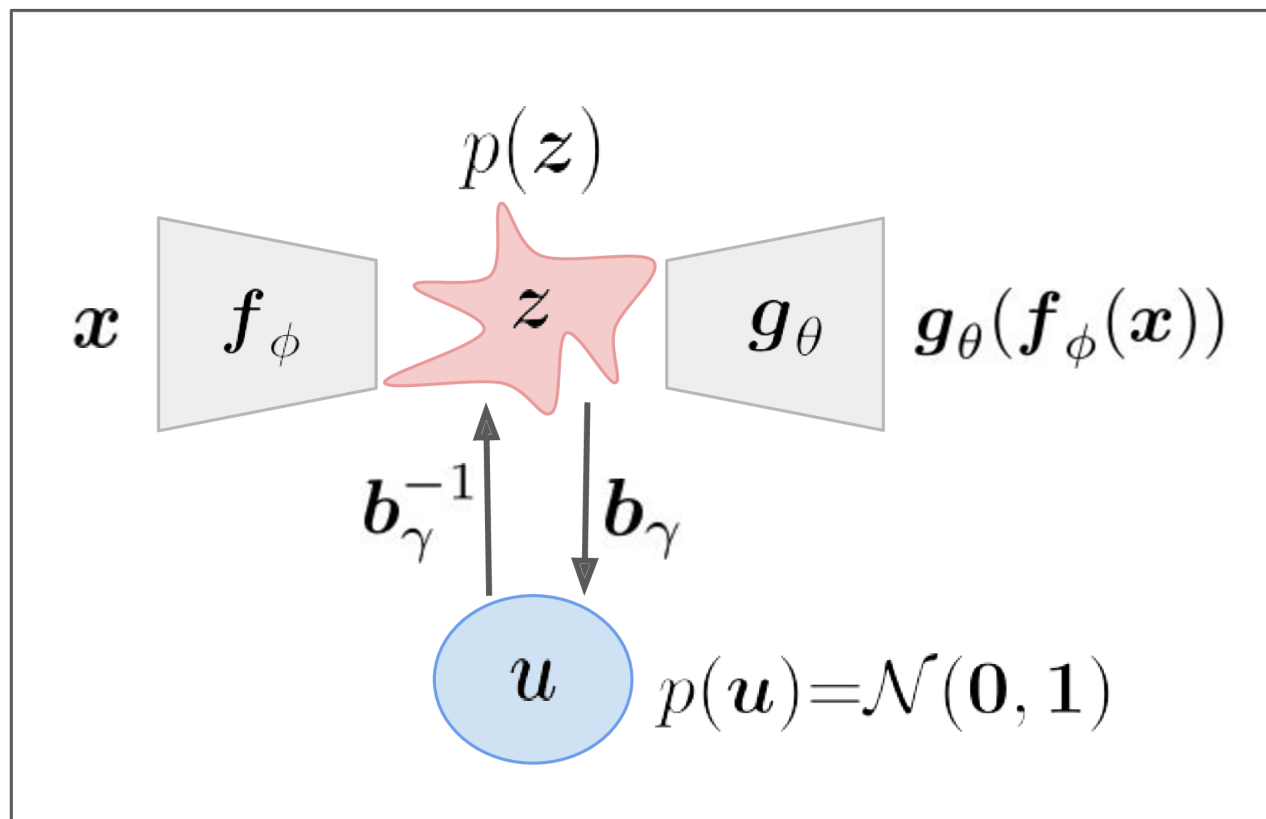
$\beta$ -VAE loss:

$$\sum_{Z \sim p(z|x)} (x - D(z))^2 + \beta \text{KL}(p(z|x) || p_0(z))$$

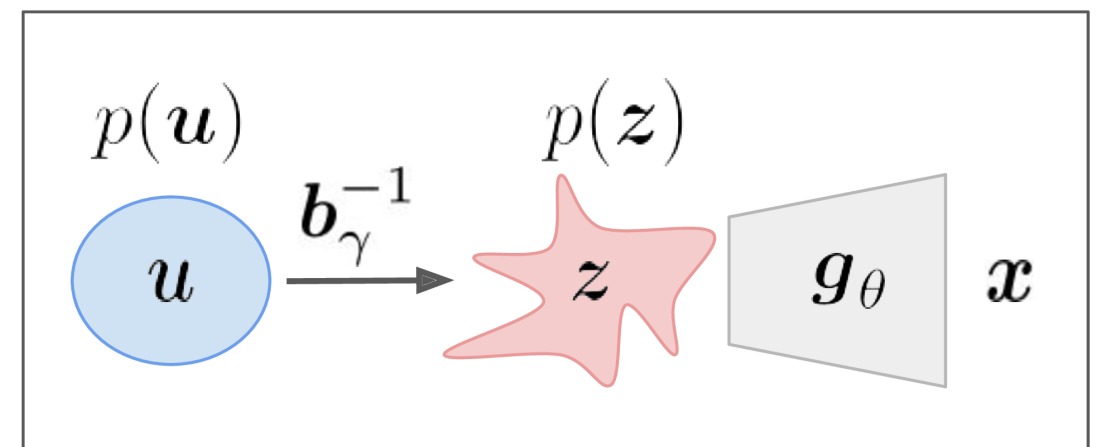
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$\beta$

2006.05479



Probabilistic AE



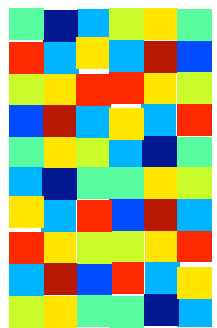
# Introduction: NFs

25

Normalizing Flows (NFs):

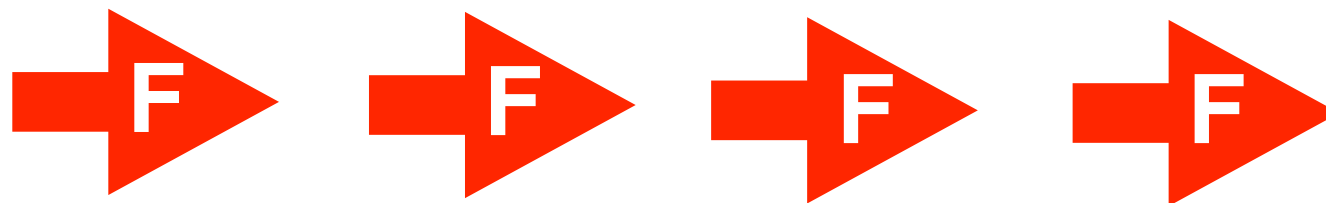
*A series of invertible transformations mapping a known density into the data density.*

$$\text{Loss: } -\log(p(x))$$



latent space

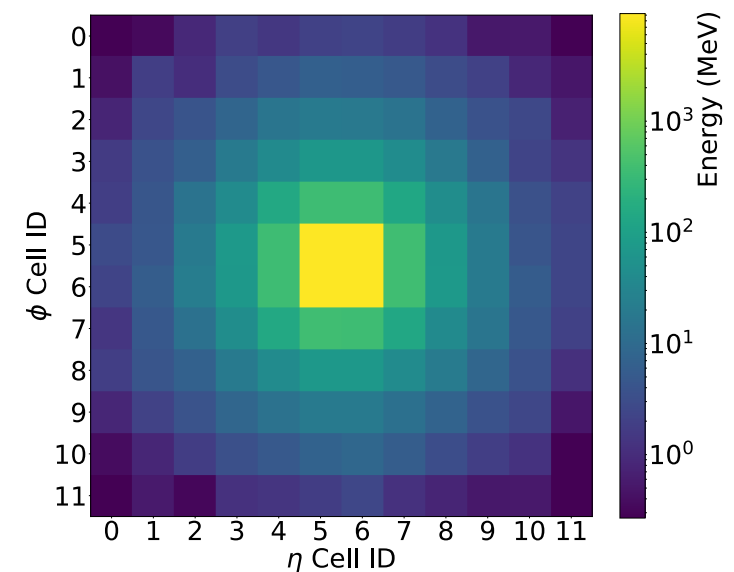
$p(z)$



*Invertible transformations with tractable **Jacobians***

$$p(x) = p(z) \left| \frac{dF^{-1}}{dx} \right|$$

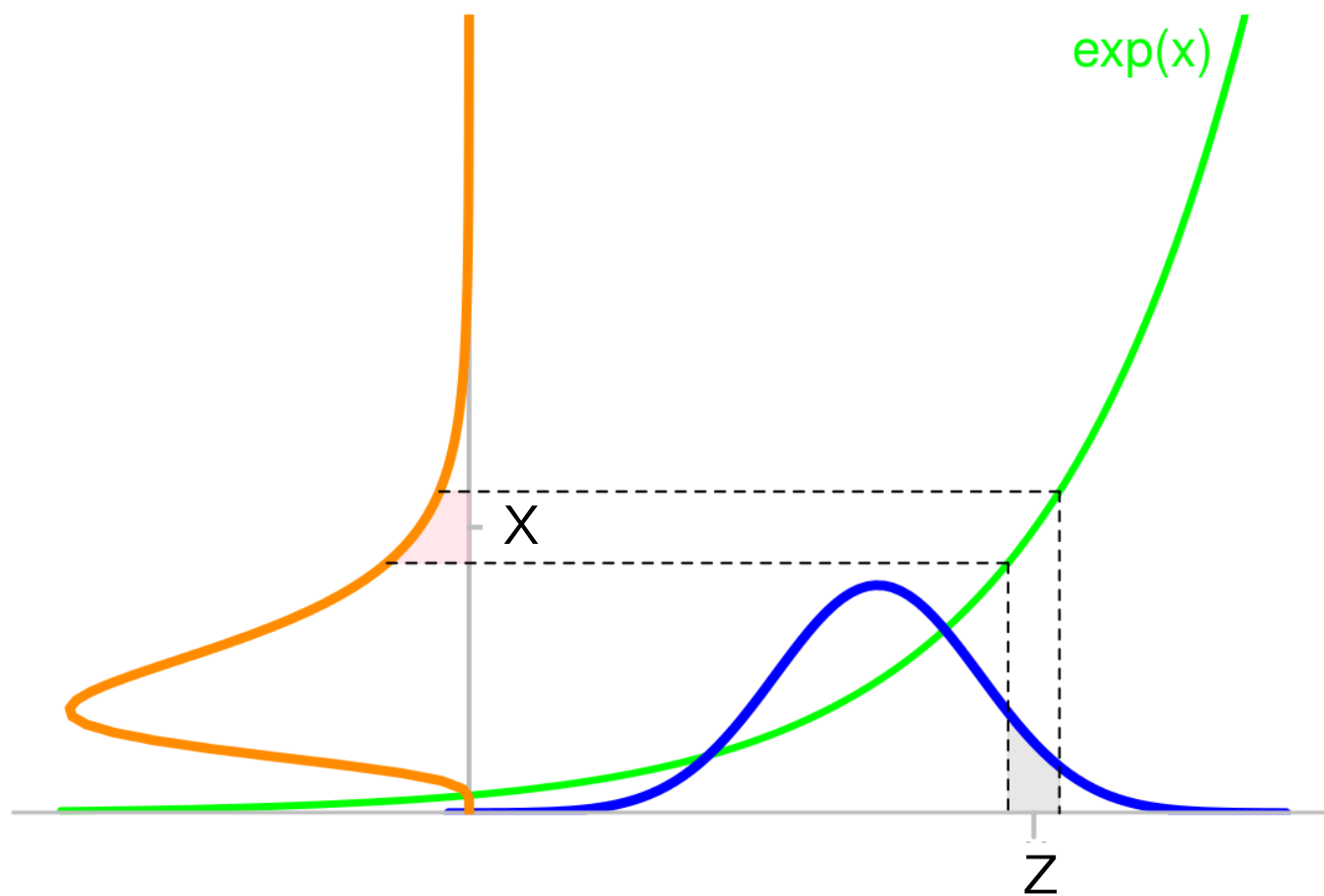
Optimize via maximum likelihood



$p(x)$

# Introduction: NFs

In case you have not seen the change of variables formula...



*The Jacobian is a matrix of partial derivatives.*

*Food for thought:  
Suppose the PDF of  $X$  is  $f$   
and the CDF is  $F$ . What is  
the density of  $F(X)$ ?*

latent  
space

*with tractable **Jacobians***

$p(z)$

$$p(x) = p(z) \left| \frac{dF^{-1}}{dx} \right|$$

$p(x)$

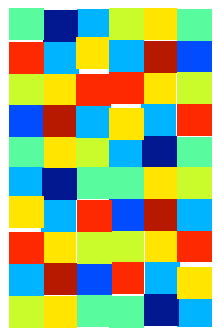
# Introduction: NFs

27

Normalizing Flows (NFs):

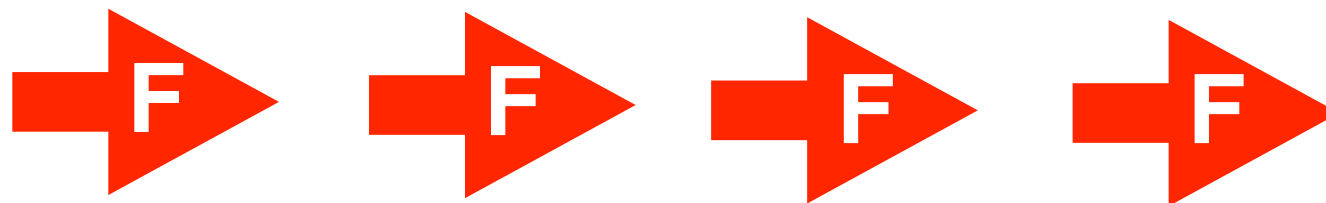
*A series of invertible transformations mapping a known density into the data density.*

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latent space

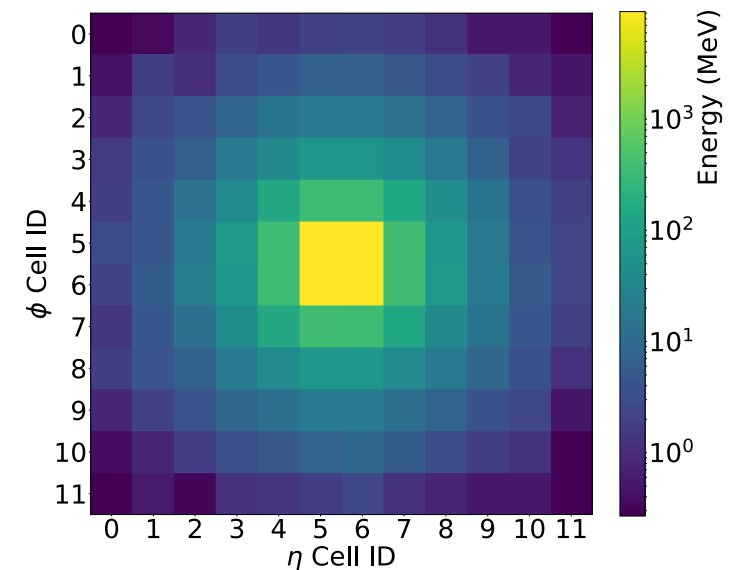
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$p(x)$

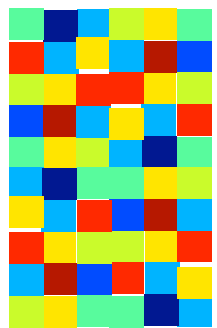
# Introduction: NFs

28

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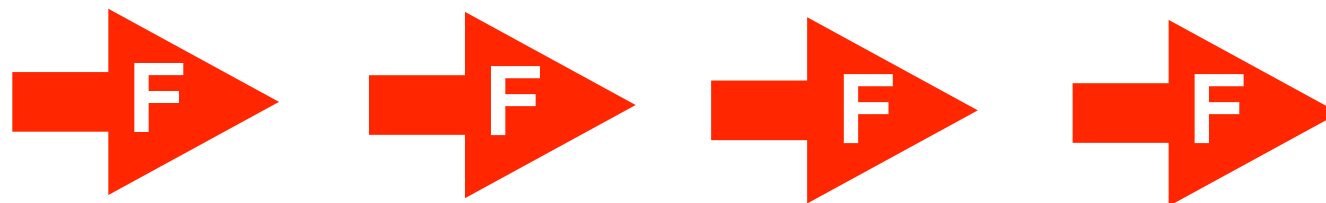
*A series of invertible transformations mapping a known density into the data density.*

Example:  $F(x_i) = z_i + \text{NN}(x_{i-1})$   
("autoregressive")



latent  
space

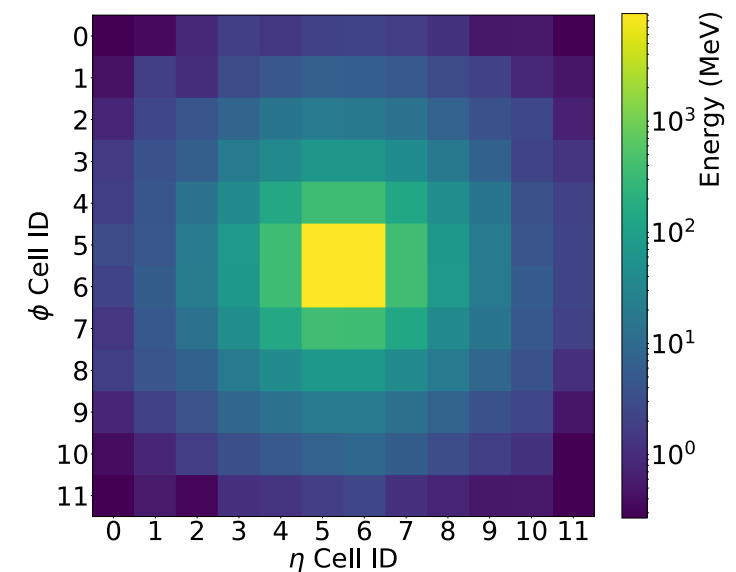
$p(z)$



*Invertible transformations  
with tractable **Jacobians***

$$p(x) = p(z) \left| \frac{dF^{-1}}{dx} \right|$$

Optimize via  
maximum likelihood



$p(x)$



## Pros:

- Usually only fast to sample OR to estimate the density
- Access to the density
- Easy to train

## Cons:

- Sometimes hard to make expressive enough
- “Generator” is highly limited in form
- Cannot learn topology

# Introduction: Score-based

30

Score-based

*Learn the gradient of the density instead of the probability density itself.*

$$\text{Loss: } |f(x) - \overset{\text{“score”}}{\nabla p(x)}|^2$$

...but since we don't know  $\nabla p$ , we make use of a trick whereby the data are perturbed and it is sufficient to match the score of the perturbing function (!)

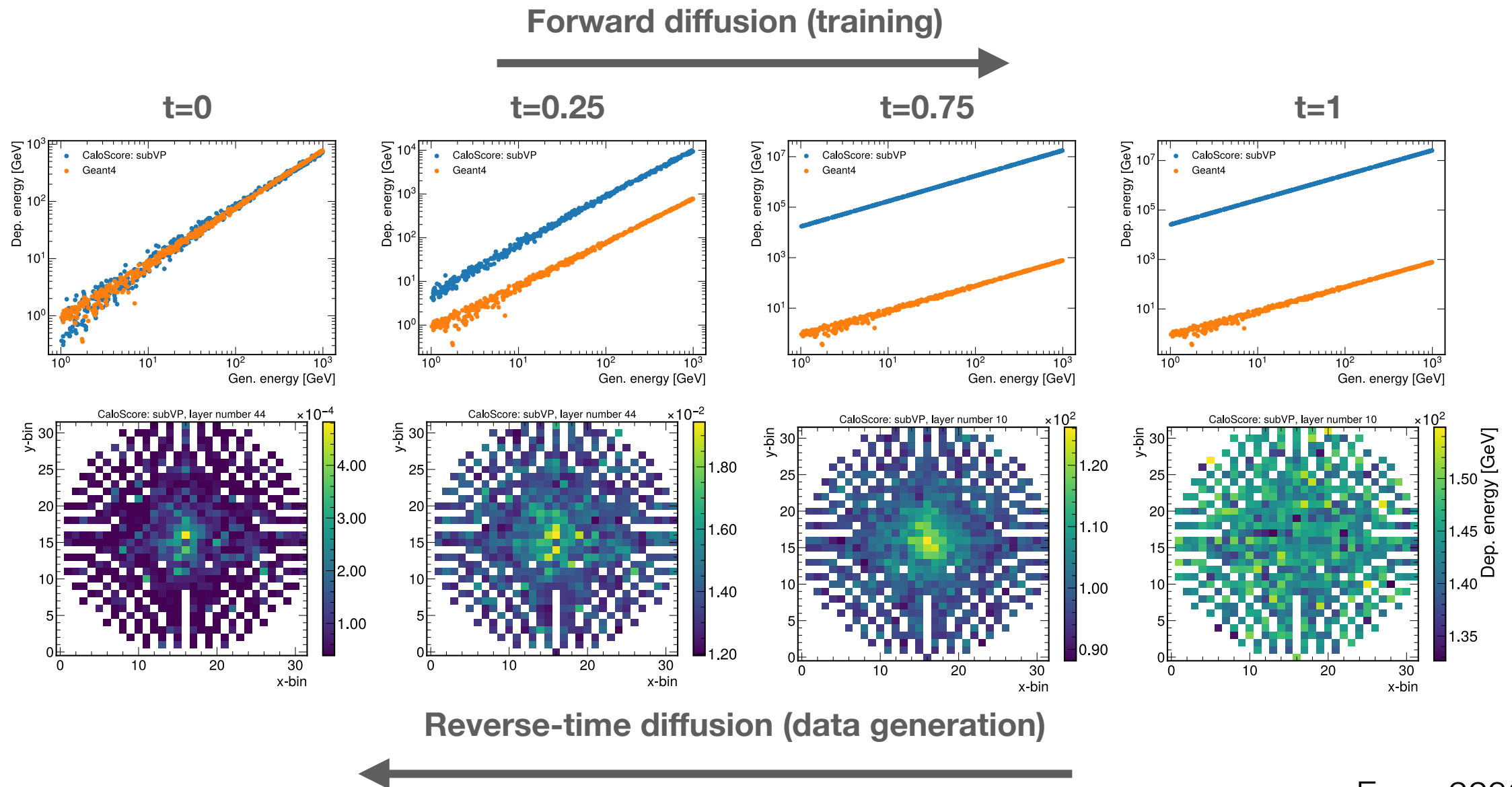
*This turns the problem into a stochastic differential equation which is the same as diffusion (sometimes this is called a “diffusion model”)*

# Introduction: Score-based

31

Score-based

*Learn the gradient of the density instead of the probability density itself.*



# Score-based: overview

32

## Pros:

- Many of the benefits of a NF, but unrestricted functions
- Access to the density with some work
- Easy to train
- Currently the “best” on the market

## Cons:

- A bit slow to train / evaluate

## Image Generation on CIFAR-10

Rank	Model	FID	Clean-FID-10k	Inception score	bits/dimension	FID-10k-test	Paper	Code	Result	Year	Tags
1	EDM-G++ (unconditional)	1.77			2.55		<a href="#">Refining Generative Process with Discriminator Guidance in Score-based Diffusion Models</a>	<a href="#">Code</a>	<a href="#">Result</a>	2022	Diffusion, Score-based
2	StyleGAN-XL	1.85					<a href="#">StyleGAN-XL: Scaling StyleGAN to Large Diverse Datasets</a>	<a href="#">Code</a>	<a href="#">Result</a>	2022	GAN
3	STF (unconditional)	1.90					<a href="#">Stable Target Field for Reduced Variance Score Estimation in Diffusion Models</a>	<a href="#">Code</a>	<a href="#">Result</a>	2023	Score-based
4	LSGM-G++ (FID)	1.94			3.42		<a href="#">Refining Generative Process with Discriminator Guidance in Score-based Diffusion Models</a>	<a href="#">Code</a>	<a href="#">Result</a>	2022	Score-based
5	PSLD (ODE)	2.10		9.93			<a href="#">Generative Diffusions in Augmented Spaces: A Complete Recipe</a>	<a href="#">Code</a>	<a href="#">Result</a>	2023	Score-based
6	LSGM (FID)	2.10			3.43		<a href="#">Score-based Generative Modeling in Latent Space</a>	<a href="#">Code</a>	<a href="#">Result</a>	2021	Score-based, VAE
7	Subspace Diffusion (NSCN++)	2.17		9.99			<a href="#">Subspace Diffusion Generative Models</a>	<a href="#">Code</a>	<a href="#">Result</a>	2022	Score-based
8	LSGM (balanced)	2.17			2.95		<a href="#">Score-based Generative Modeling in Latent Space</a>	<a href="#">Code</a>	<a href="#">Result</a>	2021	Score-based, VAE
9	NCSN++	2.20		9.73			<a href="#">Score-Based Generative Modeling through Stochastic Differential Equations</a>	<a href="#">Code</a>	<a href="#">Result</a>	2020	Score-based
10	PSLD (SDE)	2.21					<a href="#">Generative Diffusions in Augmented Spaces: A Complete Recipe</a>	<a href="#">Code</a>	<a href="#">Result</a>	2023	Score-based

# A few words about statistics

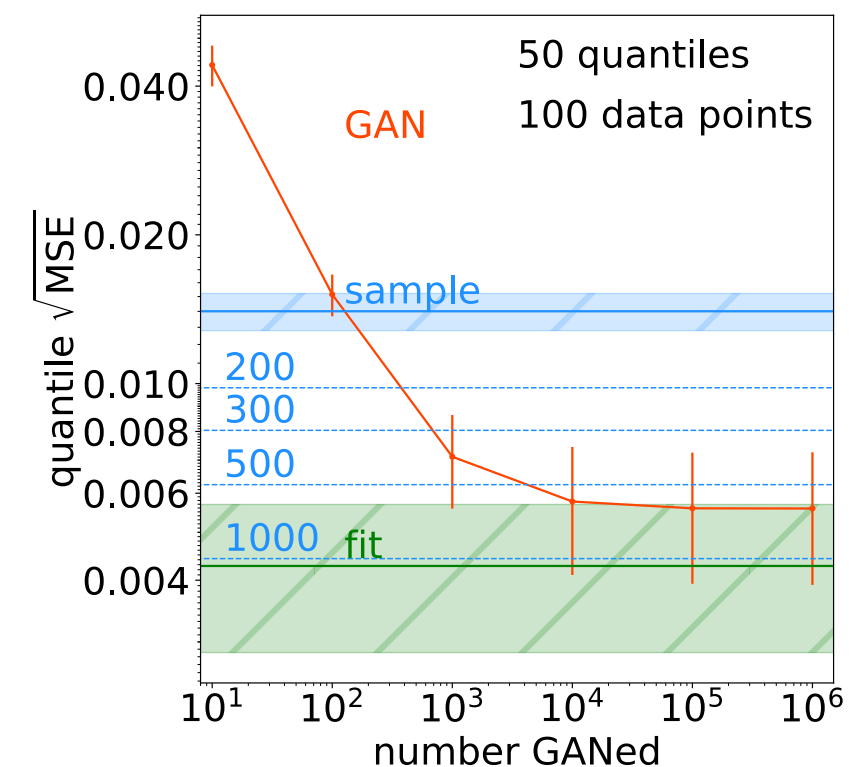
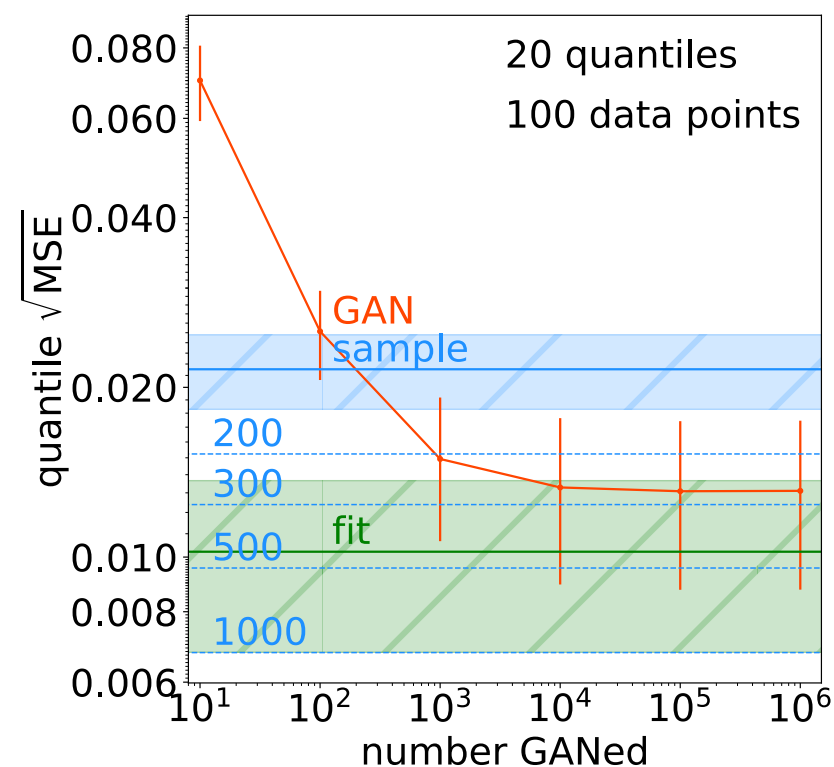
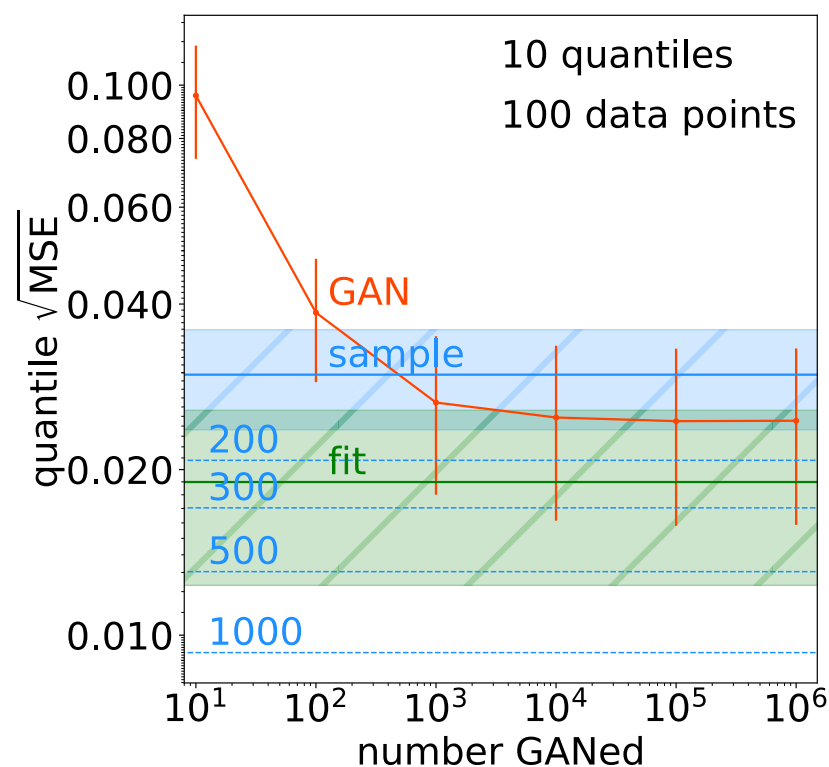




# Statistical Amplification

Performance continues to improve on many fronts. As we integrate these tools into our workflows, we need to think about uncertainties.

One question is about the **statistical power** of samples from a generative model. This depends on the implicit or explicit information we encode in the networks.

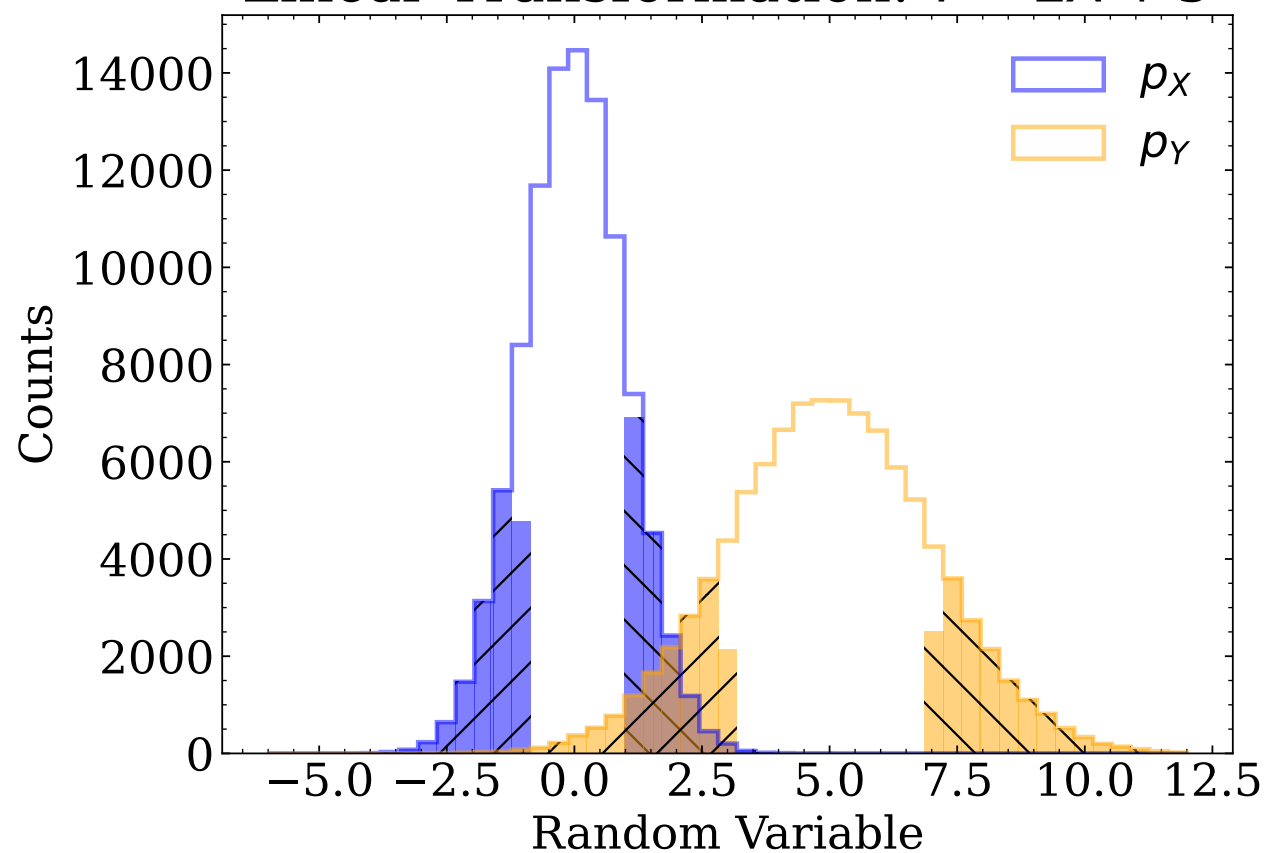


See also 1909.03081, 2002.06307, 2104.04543 (Generative Bayesian NNs), and 2107.08979 (“resampling”)

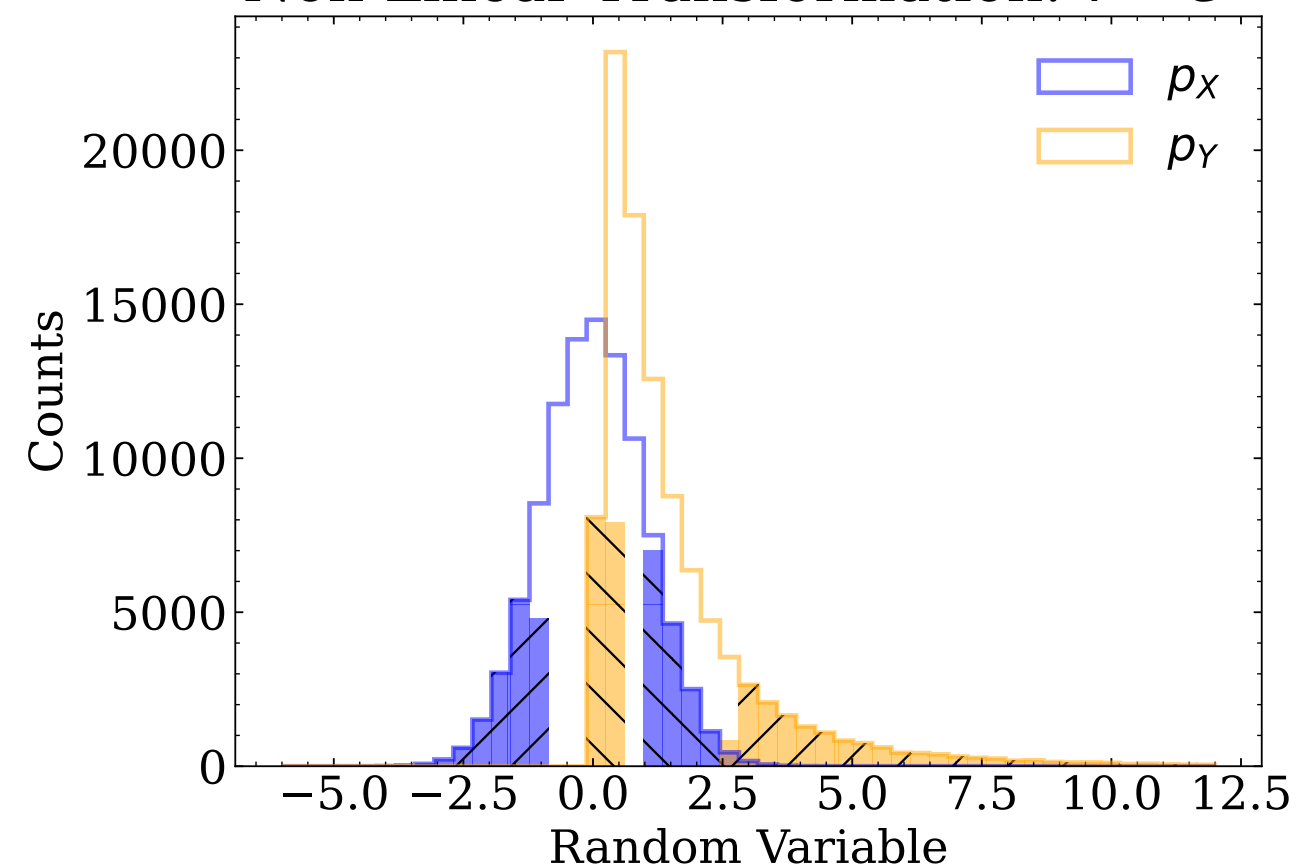
# Impact of (re)parameterization

Generative models are NOT invariant under coordinate transformations. Choose your coordinates wisely!

Linear Transformation:  $Y = 2X + 5$



Non-Linear Transformation:  $Y = e^{-X}$



# Applications

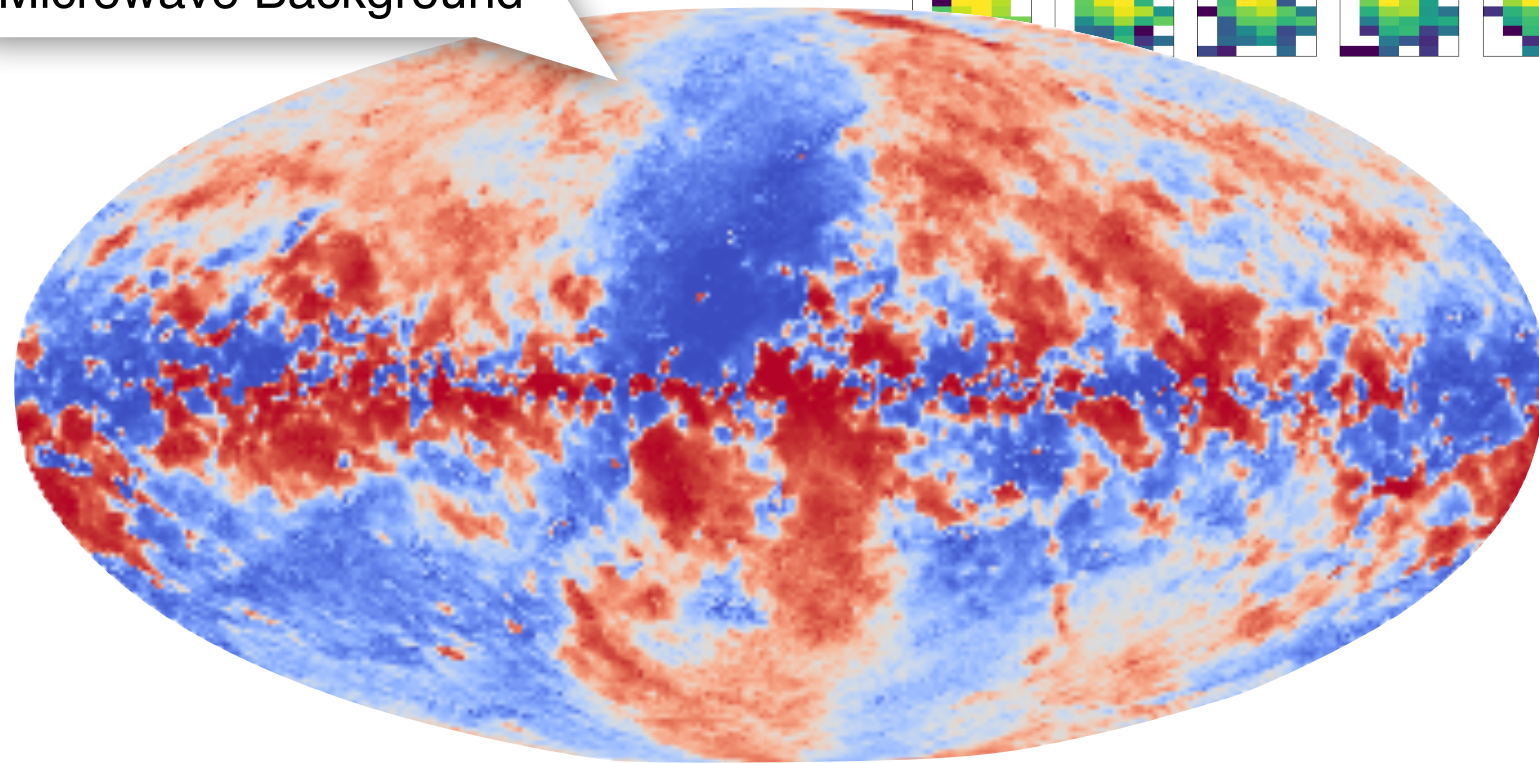




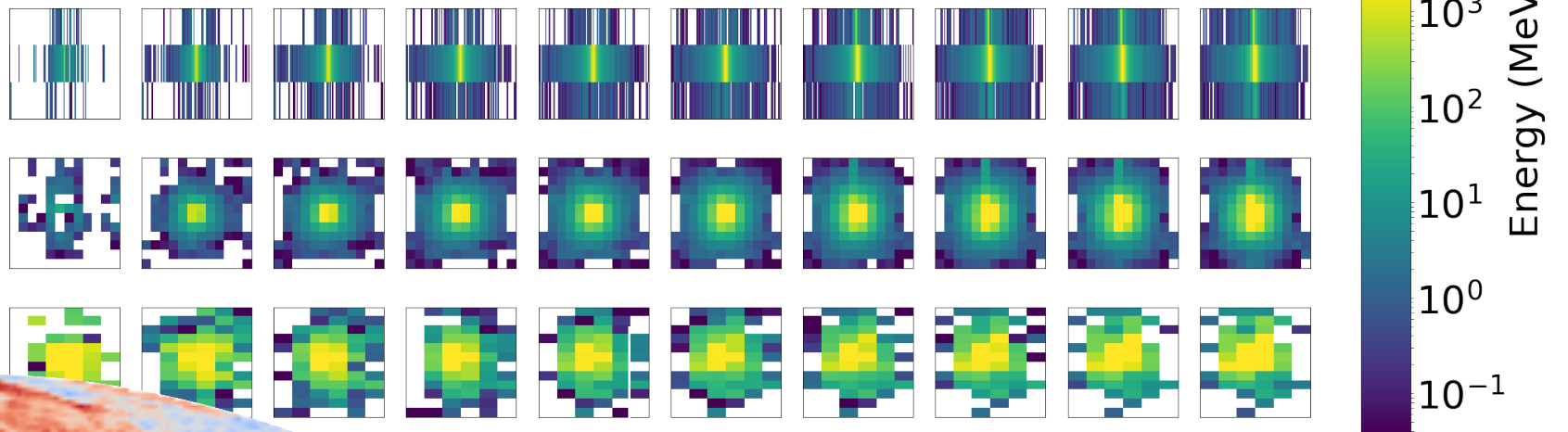
# Generative Models for Particle/Nuclear/Astro

**All of these pictures are fake!**

Synthetic Galactic radiation for Cosmic Microwave Background

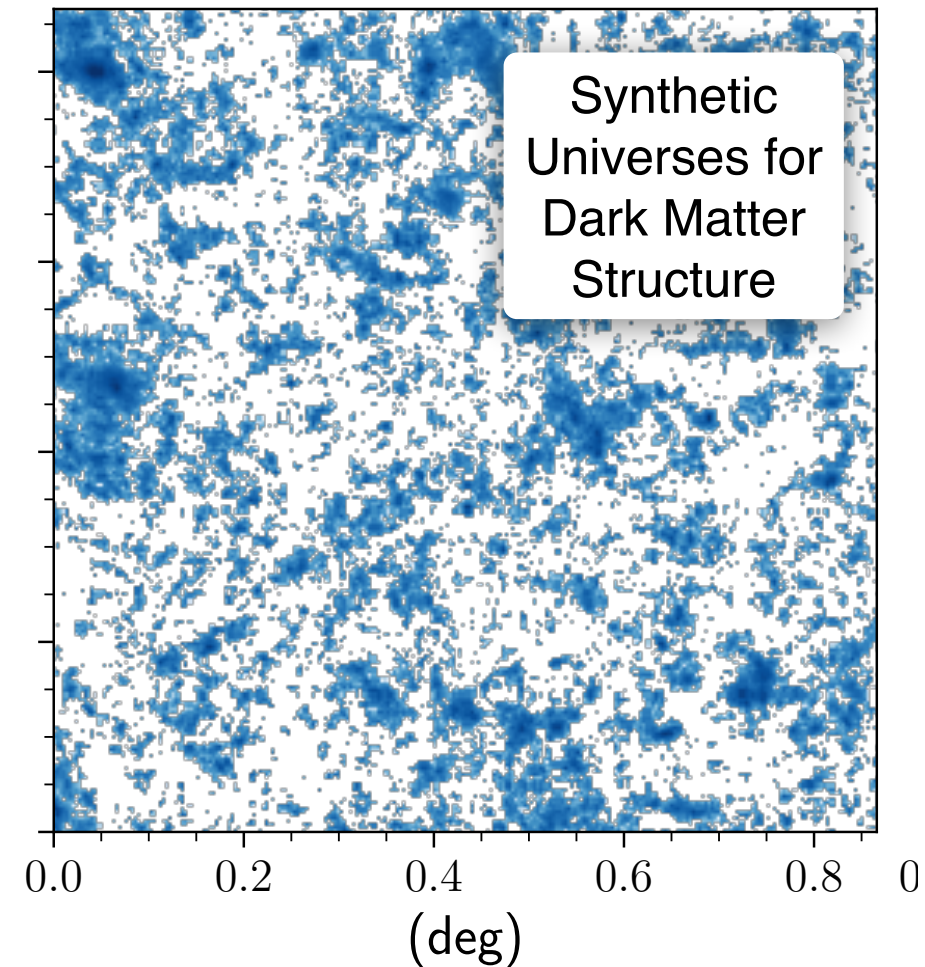


Material Interactions with High Energy Particles

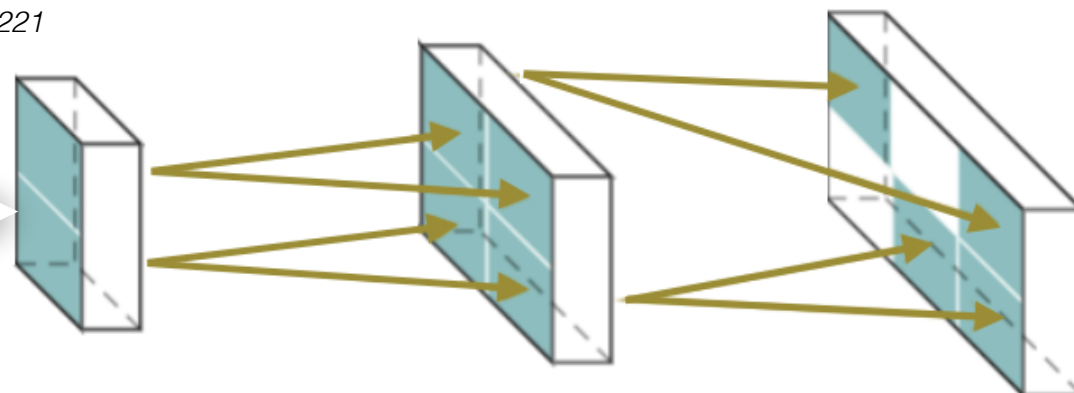


M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

Synthetic Universes for Dark Matter Structure



The Structure of Radiation in the Quantum Strong Force



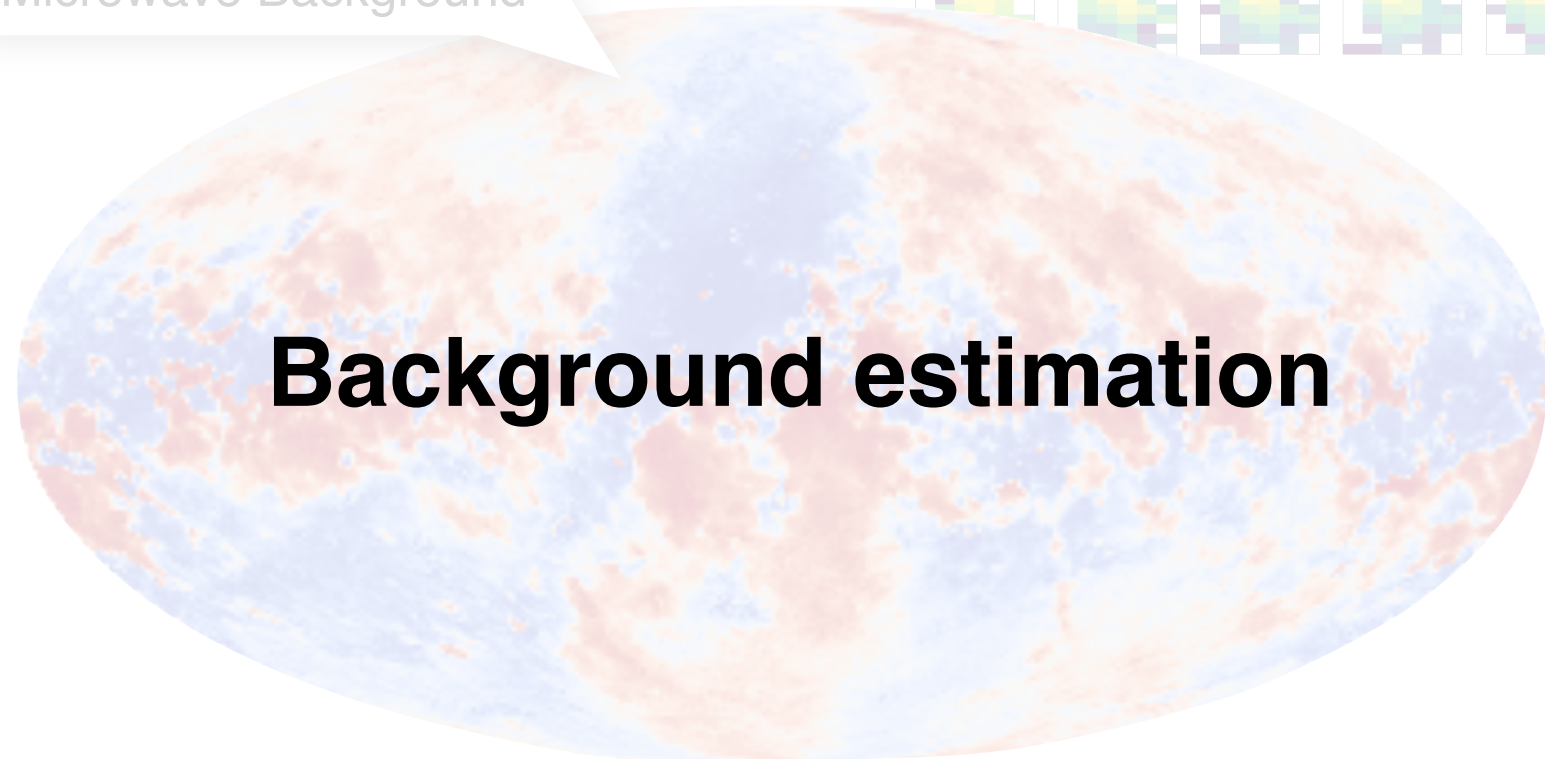
Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582

M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

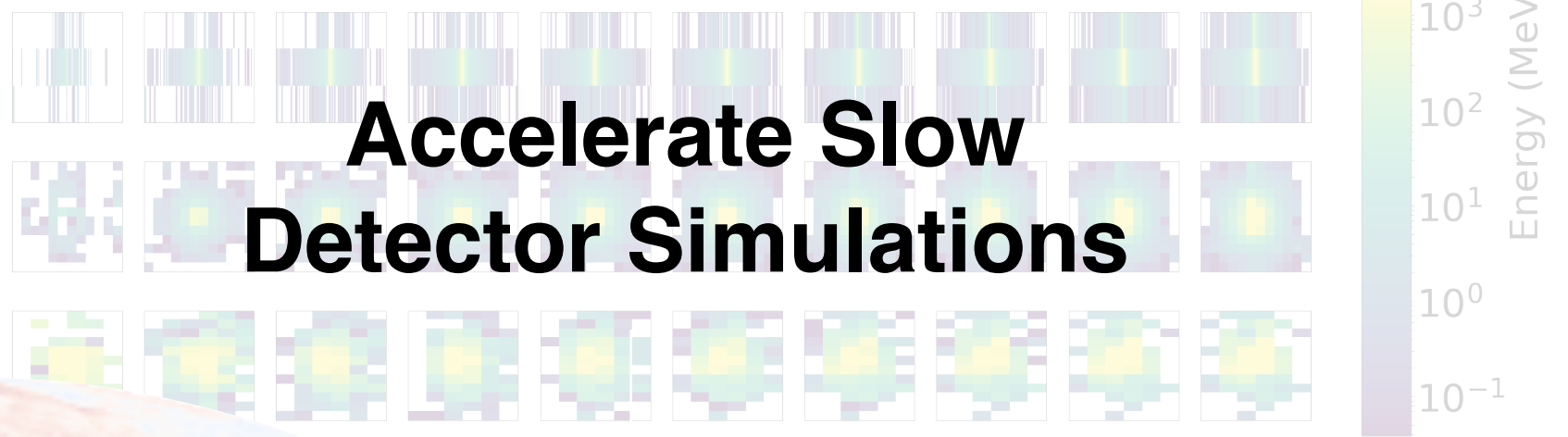
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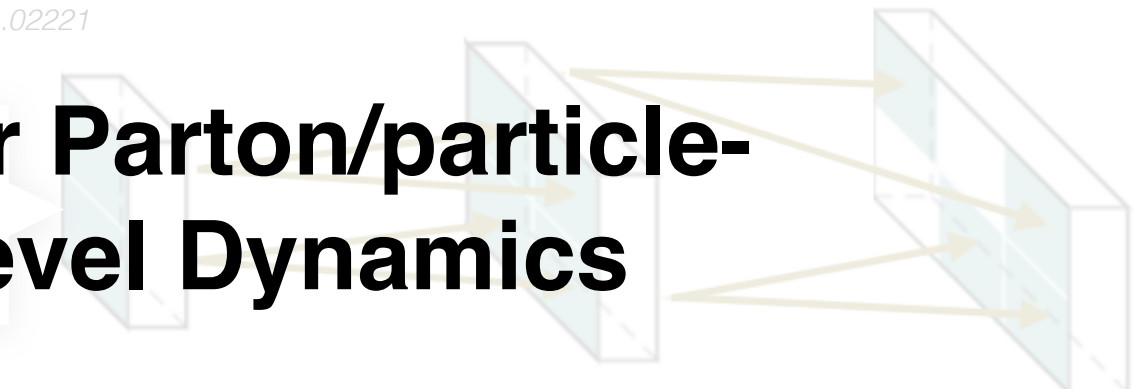


**Accelerate Slow Detector Simulations**

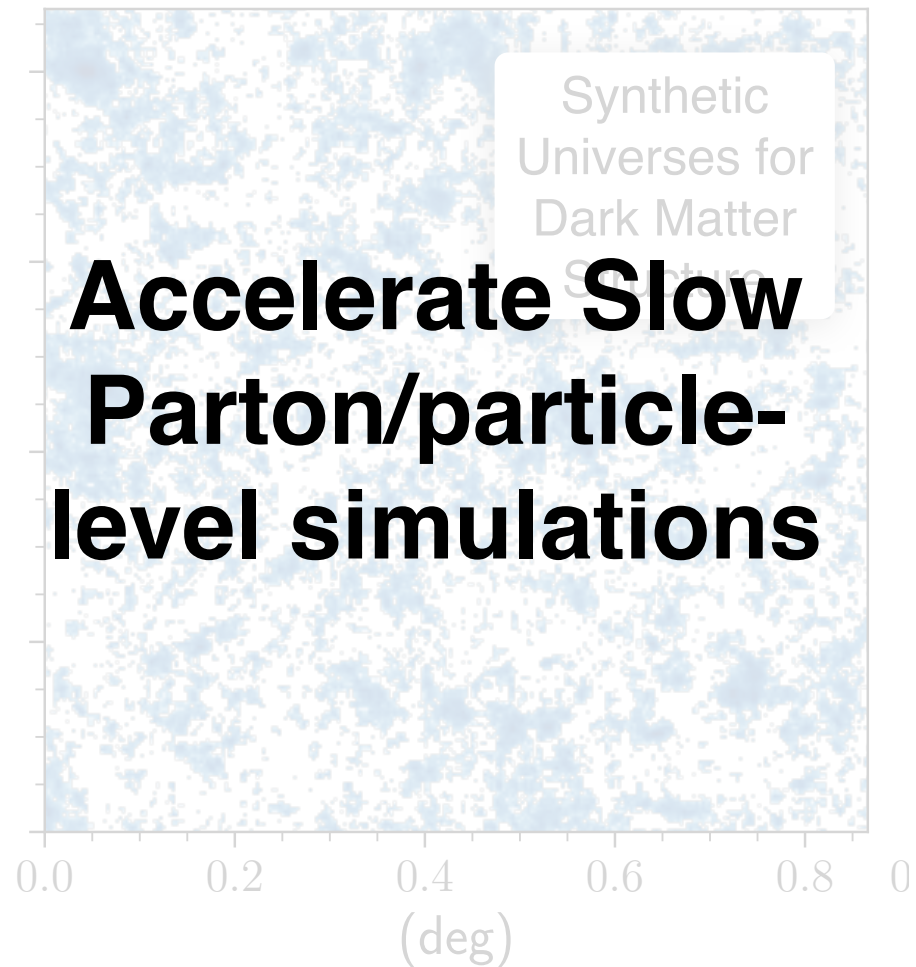
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The Structure of Radiation in the Quantum Strong Force

**Infer Parton/particle-level Dynamics**



Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582



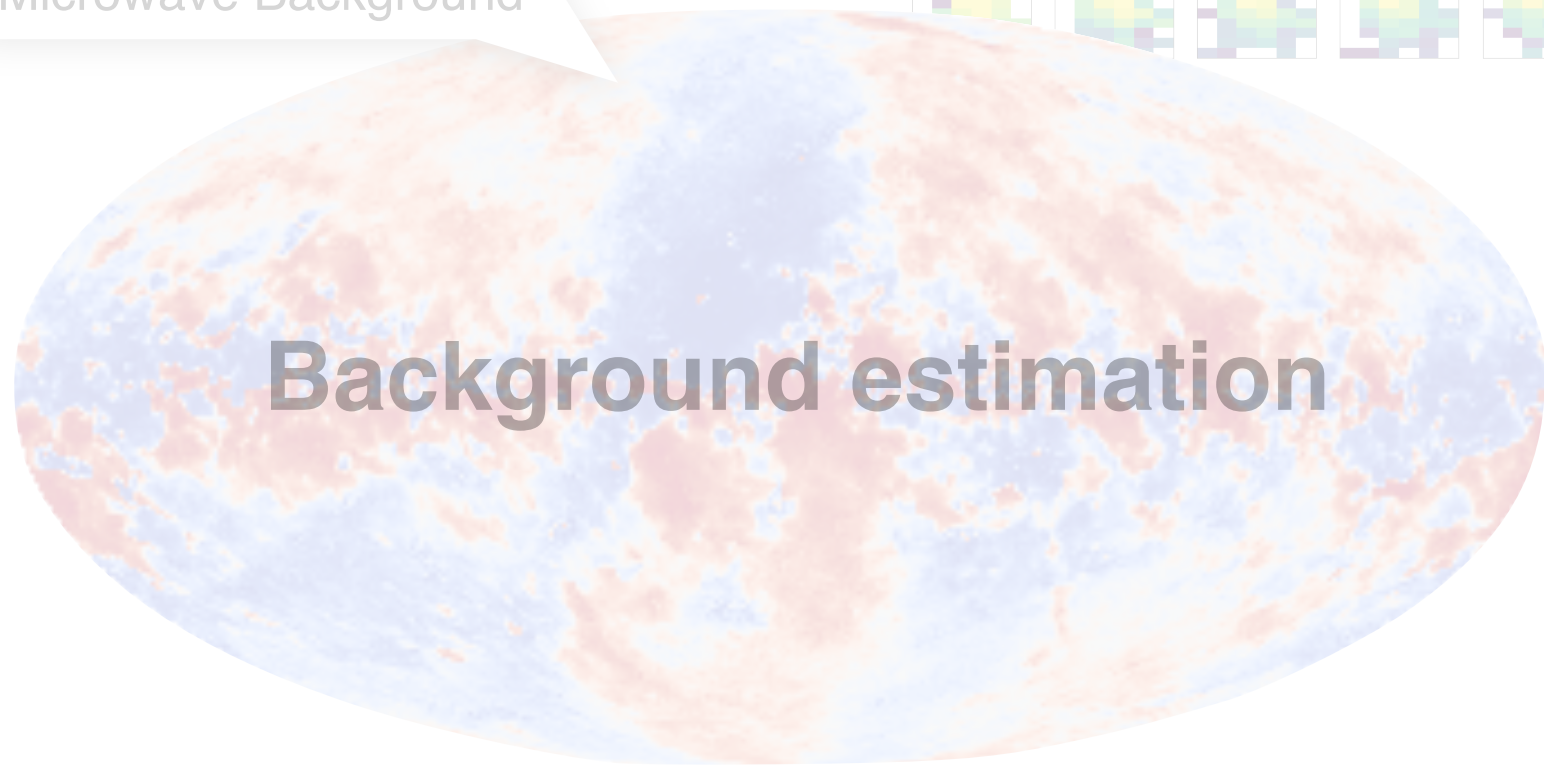
M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)



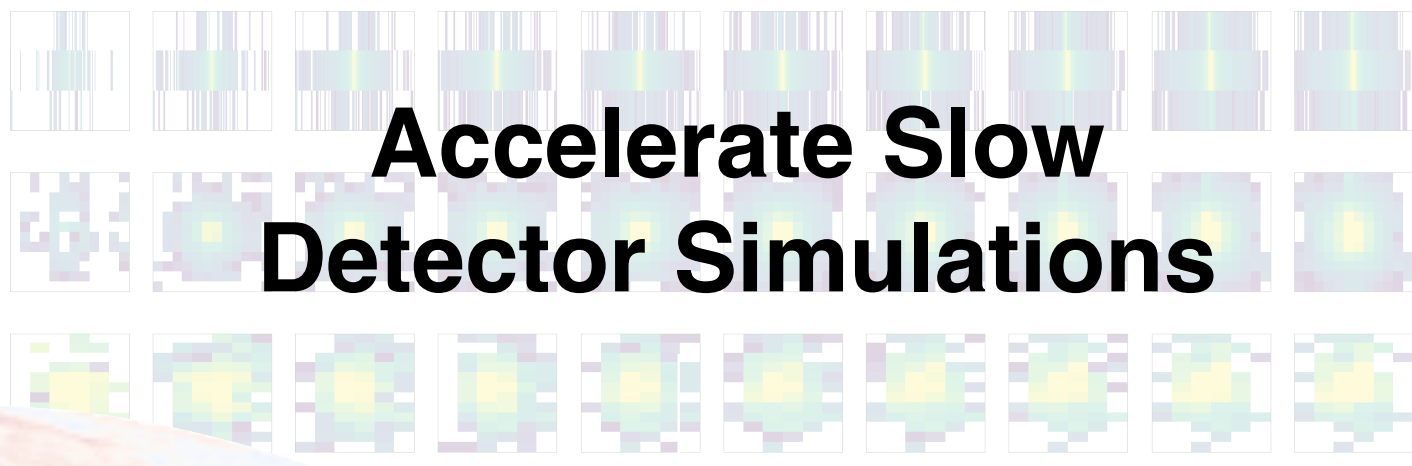
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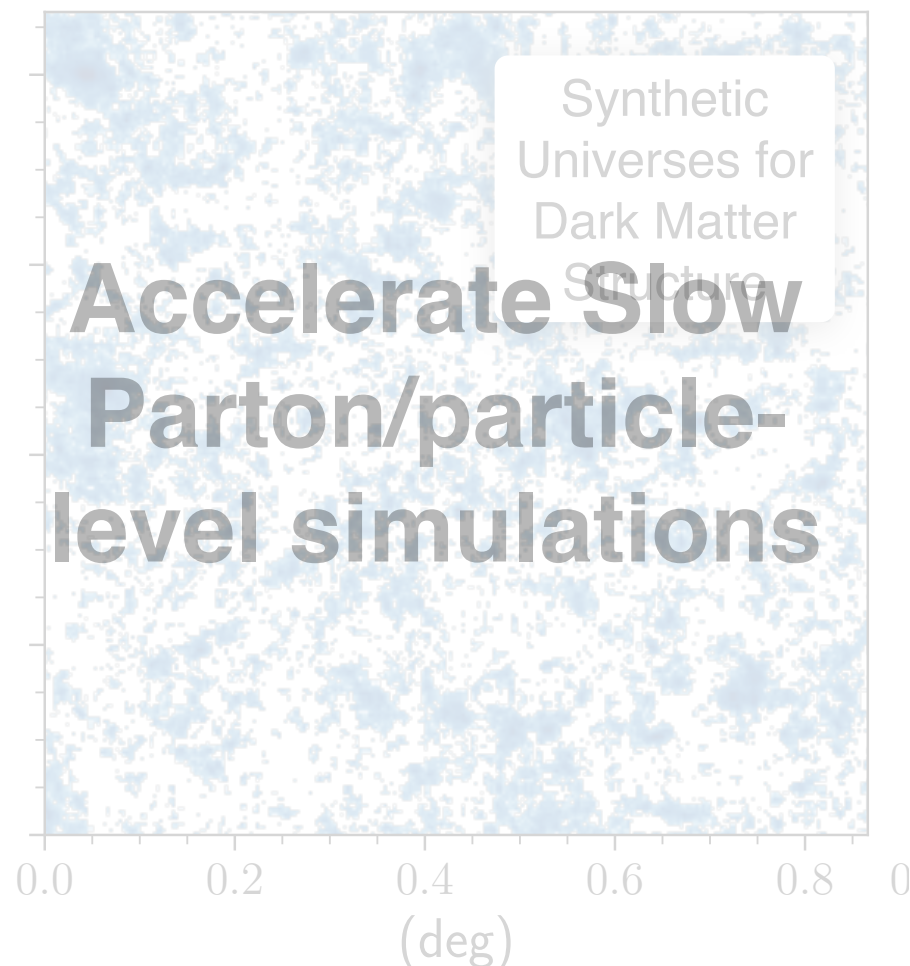


Material Interactions with High Energy Particles



Accelerate Slow Detector Simulations

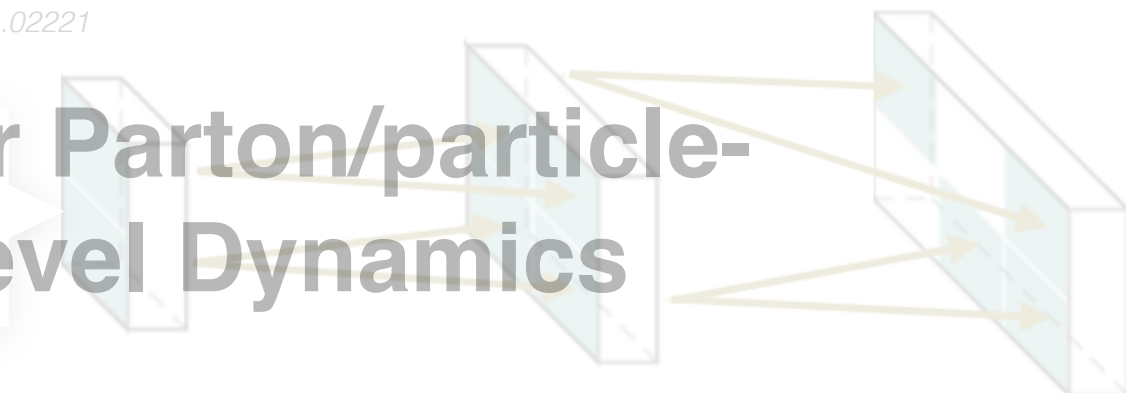
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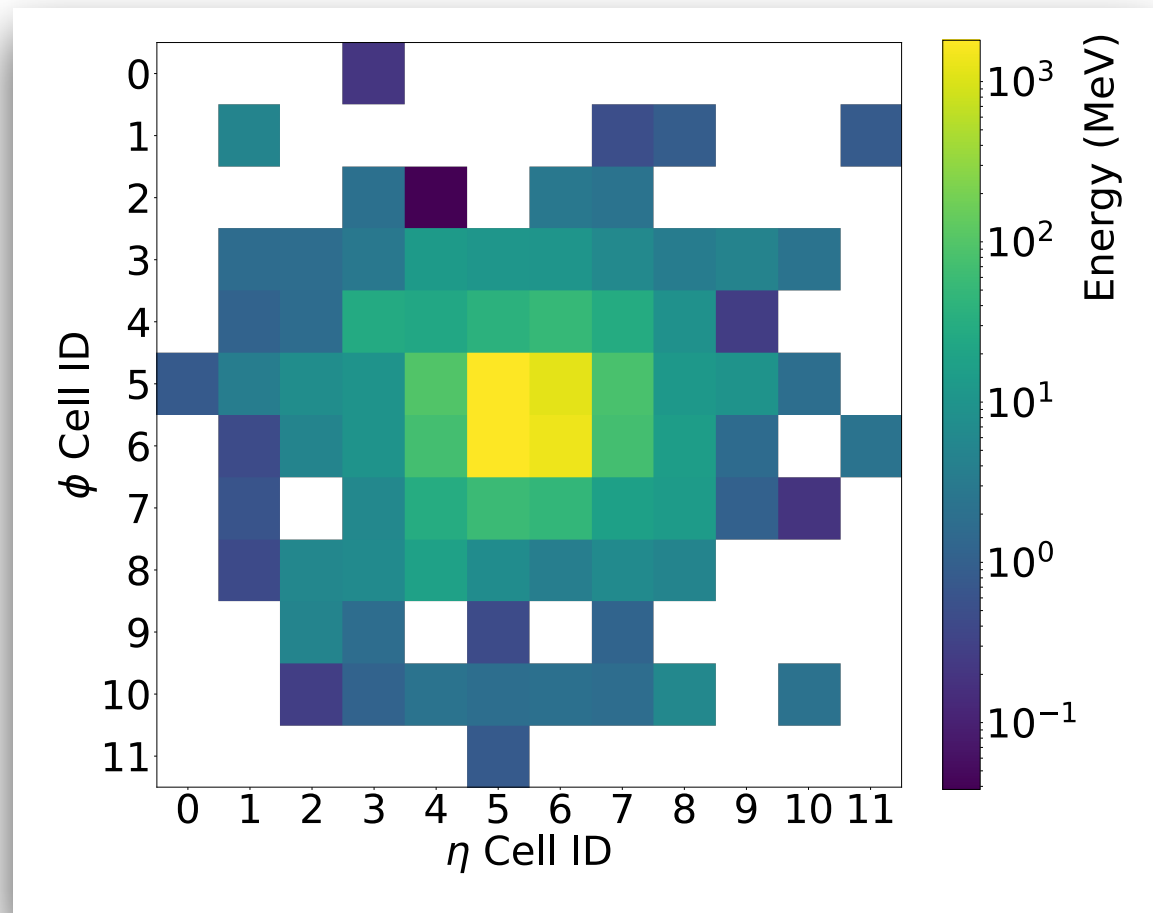


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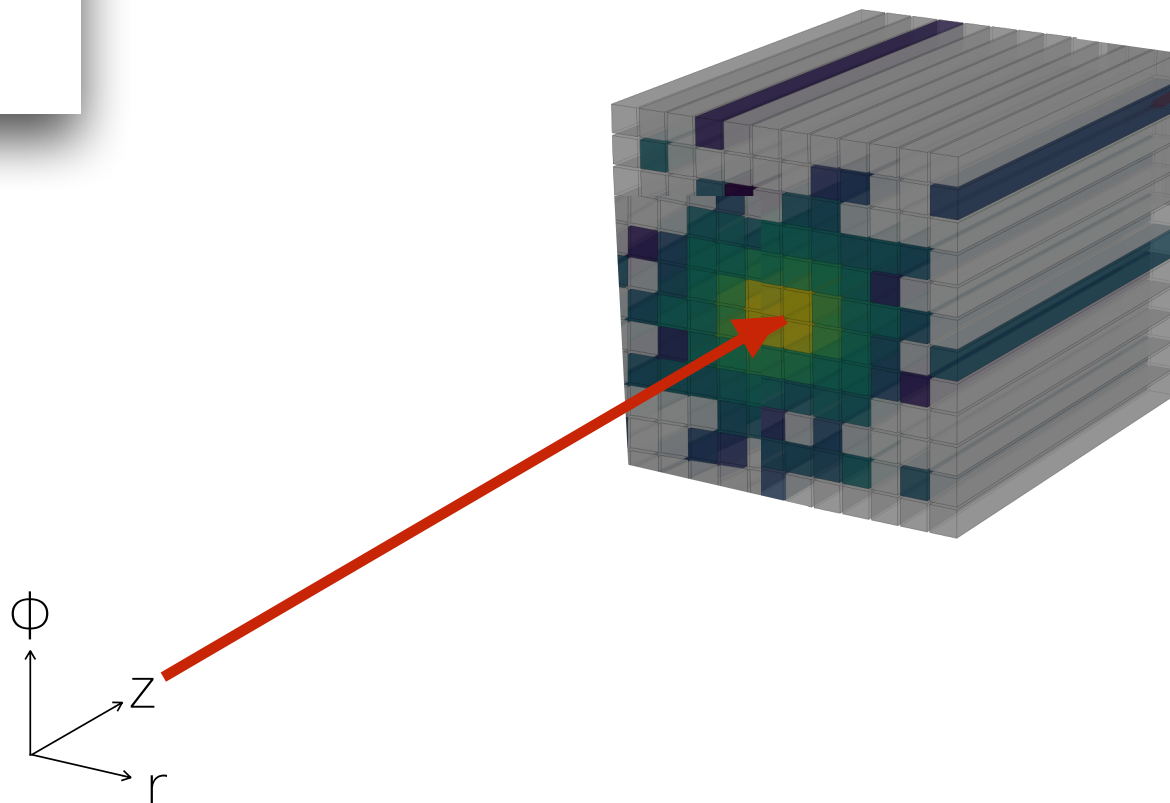
# Accelerating Detector Simulations

40



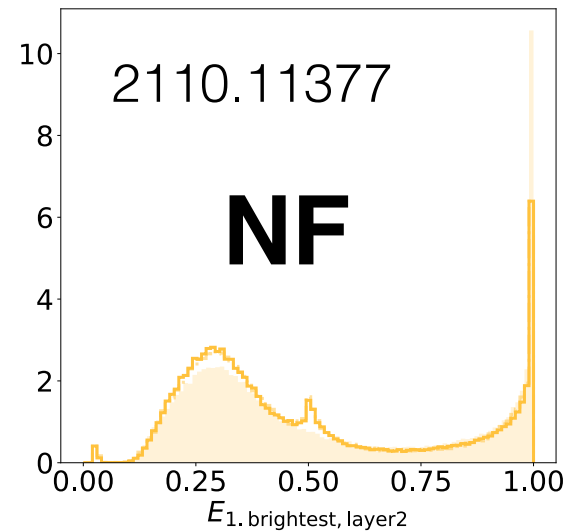
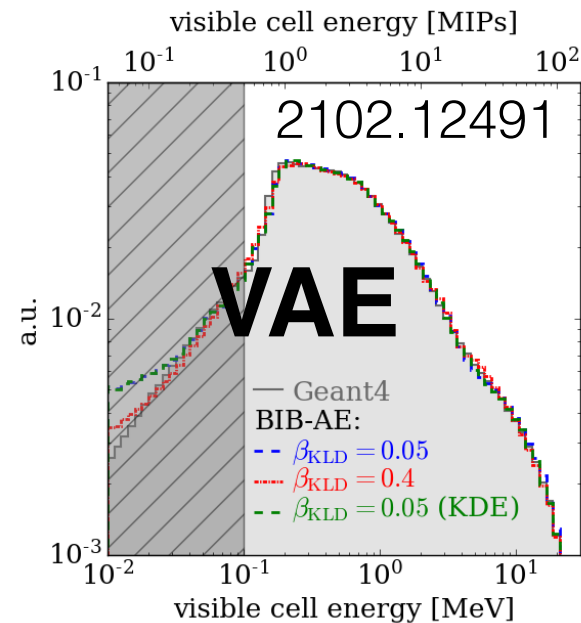
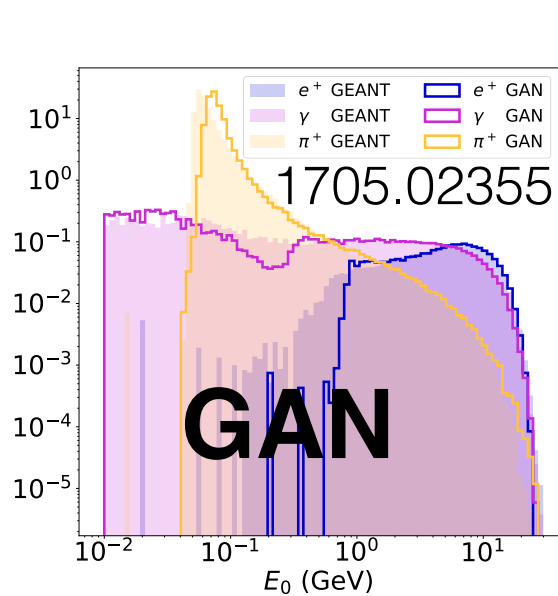
Calorimeters are often the slowest to simulate  
*stopping particles requires simulating interactions of all energies*

Grayscale images:  
Pixel intensity =  
energy deposited

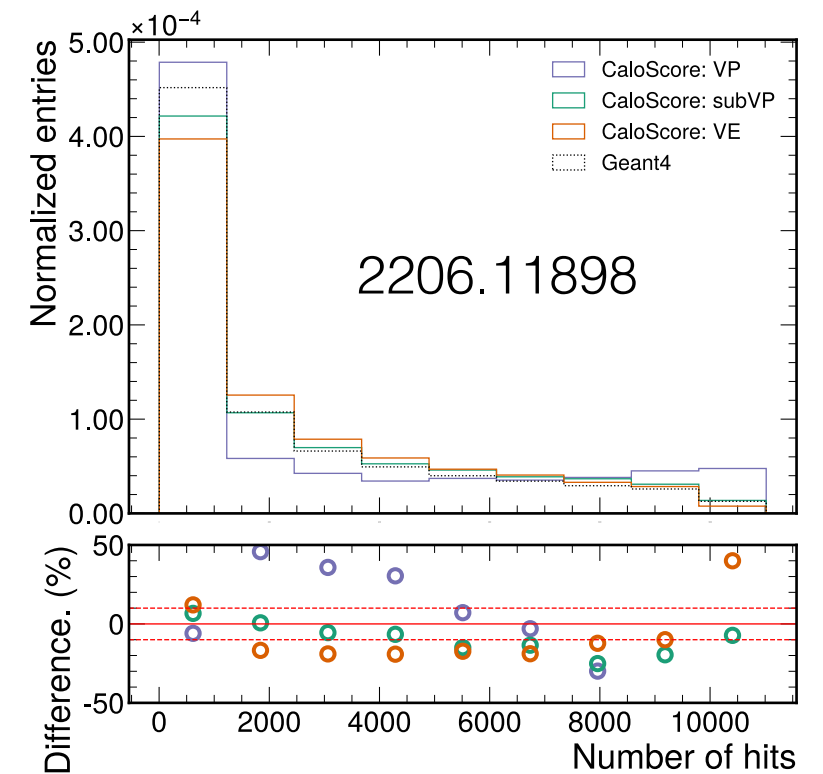
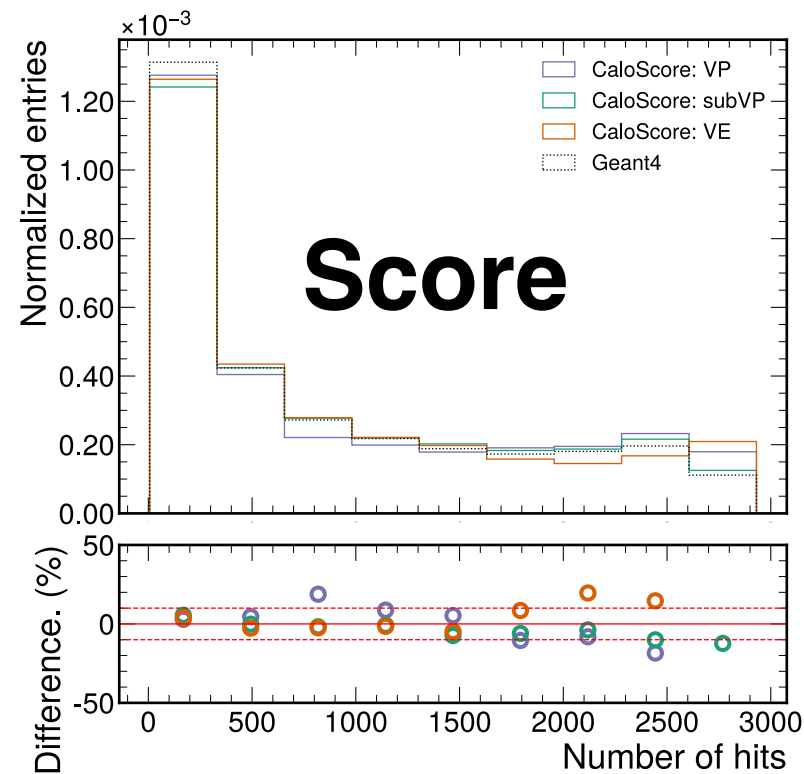
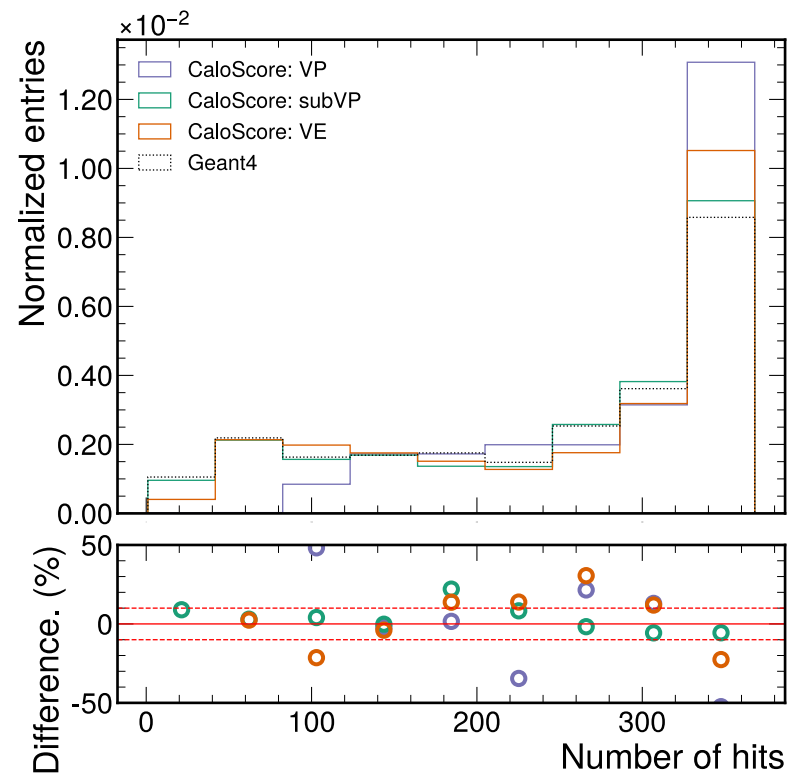


# Calorimeter ML Surrogate Models

41



Many papers on this subject - see the living review for all

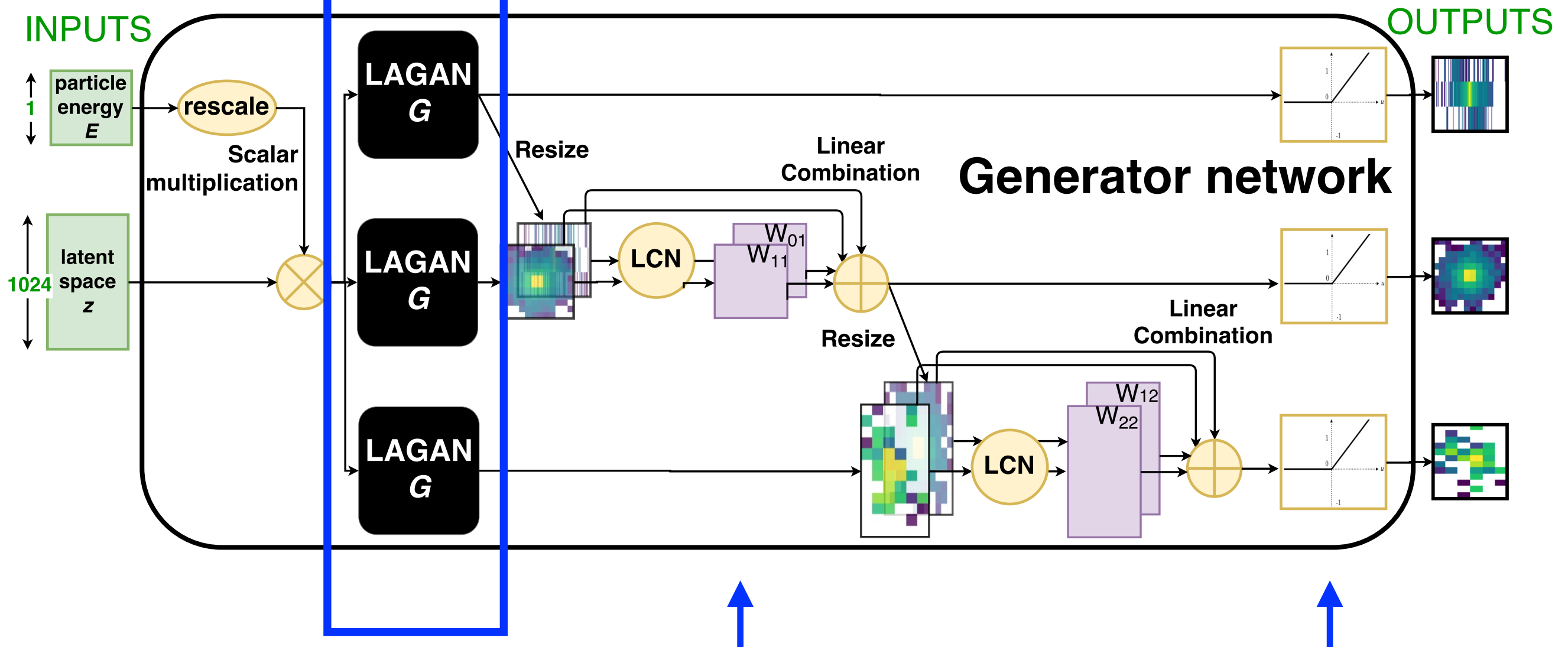


See also <https://calochallenge.github.io/homepage/>

# Introducing CaloGAN

One image per  
calo layer

One network per particle type;  
input particle energy



LA = Locally  
Aware, like a CNN

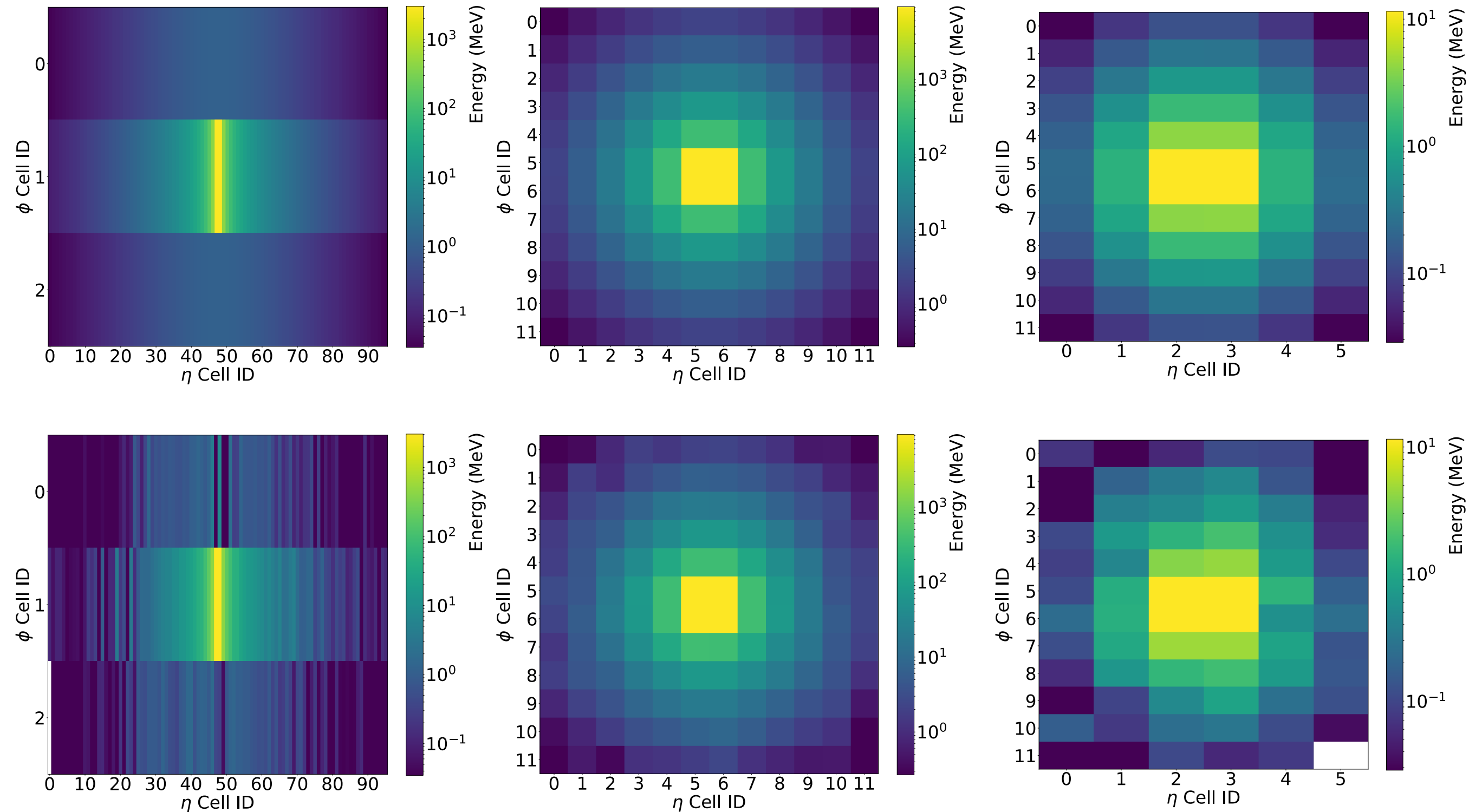
use layer  $i$  as  
input to layer  $i+1$

ReLU to  
encourage sparsity



# Performance: average images

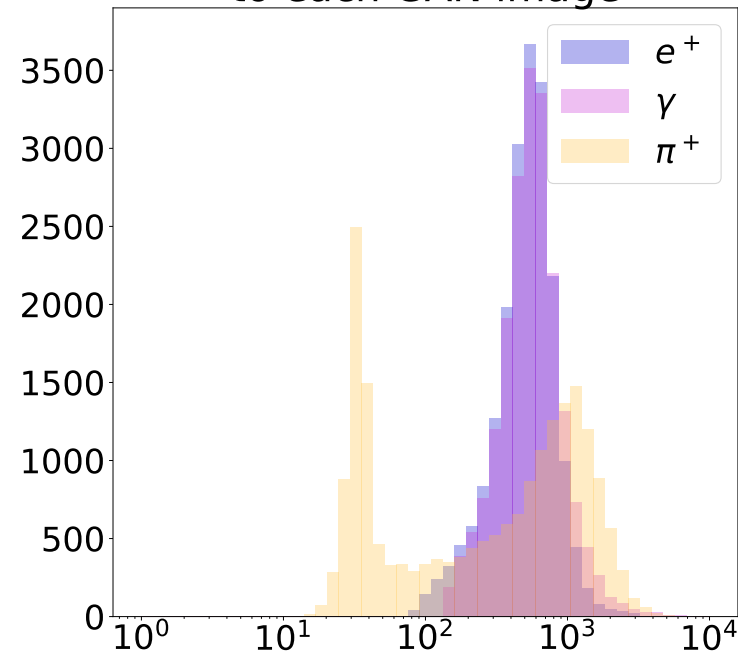
## Geant4



## CaloGAN

# Test for “overtraining”

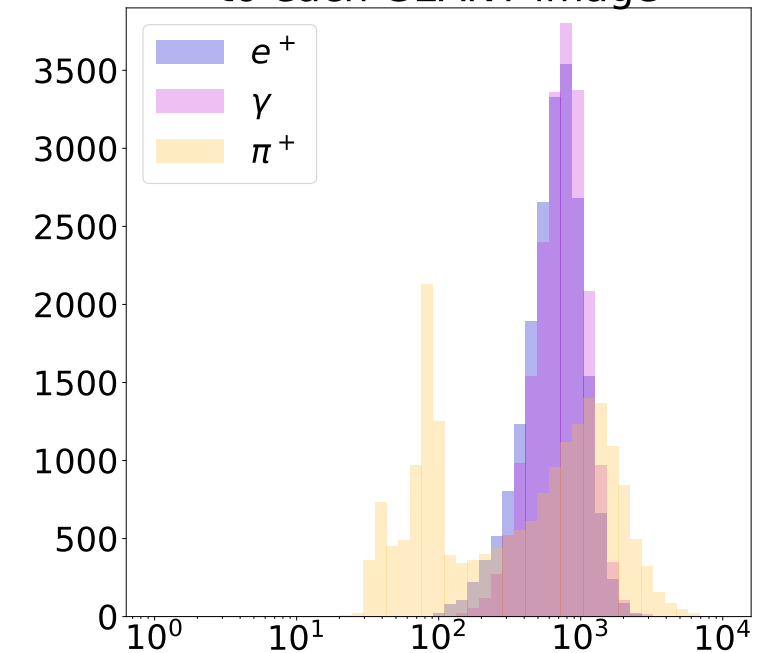
Nearest GEANT neighbour  
to each GAN image



not  
memorizing

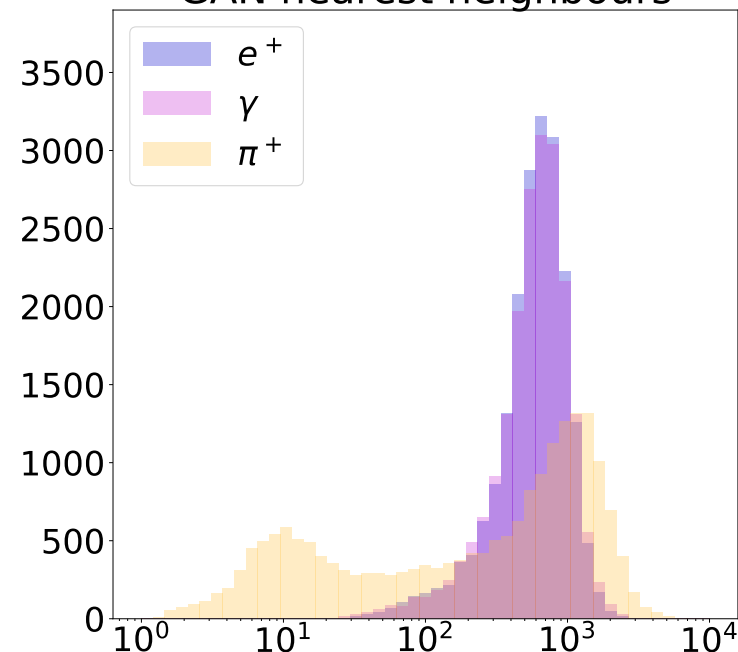


Nearest GAN neighbour  
to each GEANT image



A key challenge in training GANs is the diversity of generated images.  
This does not seem to be a problem for CaloGAN.

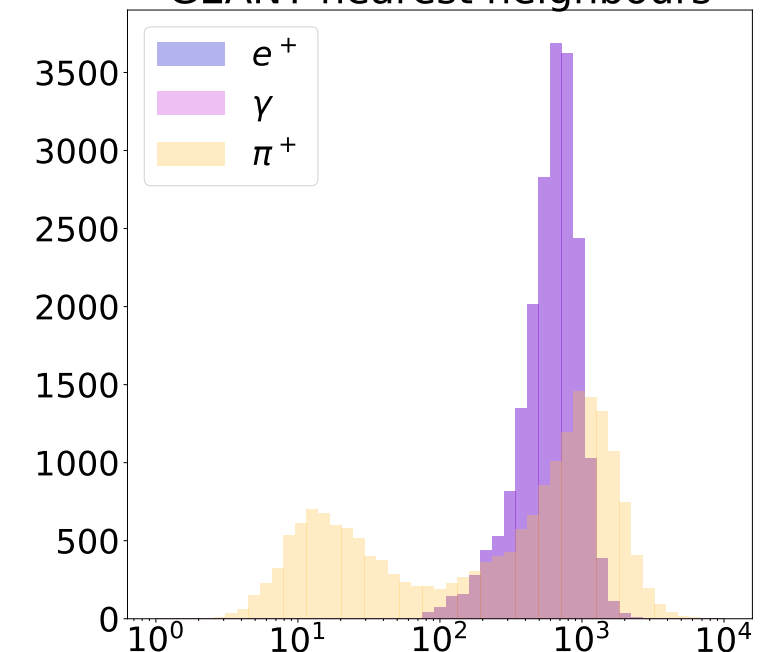
GAN nearest neighbours



no mode  
collapse



GEANT nearest neighbours

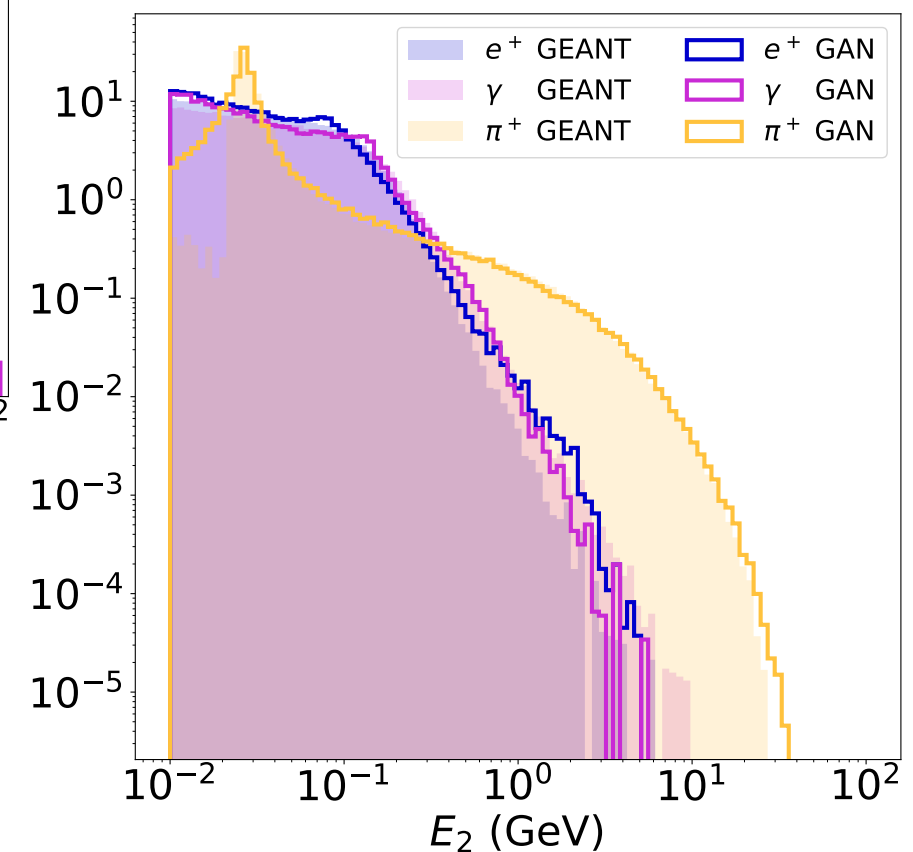
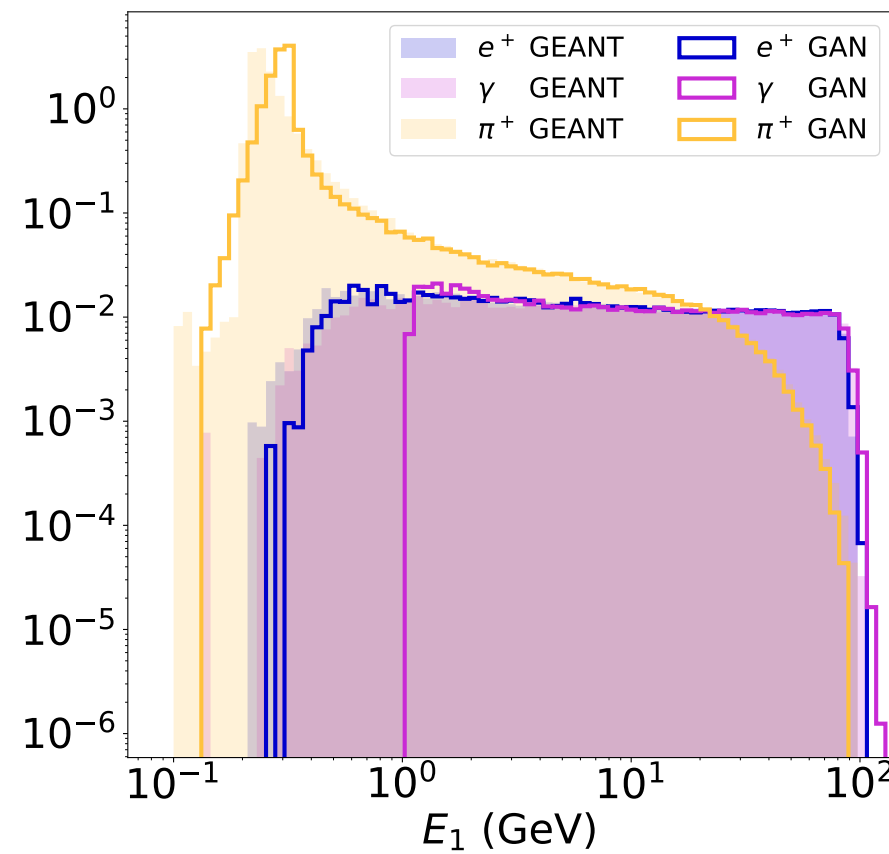
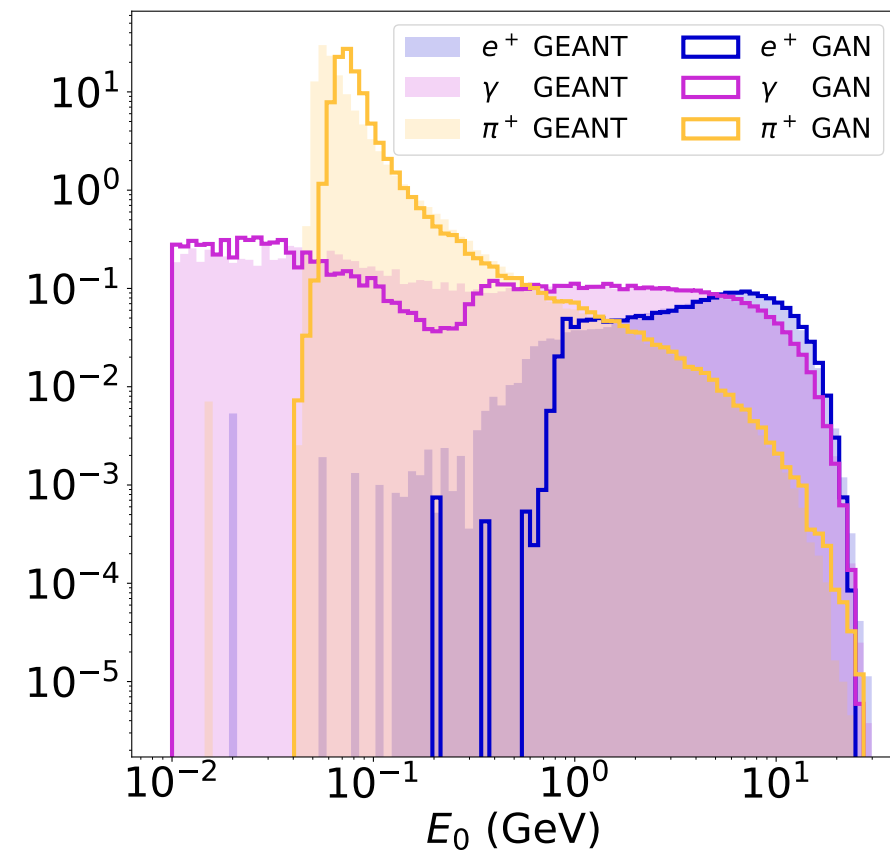






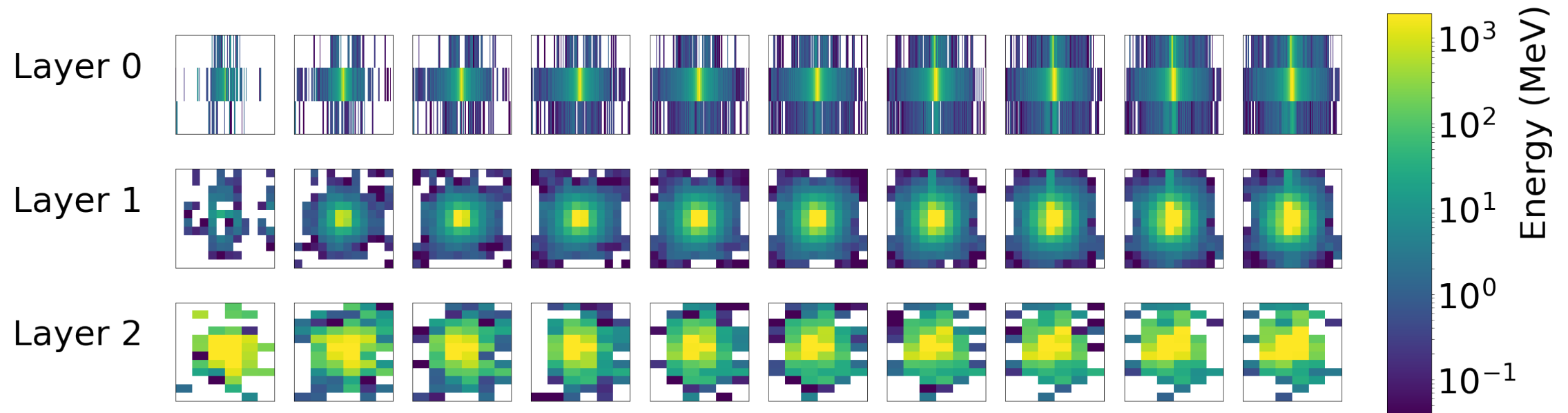
# Performance: energy per layer

Pions deposit much less energy in the first layers; leave the calorimeter with significant energy

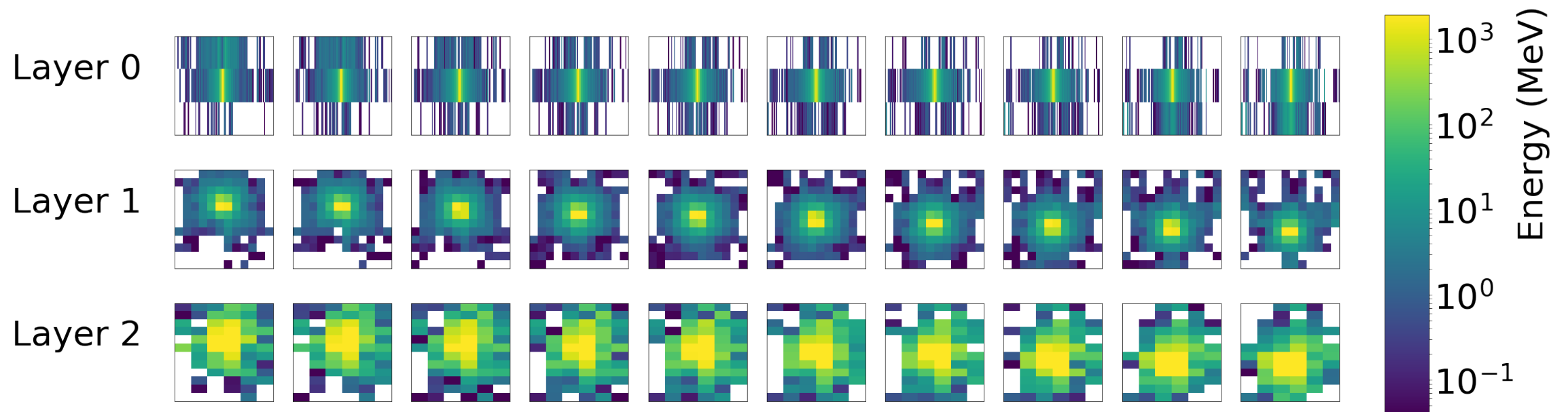


# Conditioning

Fix noise, scan latent variable corresponding to energy



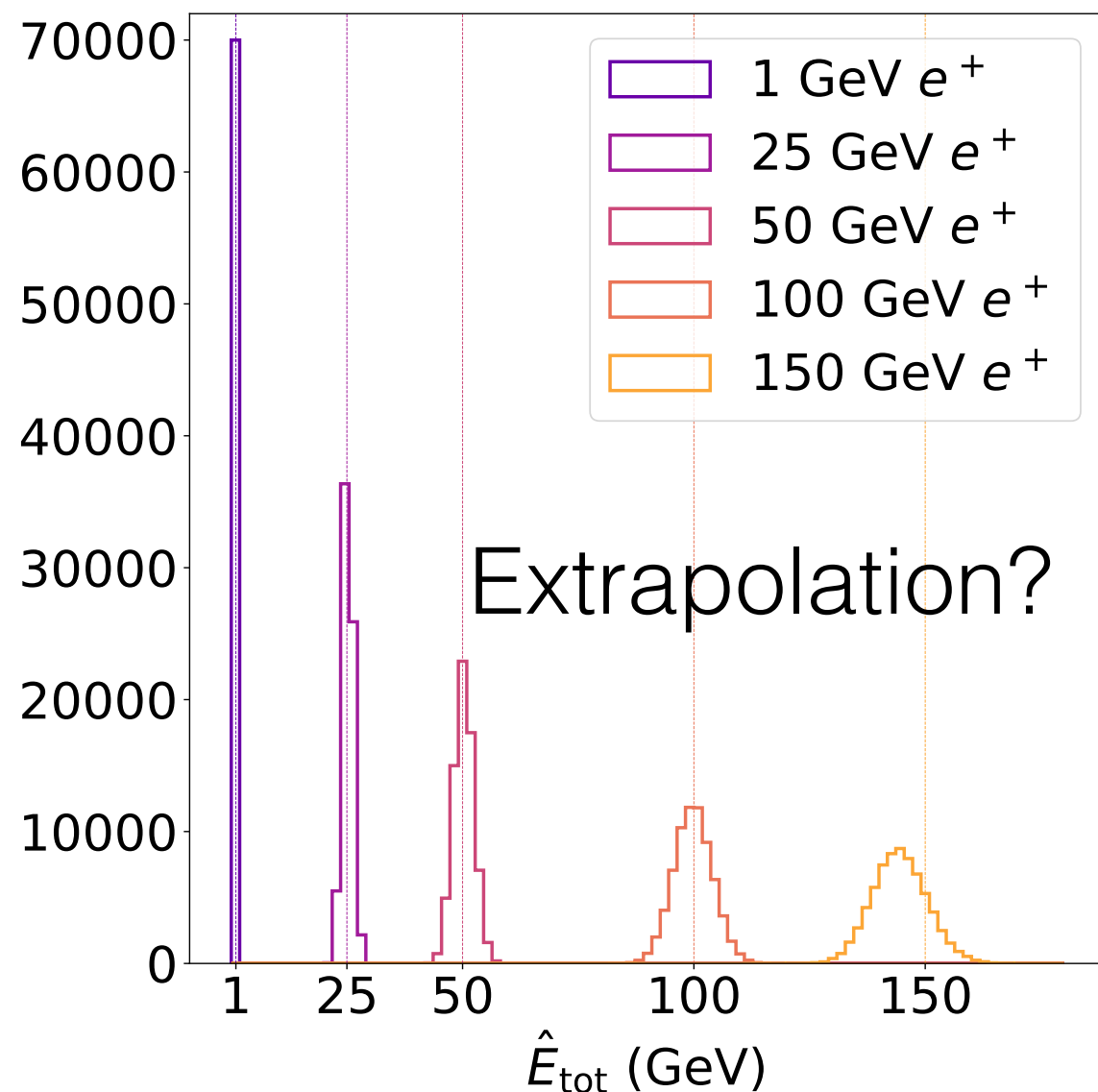
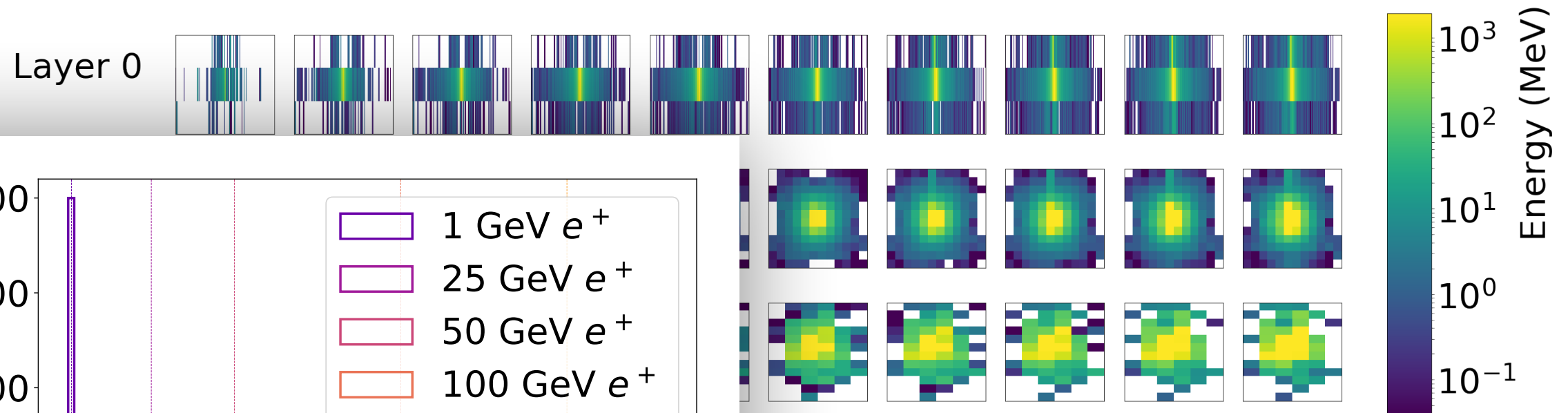
Fix noise, scan latent variable corresponding to x-position



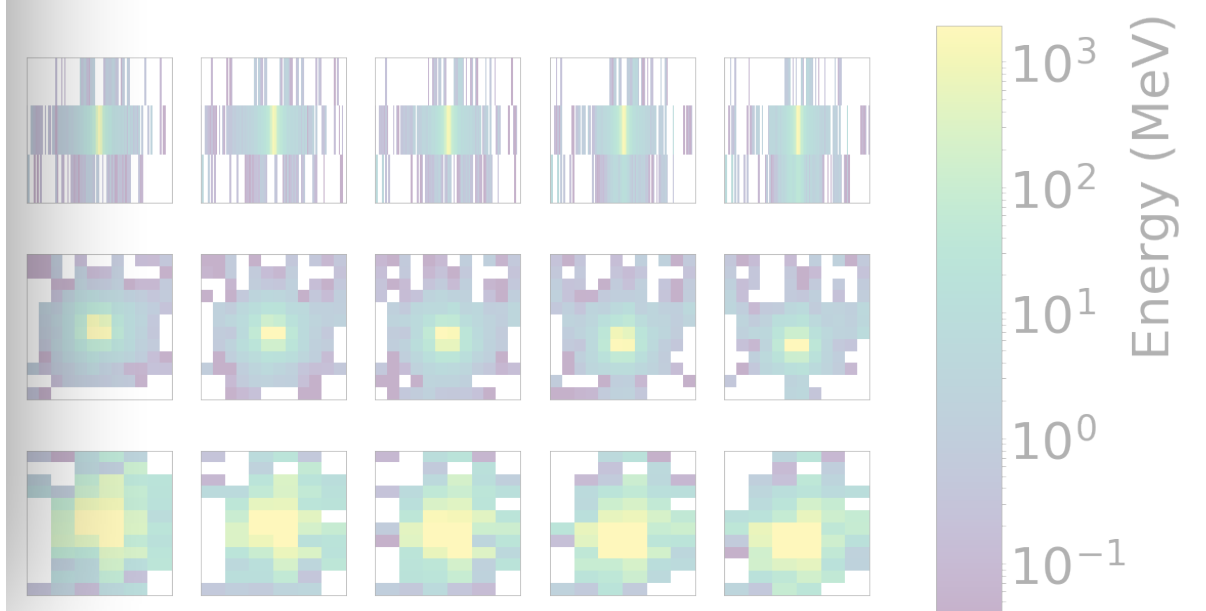


# Conditioning

Fix noise, scan latent variable corresponding to energy



... corresponding to x-position



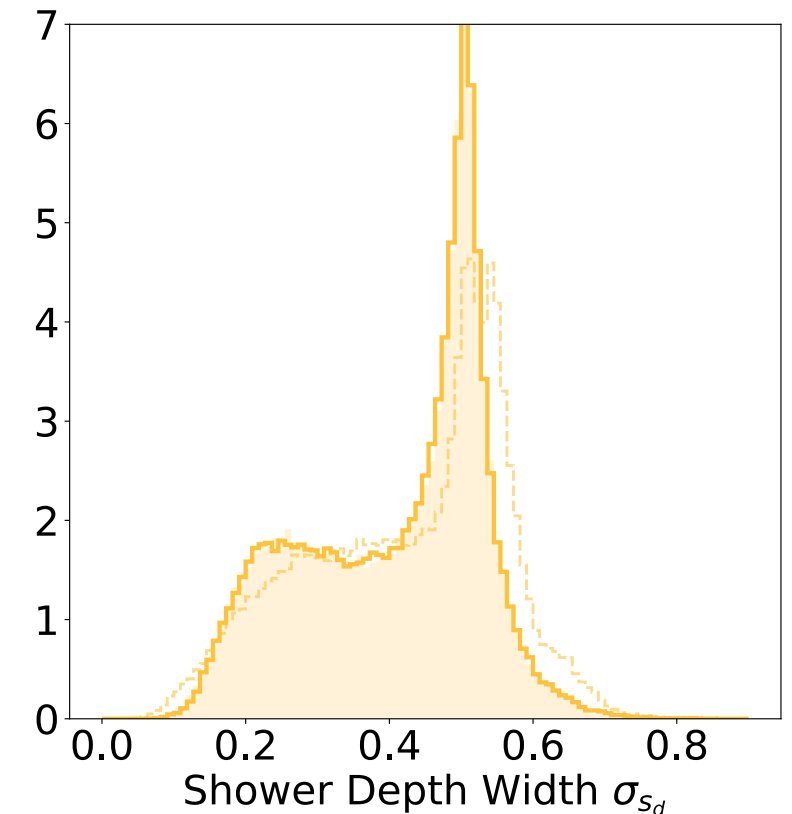
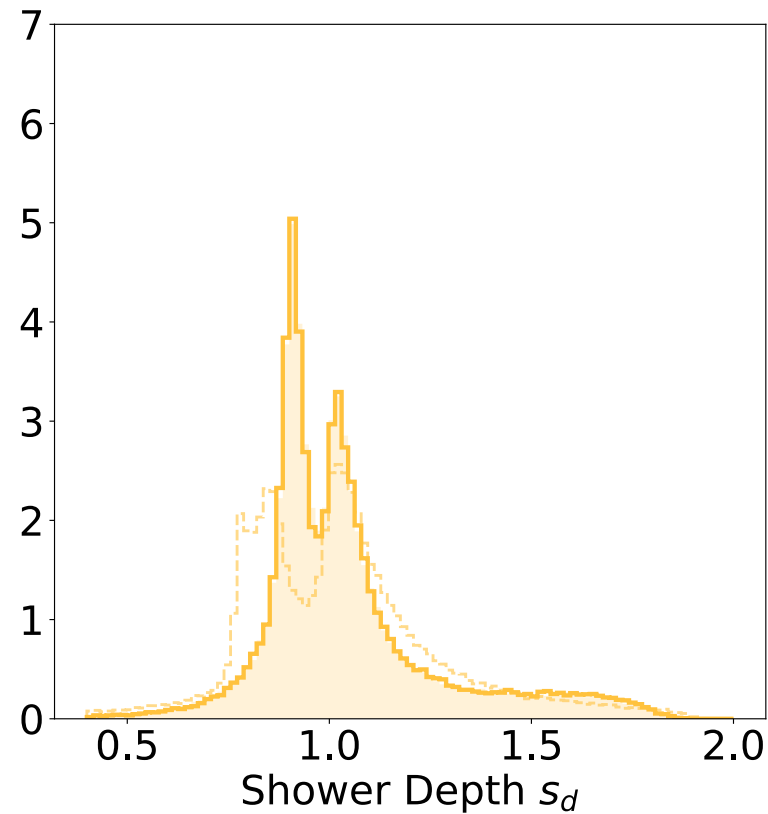
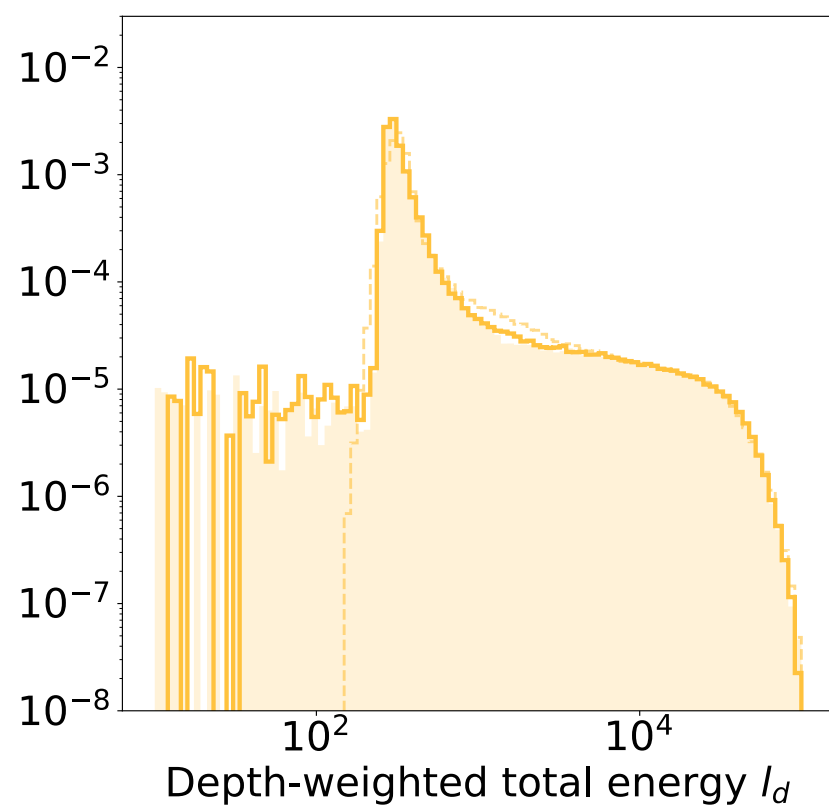
# Timing

Generation Method	Hardware	Batch Size	milliseconds/shower
GEANT4	CPU	N/A	1772 ←
CALOGAN	CPU <i>Intel Xeon E5-2670</i>	1	13.1
		10	5.11
		128	2.19
		1024	2.03
	GPU <i>NVIDIA K80</i>	1	14.5
		4	3.68
		128	0.021
		512	0.014
		1024	0.012 ←


(clearly these numbers have changed as both technologies have improved - this is simply meant to be qualitative & motivating!)

# Current State of the art

Generative models have gotten much better; **flow models** are particularly promising. Added bonus: have an explicit density.



  $\pi^+$  GEANT

  $\pi^+$  CaloGAN

  $\pi^+$  CaloFlow



# Current State of the art

Generative models have gotten much better: **flow models** are

AUC / JSD		DNN	
		vs. CALOGAN	vs. CALOFlow
$e^+$	unnormalized	1.000(0) / 0.993(1)	0.847(8) / 0.345(12)
	normalized	1.000(0) / 0.997(0)	0.869(2) / 0.376(4)
$\gamma$	unnormalized	1.000(0) / 0.996(1)	0.660(6) / 0.067(4)
	normalized	1.000(0) / 0.994(1)	0.794(4) / 0.213(7)
$\pi^+$	unnormalized	1.000(0) / 0.988(1)	0.632(2) / 0.048(1)
	normalized	1.000(0) / 0.997(0)	0.751(4) / 0.148(4)

Output is nearly indistinguishable from Geant4 !

*AUC = 1 means easily distinguishable, AUC = 0.5 means not distinguishable*

Depth-weighted total energy  $I_d$

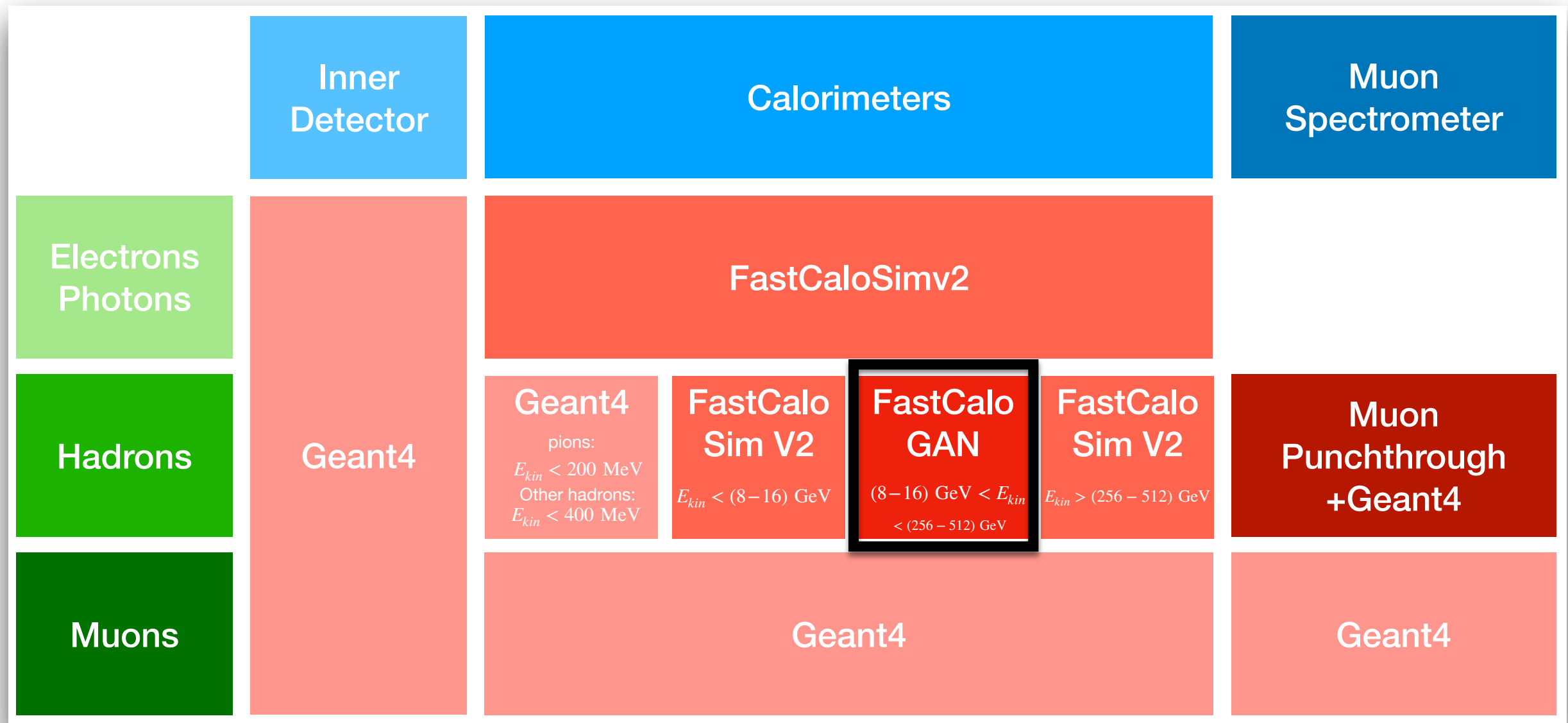
Shower Depth  $s_d$

Shower Depth Width  $\sigma_{s_d}$

   $\pi^+$  GEANT    
    $\pi^+$  CaloGAN    
    $\pi^+$  CaloFlow



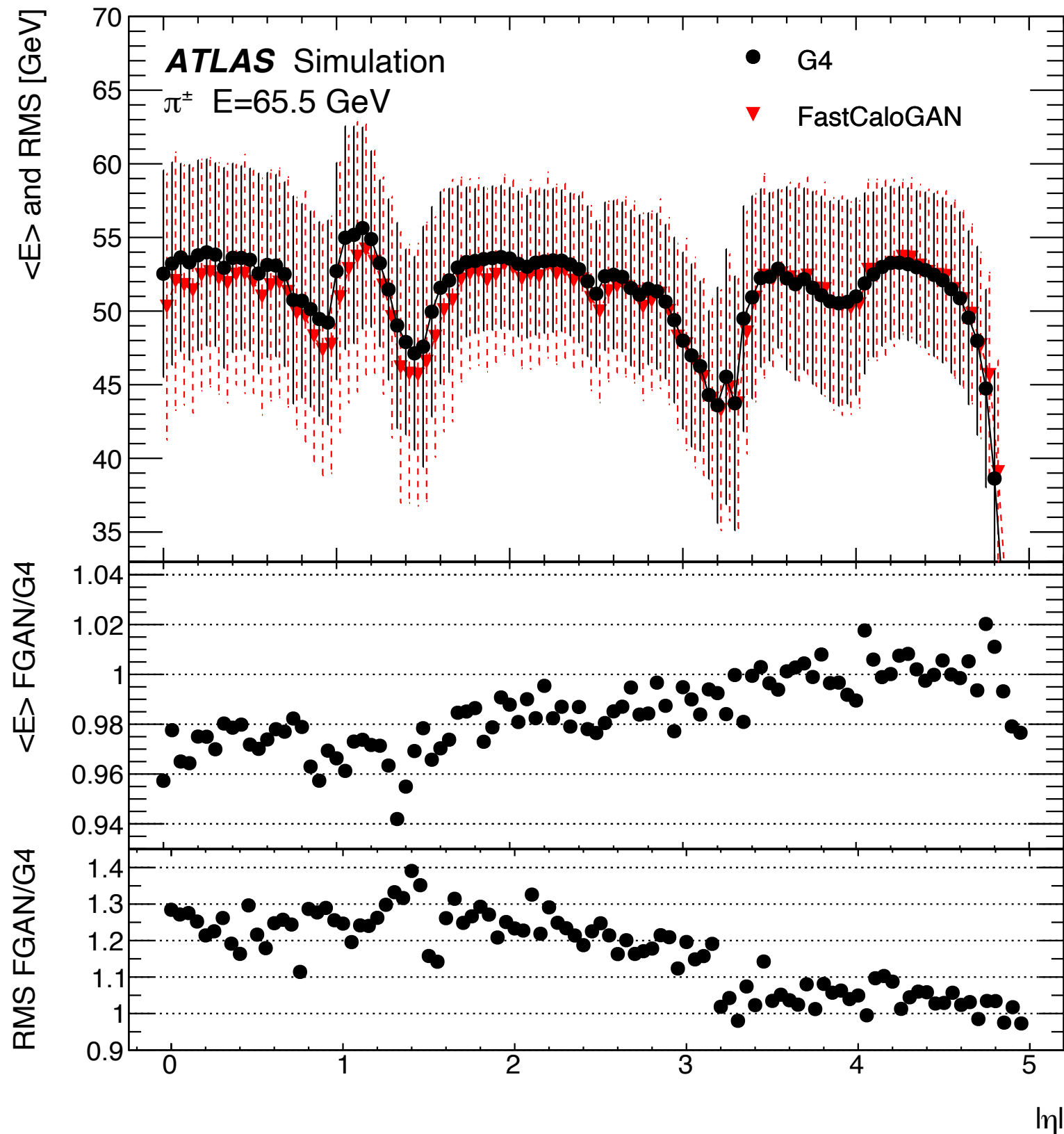
# Integration into real detector sim.



The ATLAS Collaboration fast simulation (AF3) now includes a GAN at intermediate energies for pions



# Integration into real detector sim.

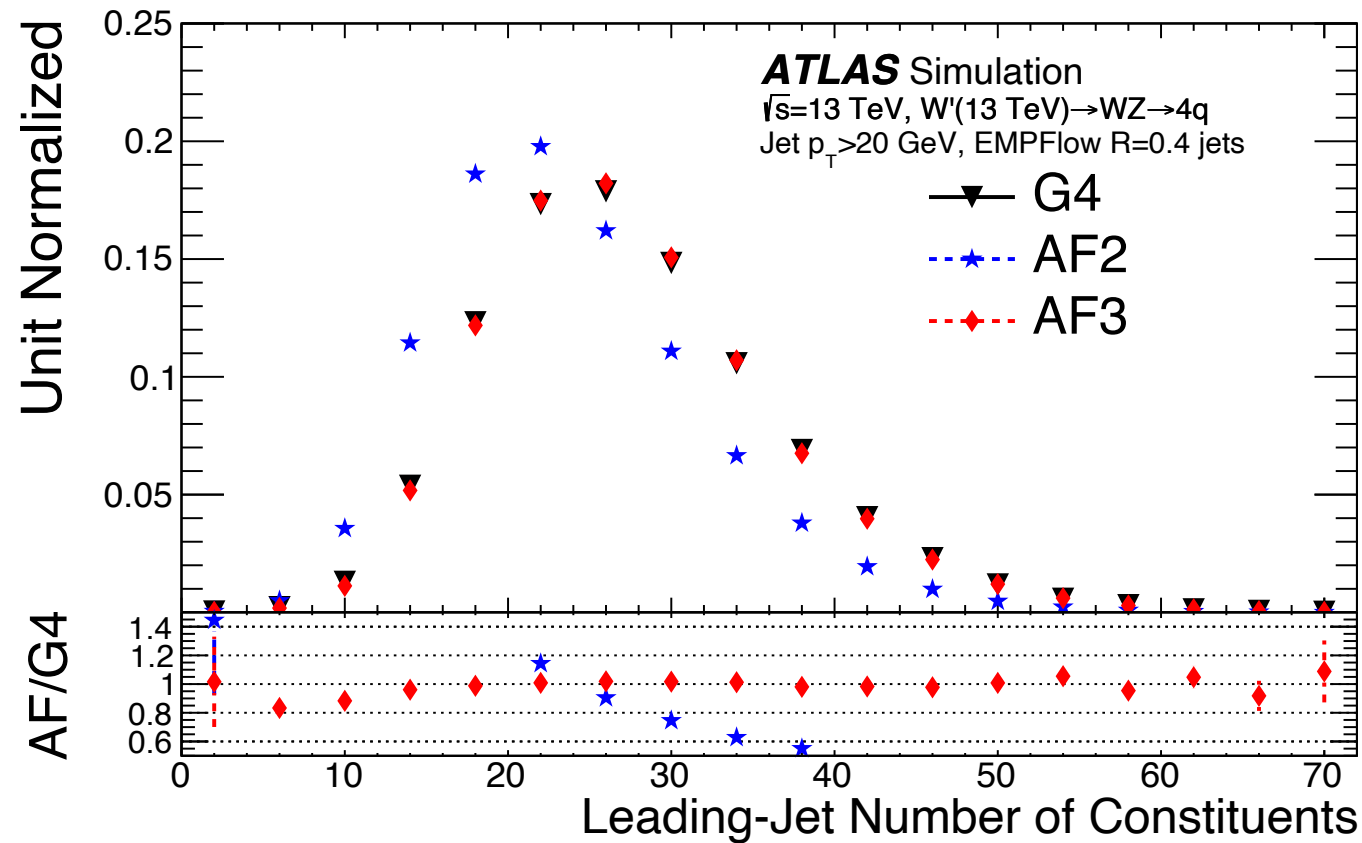


The GAN architecture is relatively simple, but it is able to match the energy scale and resolution well.

There is one GAN per  $\eta$  slice

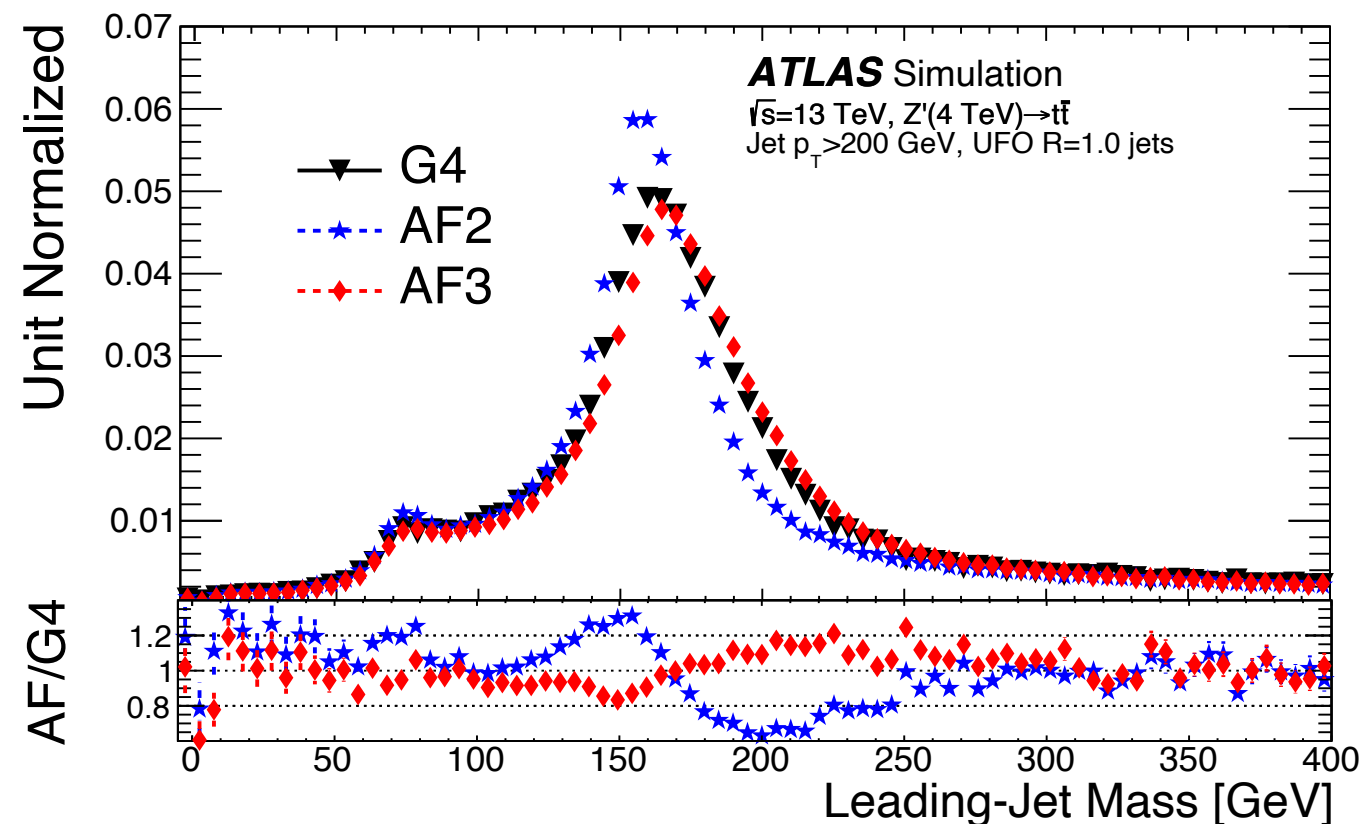


# Integration into real detector sim.

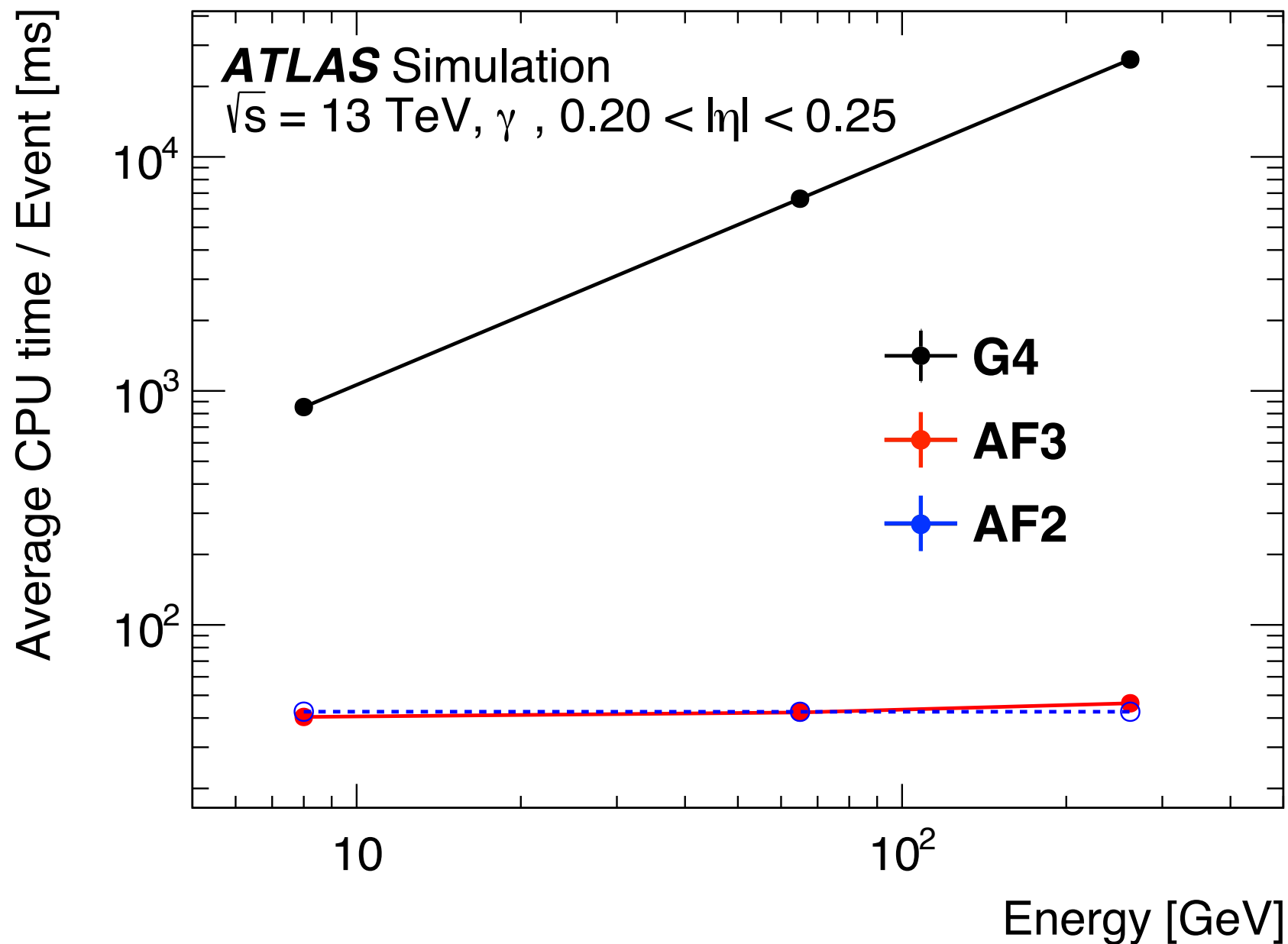


The new fast simulation (**AF3**) significantly improves jet substructure with respect to the older one (**AF2**)

Ideally, the same calibrations derived for full sim. (Geant4-based) can be applied to the fast sim.



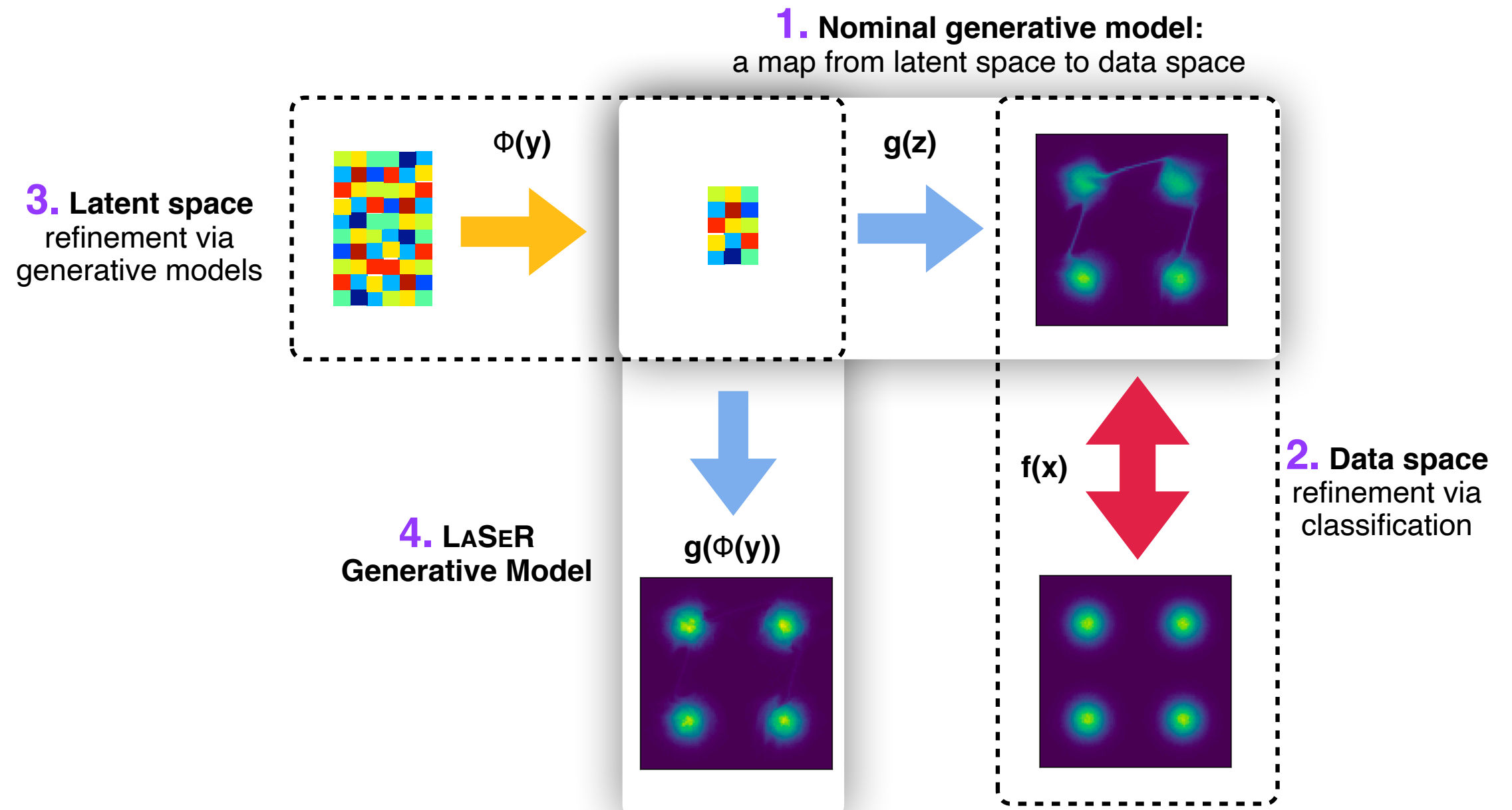
# Integration into real detector sim.



As expected, the fast sim. timing is independent of energy, while Geant4 requires more time for higher energy.

# Refining Simulations

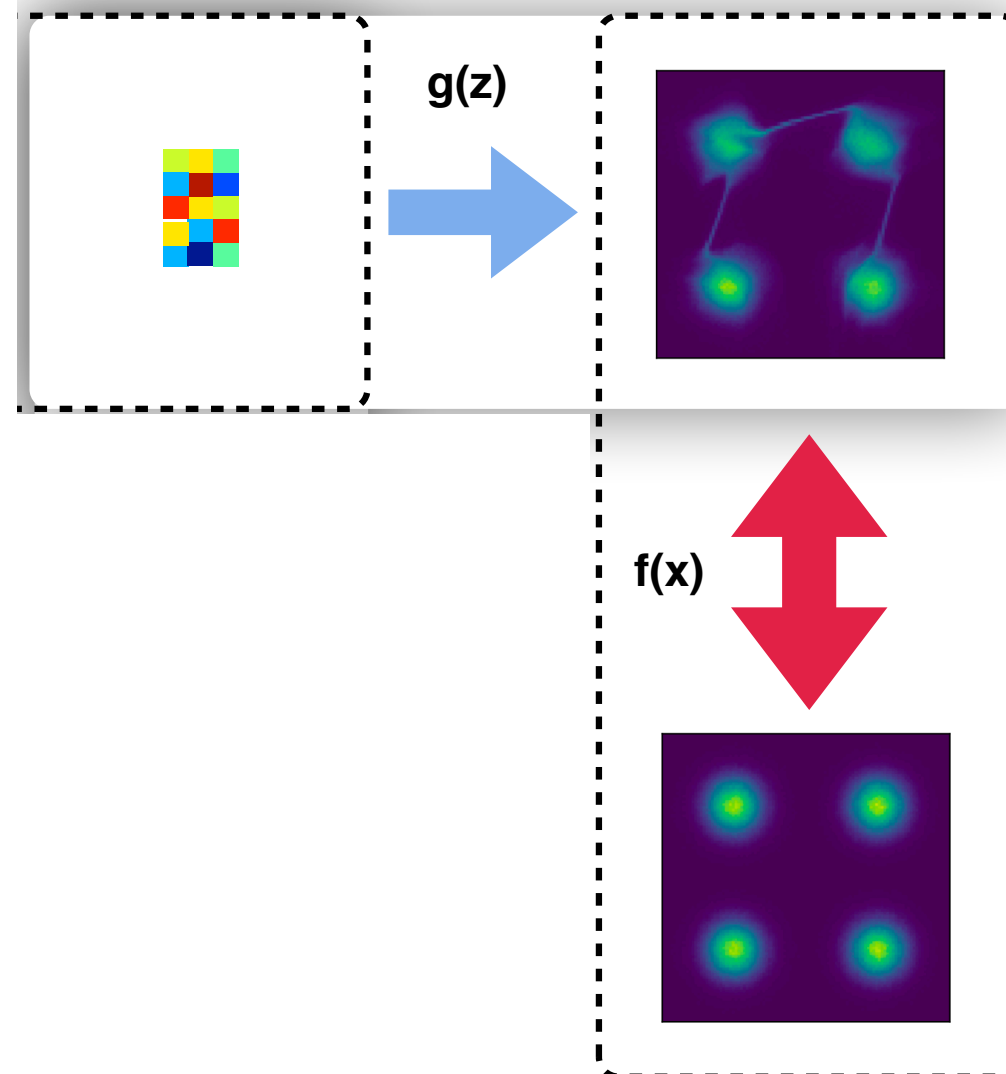
As we move towards precision, we may need to complement primary generative models with post-hoc correction models (e.g. via reweighting)



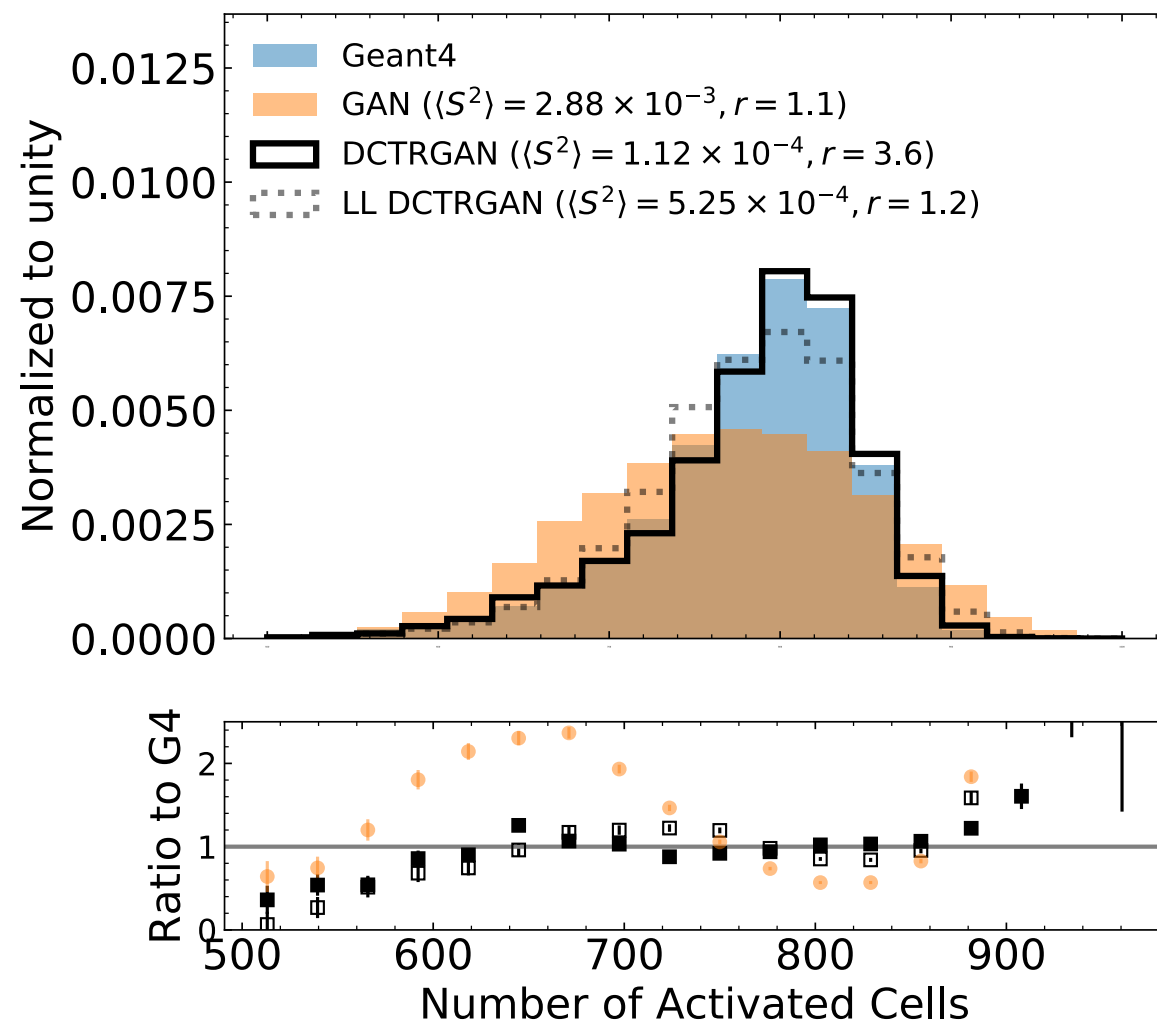
# Refining Simulations

As we move towards precision, we may need to complement primary generative models with post-hoc correction models (e.g. via reweighting)

**1. Nominal generative model:**  
a map from latent space to data space



**2. Data space**  
refinement via  
classification

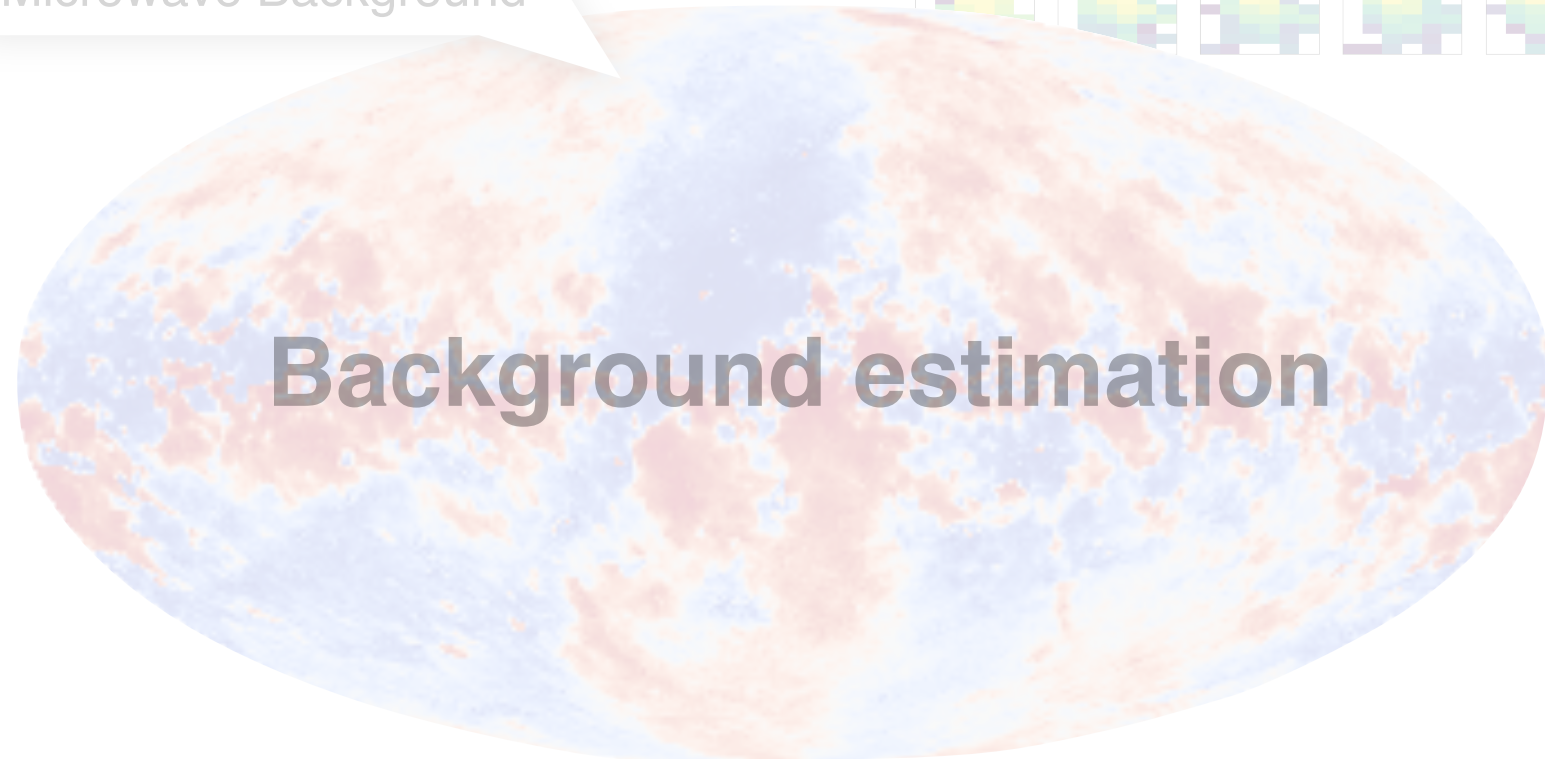


# Generative Models for Particle/Nuclear/Astro



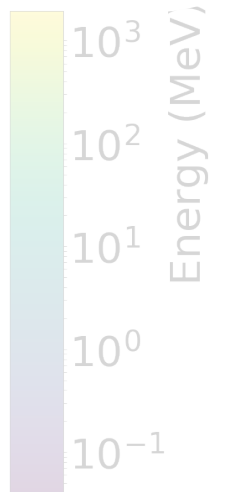
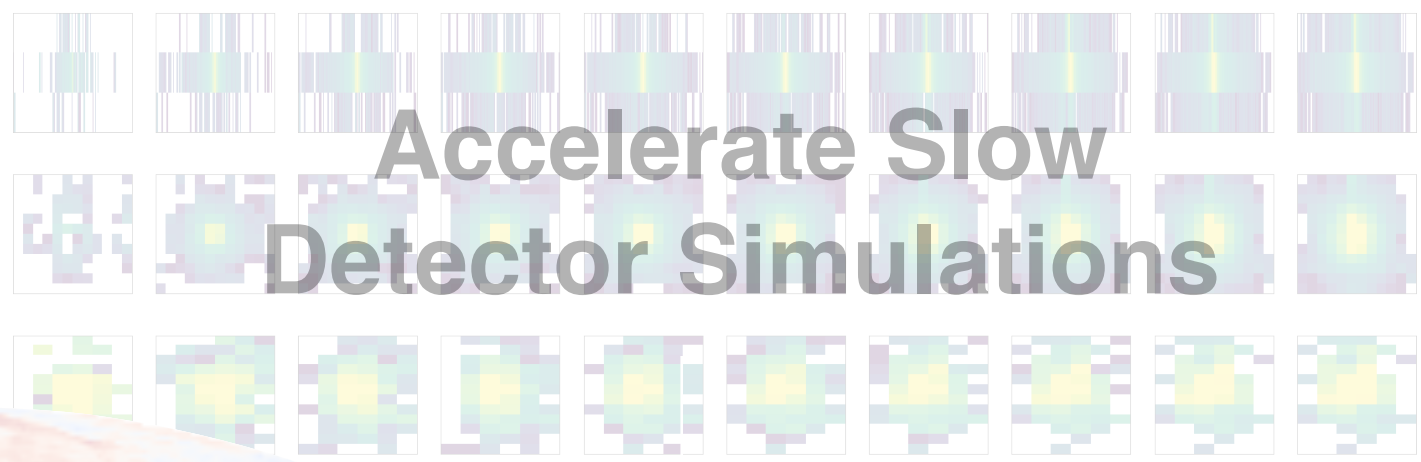
All of these pictures are fake!

Synthetic Galactic radiation for Cosmic Microwave Background



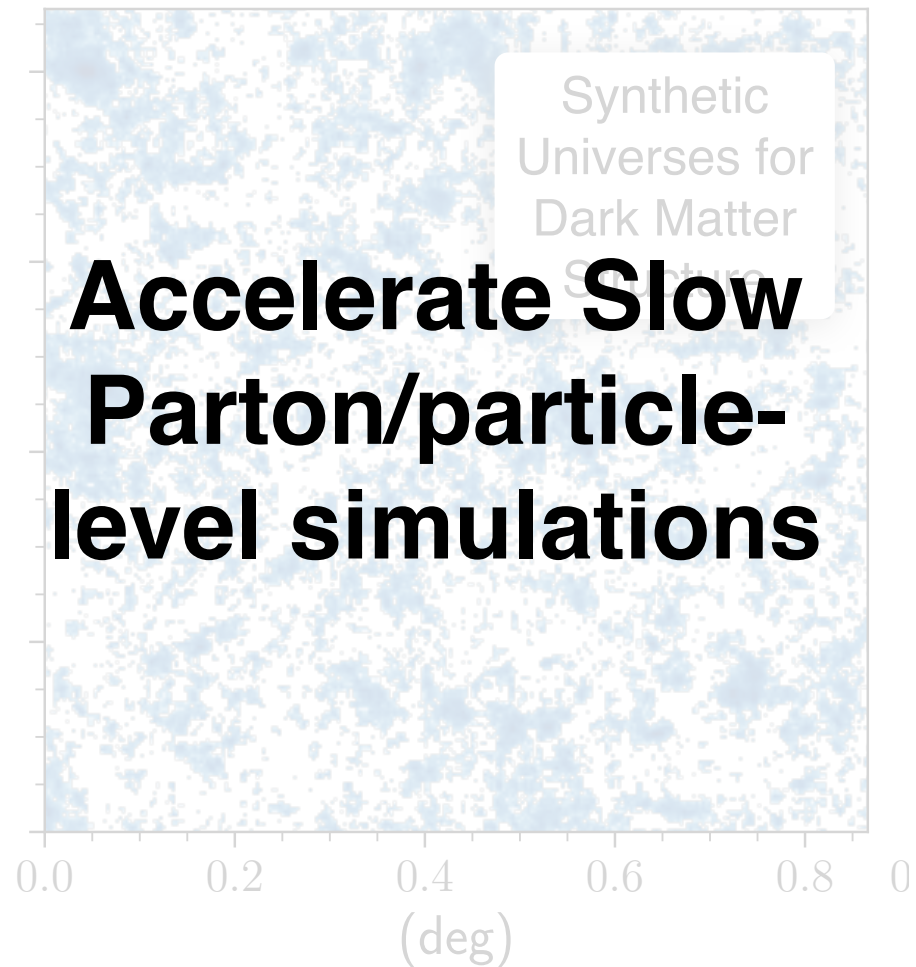
Background estimation

Material Interactions with High Energy Particles



Accelerate Slow  
Detector Simulations

M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

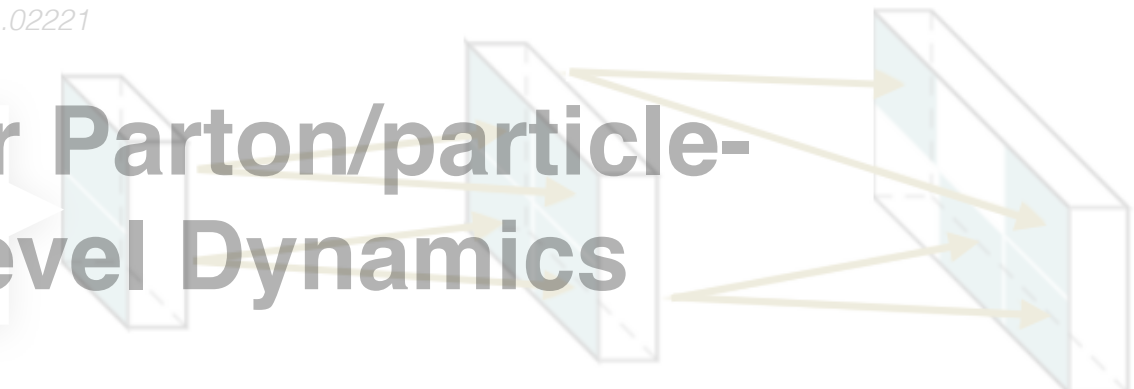


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M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

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Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582

## Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

1701.05927

\*these are just representative examples - see Living Review, 2102.02770

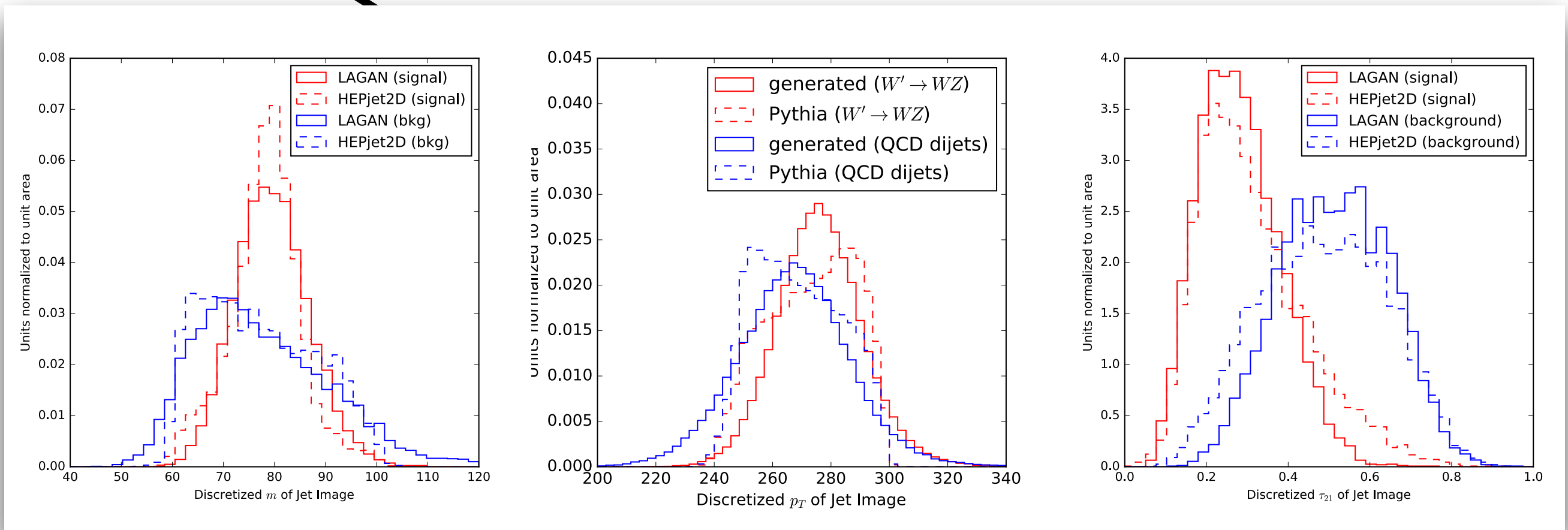
# Accelerating Parton/Particle Sim.\*

59

## Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

1701.05927



LA = Locally aware; somewhere between a DNN and a CNN

**Weight sharing across space**

\*these are just representative examples - see Living Review, 2102.02770



# Accelerating Parton/Particle

Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

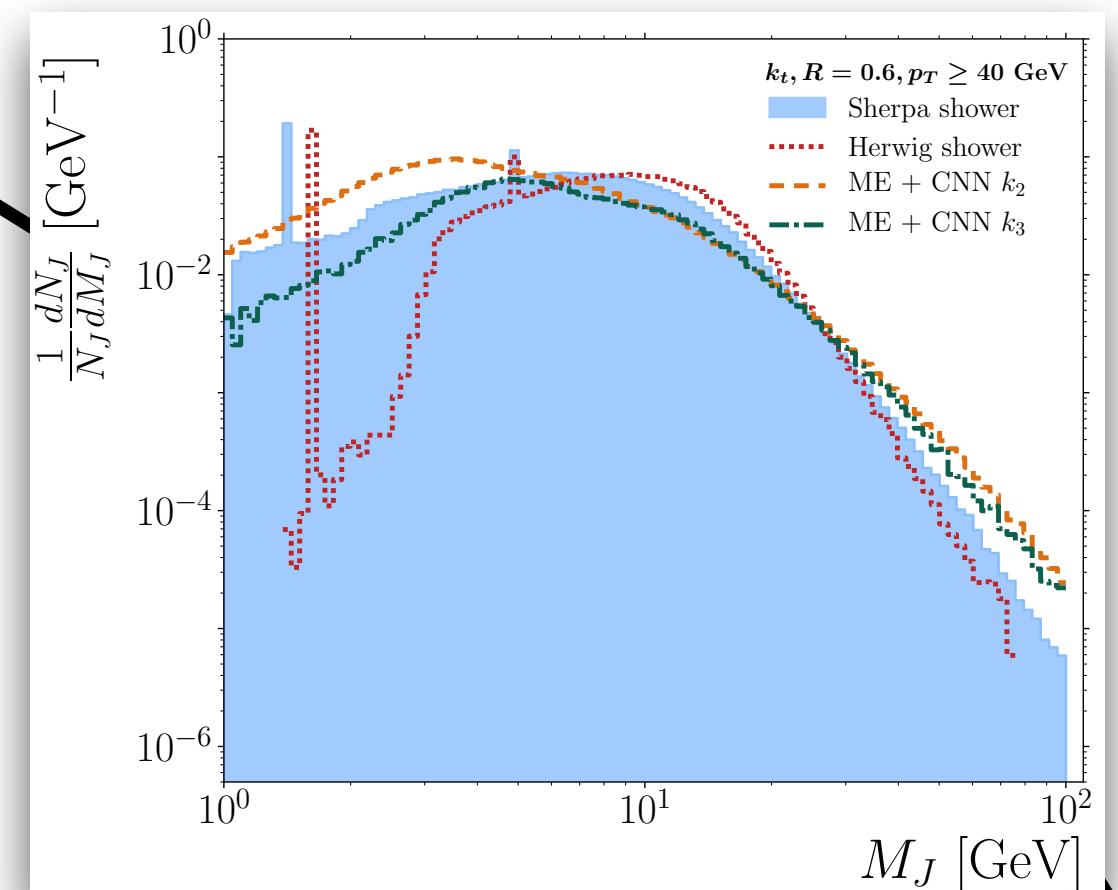
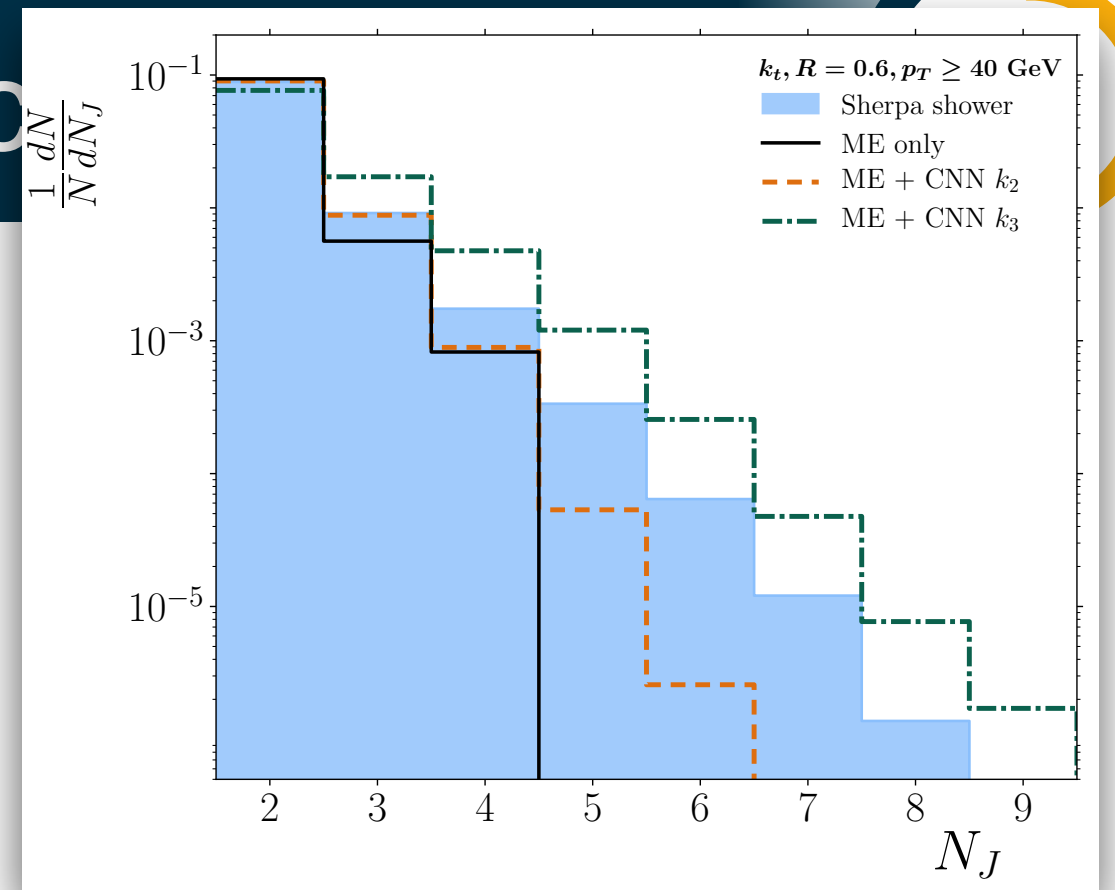
1701.05927

Scale invariant  
images with AEs

J. Monk

1807.03685

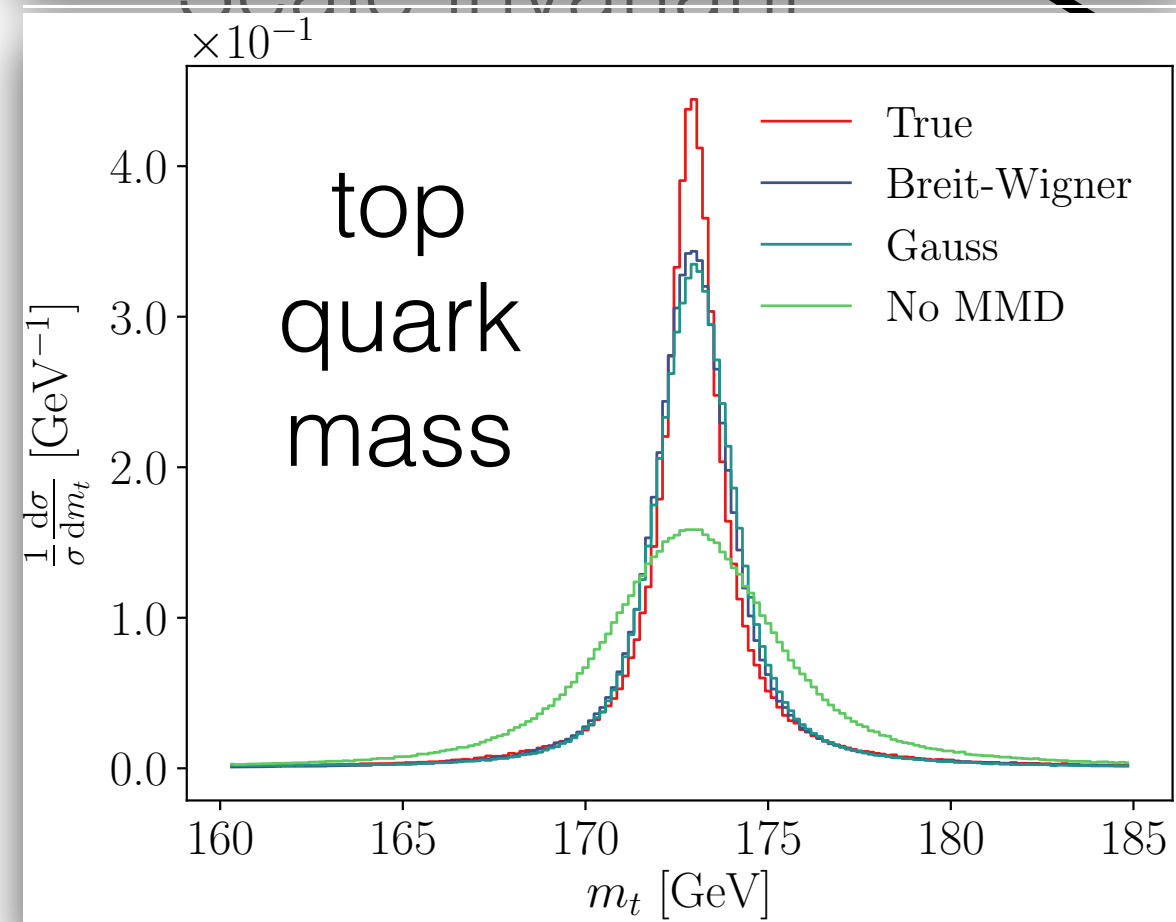
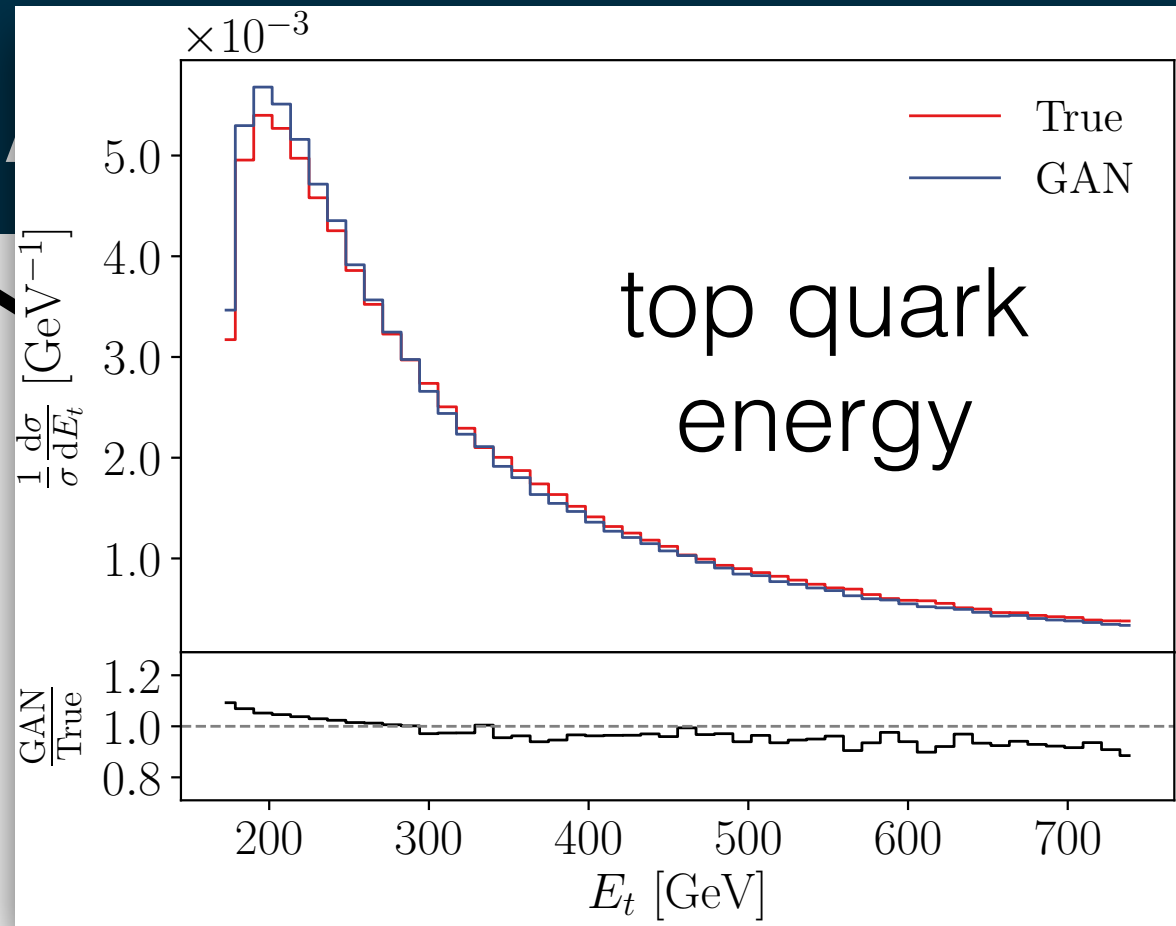
**Weight sharing across space + “time”**



\*these are just representative examples - see Living Review, 2102.02770



# Particle Sim.\*



MMD = maximum mean discrepancy

Fixed number of 4-vectors, allow for intermediate resonances

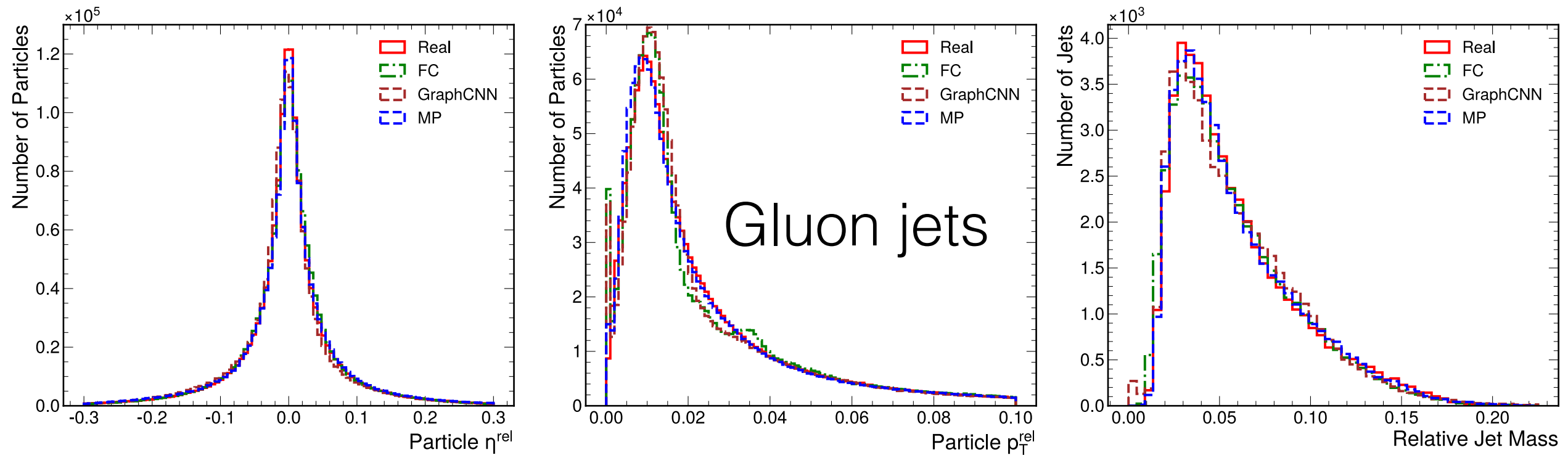
A. Butter, T. Plehn, R. Winterhalder

1907.03764

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M. Paganini, L. de Oliveira, B. Nachman



Variable-length  
output with graphs

R. Kansal et al.

2106.11535

\*these are just representative examples - see Living Review, 2102.02770

# Accelerating Parton/Particle Sim.\*

63

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1701.05927

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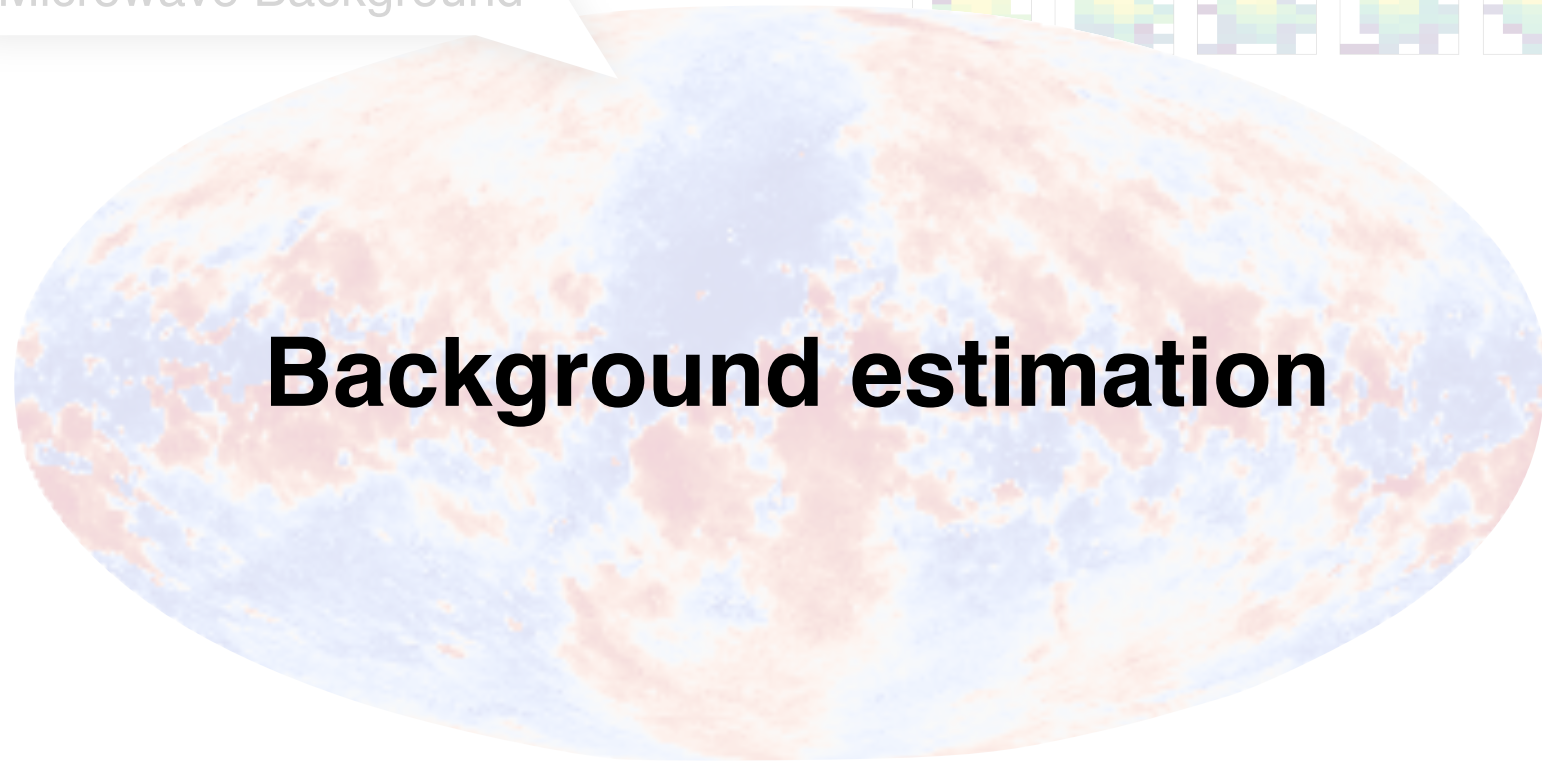
?

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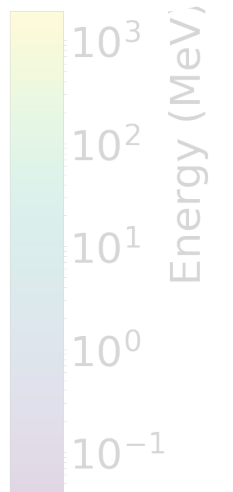
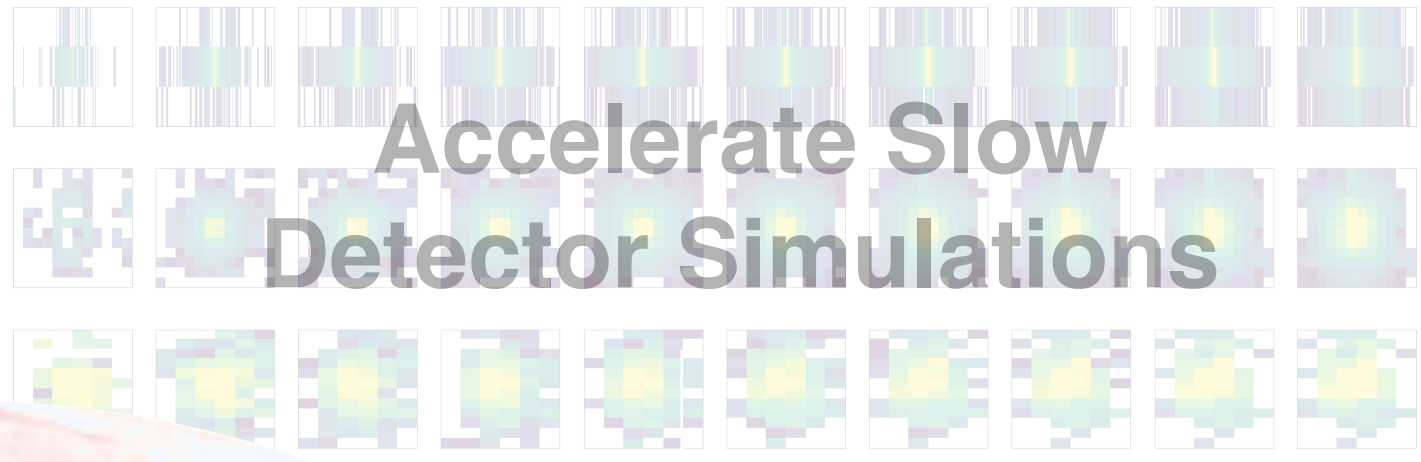
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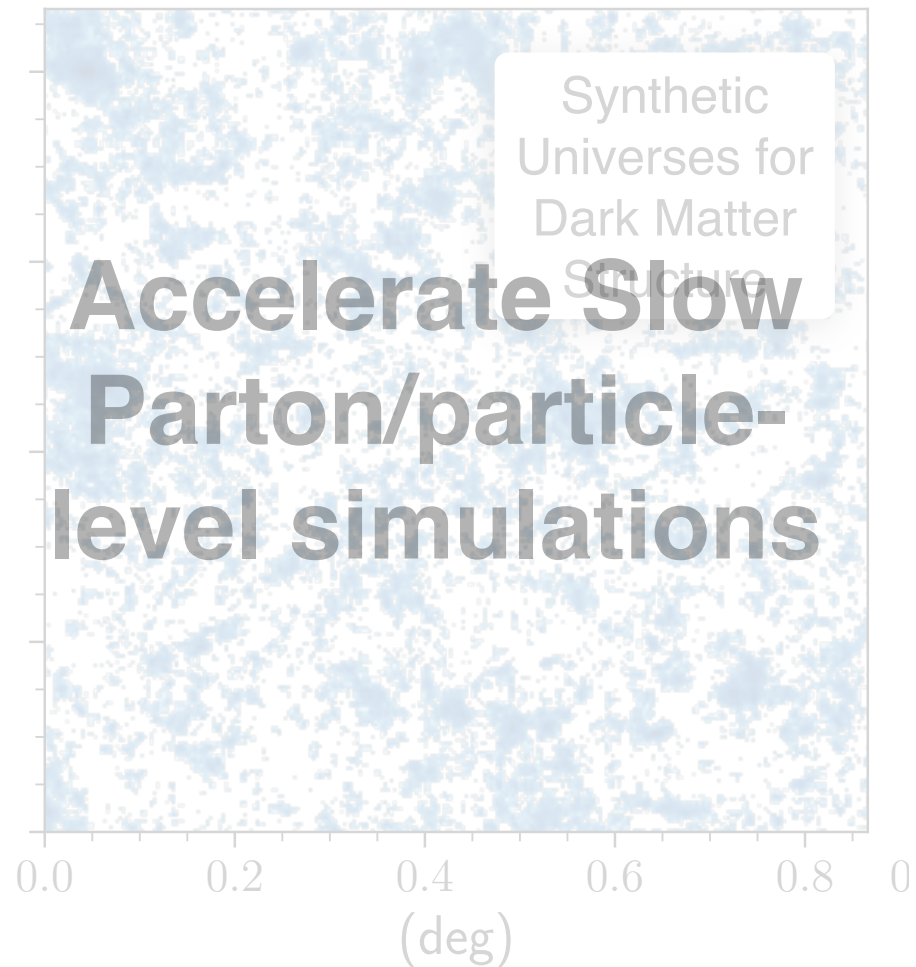
Synthetic Galactic radiation for Cosmic Microwave Background



Material Interactions with High Energy Particles



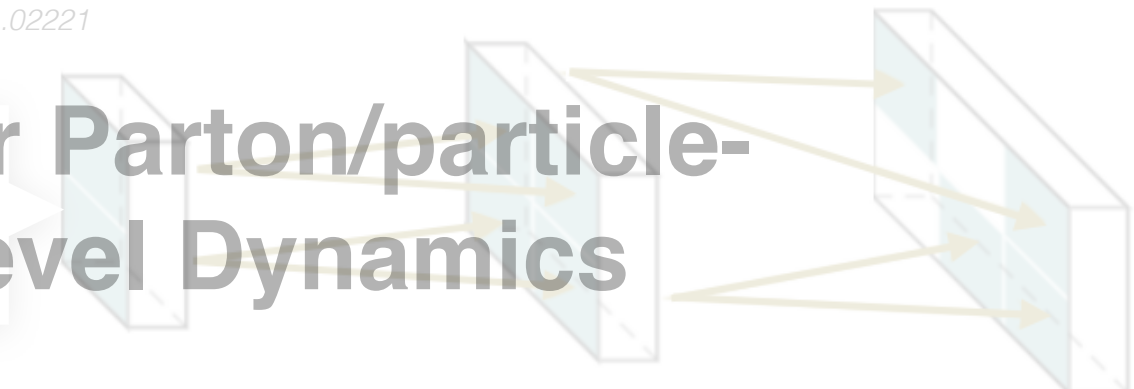
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# Background Estimation

65

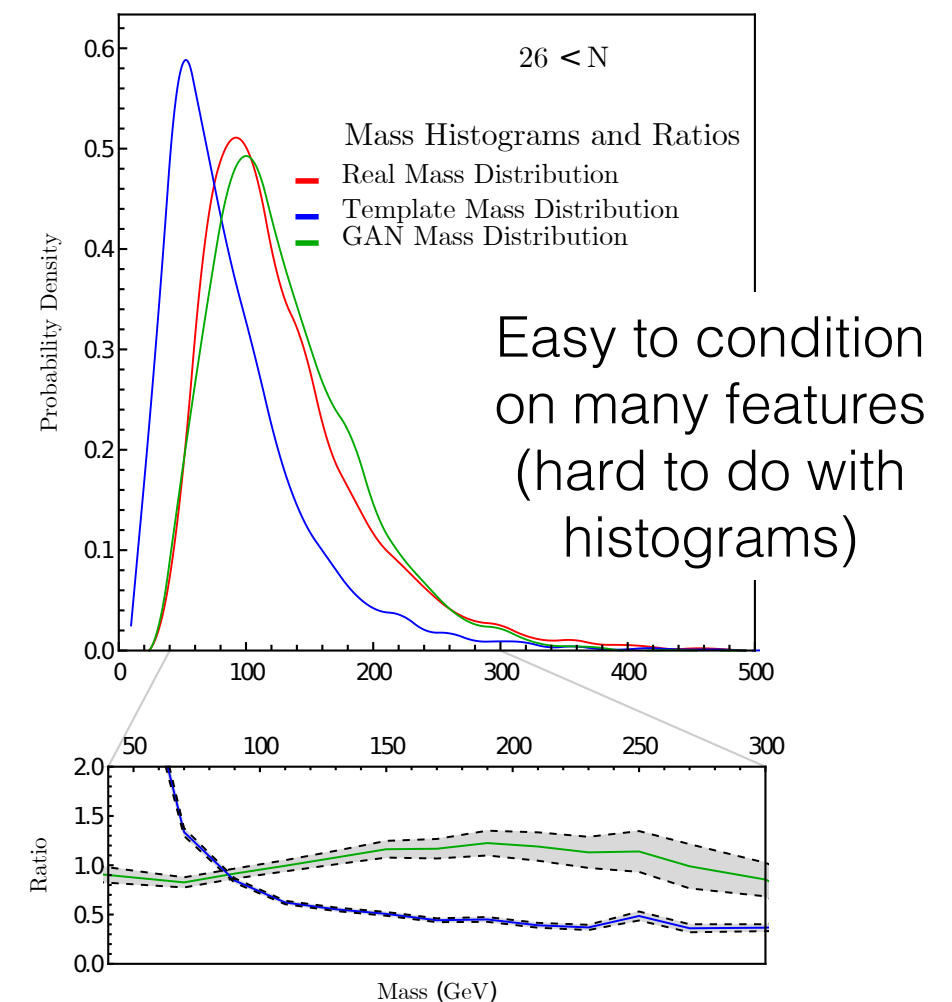
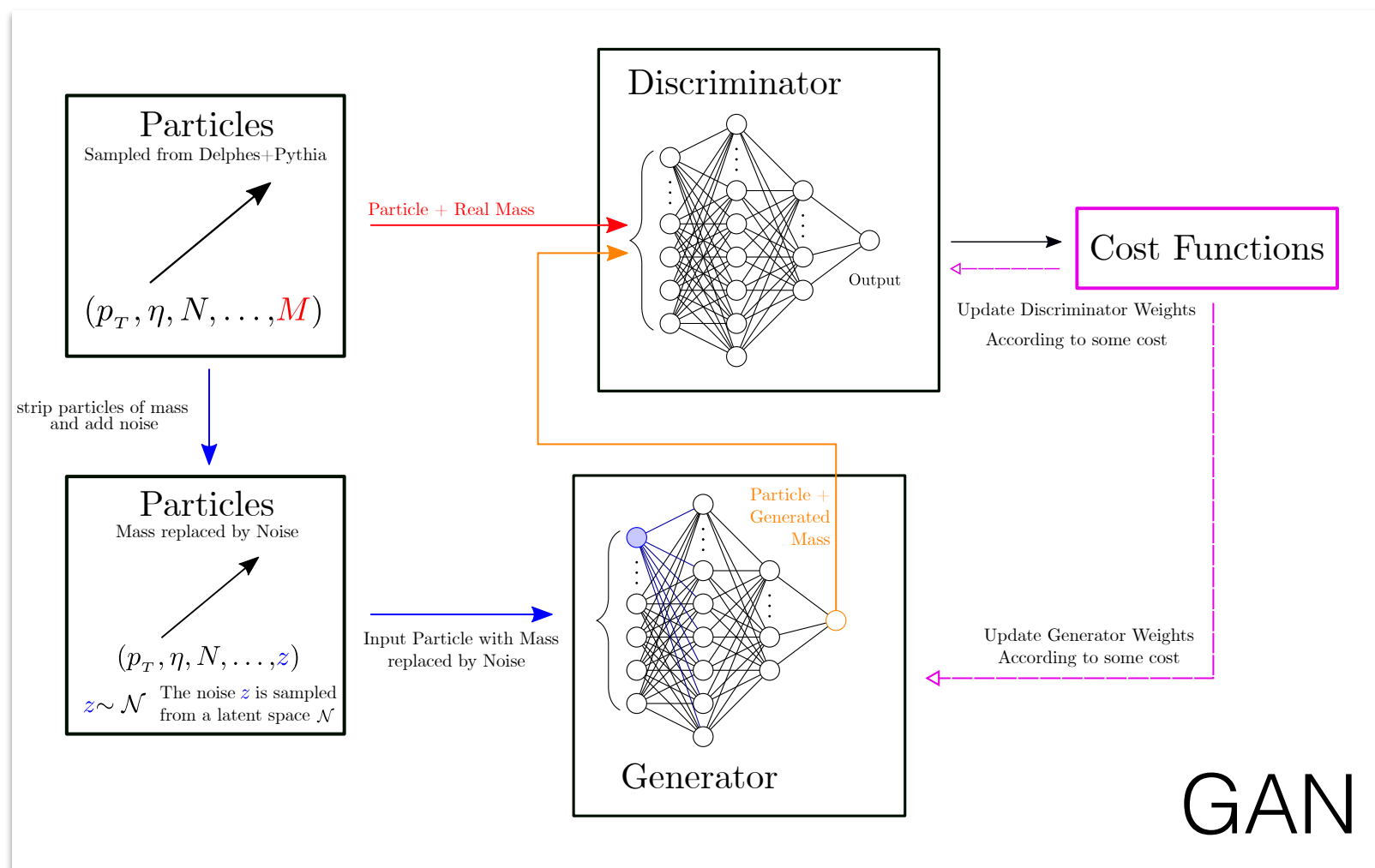
Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

N.B. everything in I've shown before this,  
we trained on simulation, not on data (!)

# Background Estimation

Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

Example 1: unbinned templates for QCD jets to extrapolate in jet multiplicity



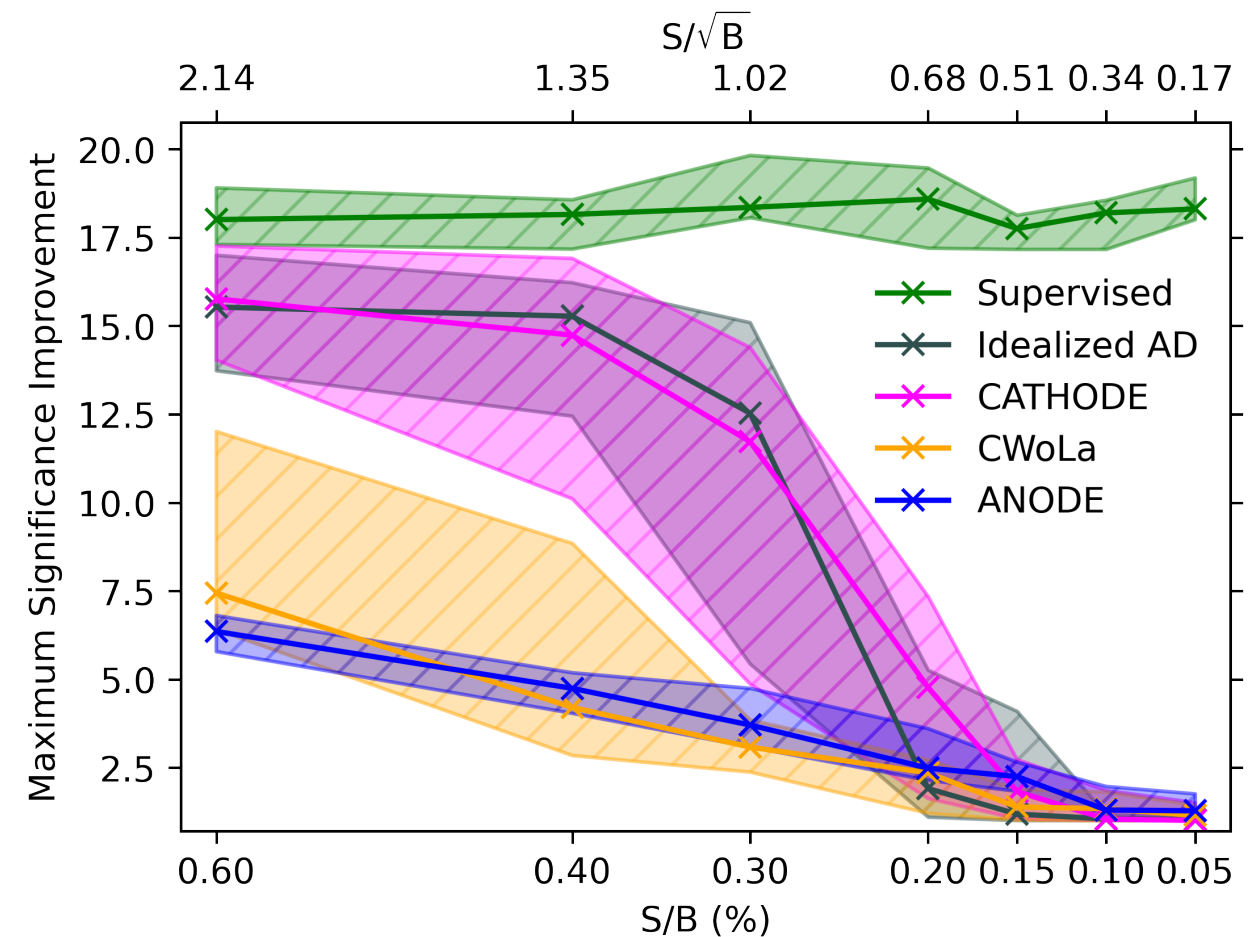
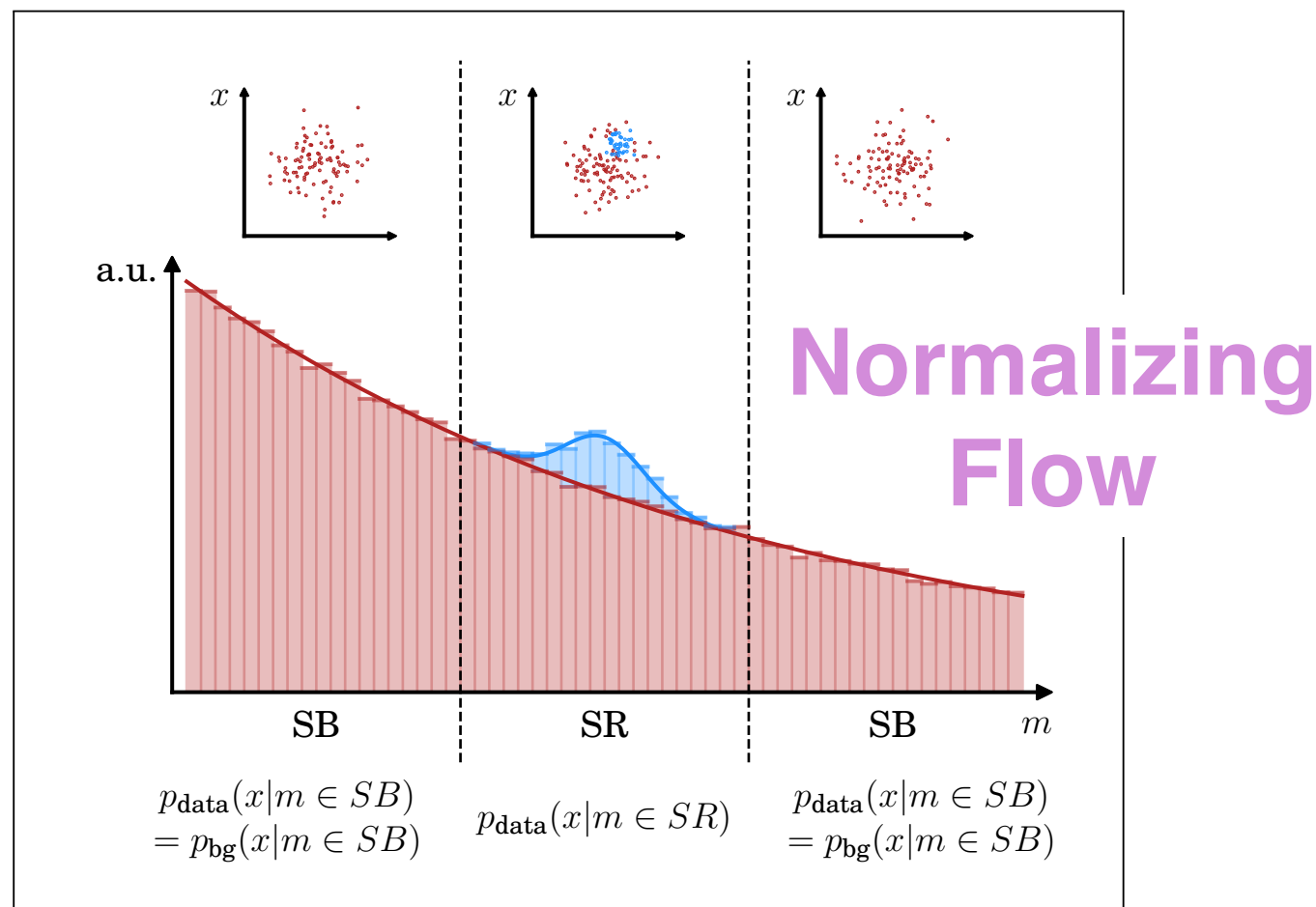




# Background Estimation

Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

Example 2: unbinned templates for QCD jets to extrapolate in dijet mass

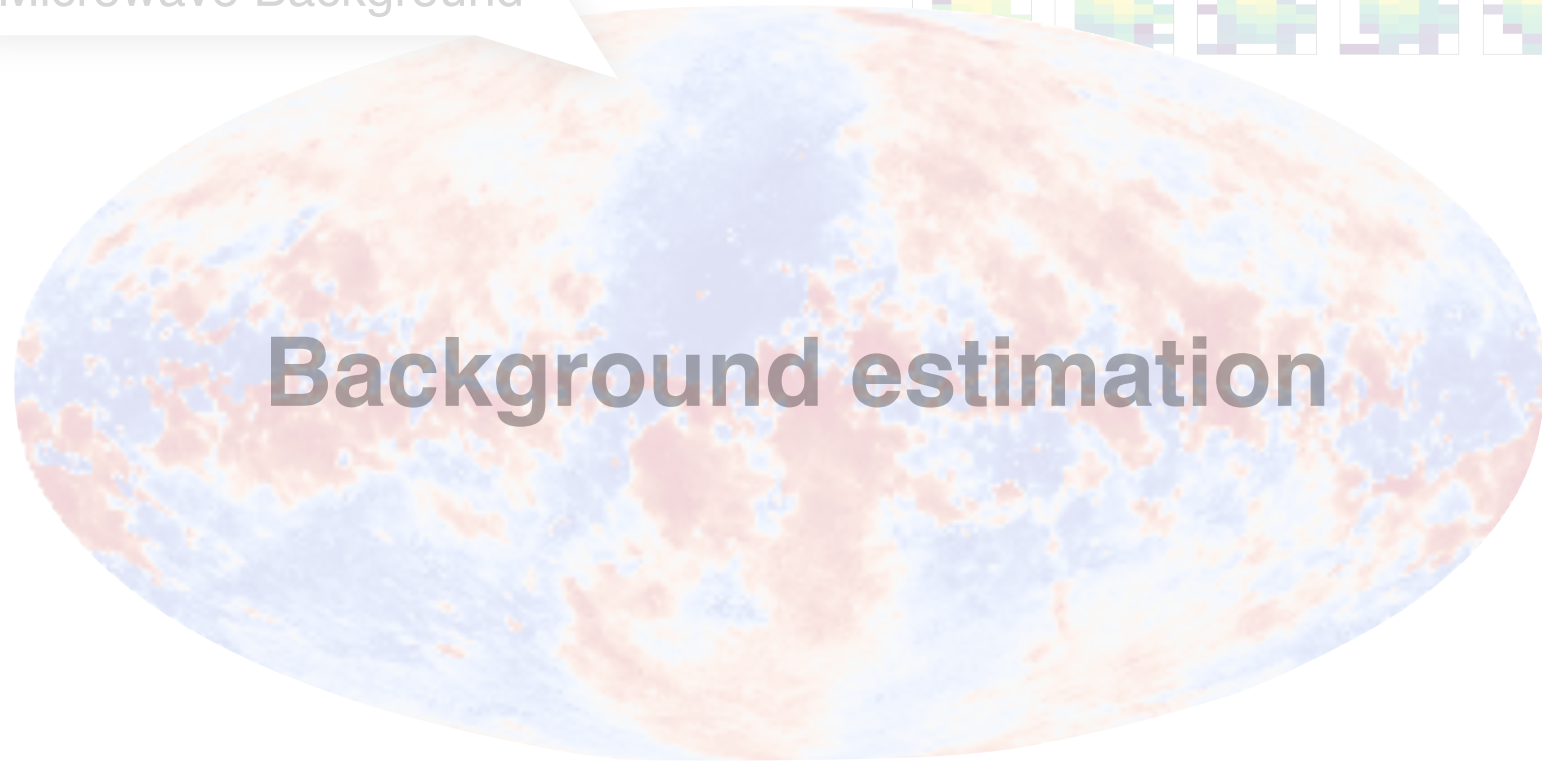


**Anomaly detection**

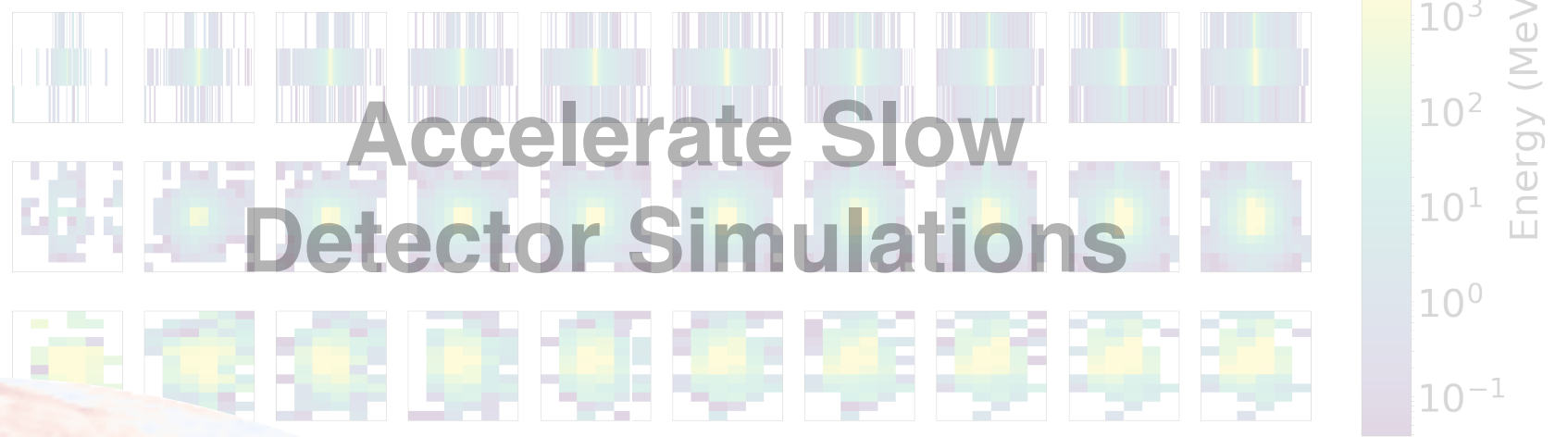
# Generative Models for Particle/Nuclear/Astro

All of these pictures are fake!

Synthetic Galactic radiation for Cosmic Microwave Background



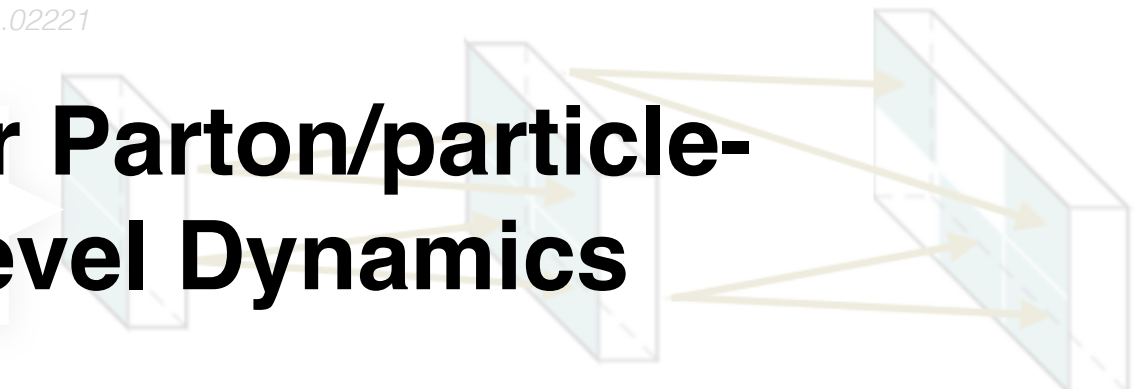
Material Interactions with High Energy Particles



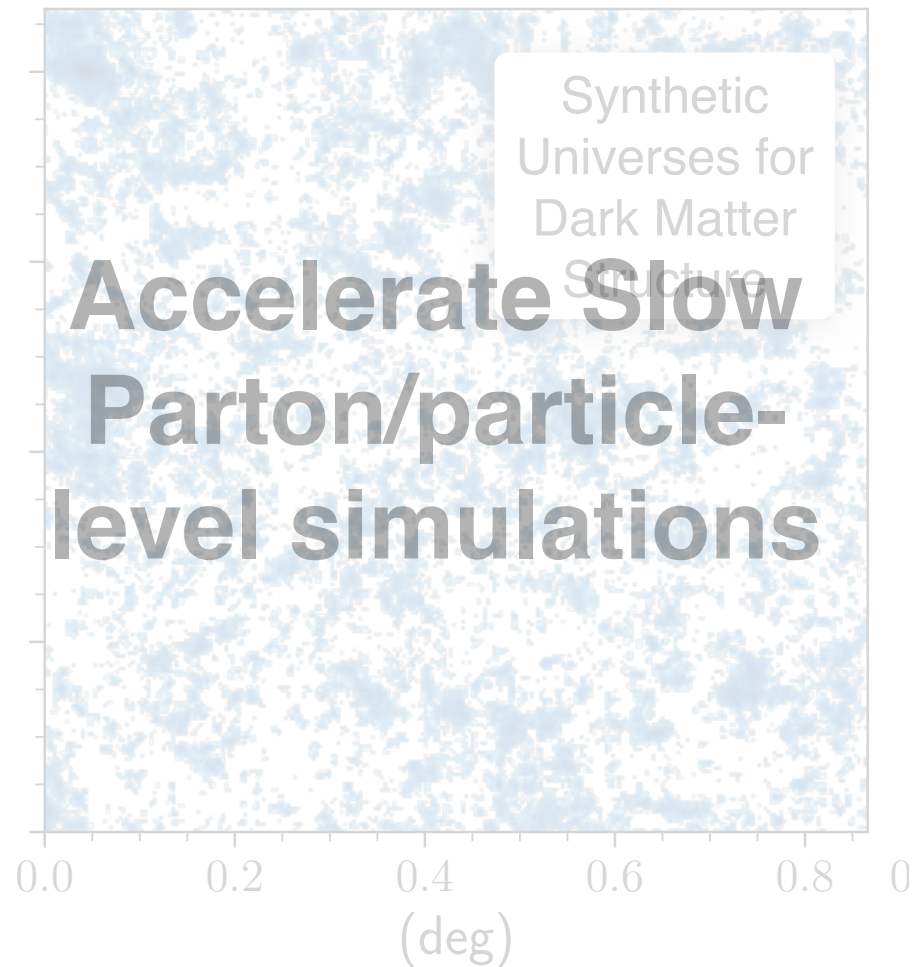
M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

The Structure of Radiation in the Quantum Strong Force

Infer Parton/particle-level Dynamics



Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582



M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)



# Inferring Parton/particle-level Dynamics



69

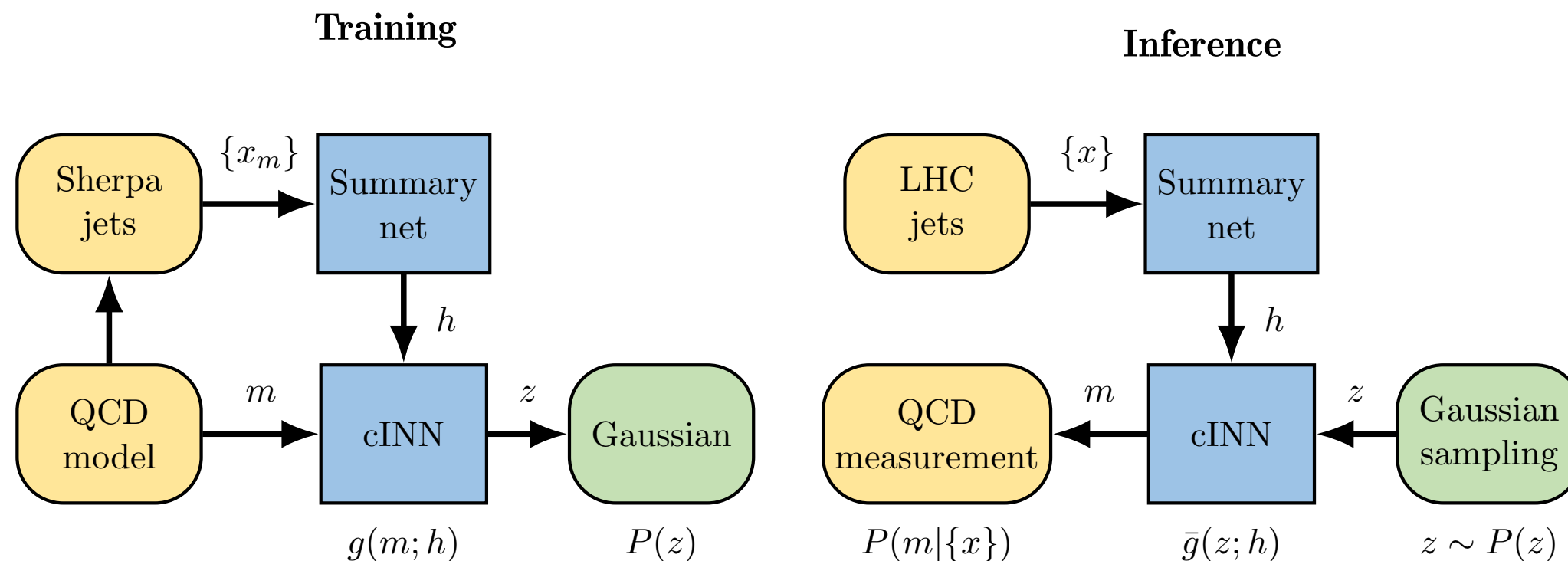
Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)



# Inferring Parton/particle-level Dynamics

Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)

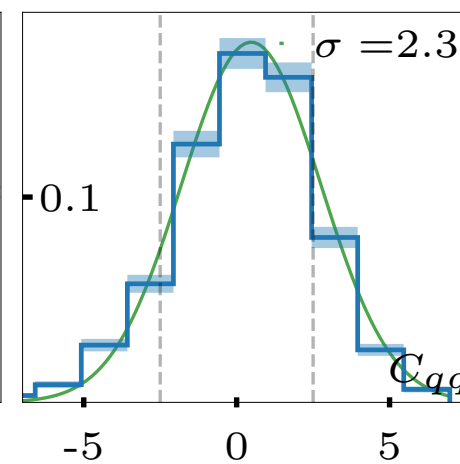
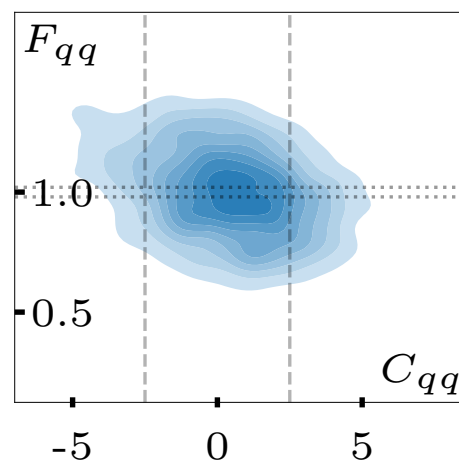
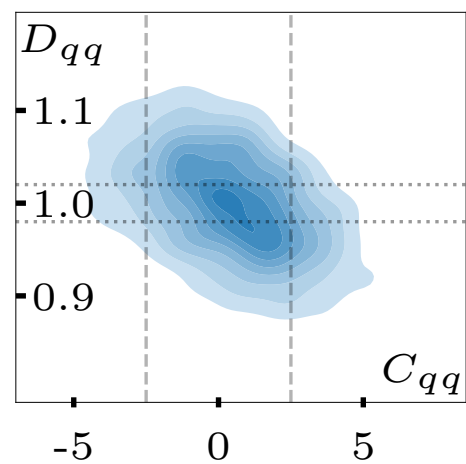
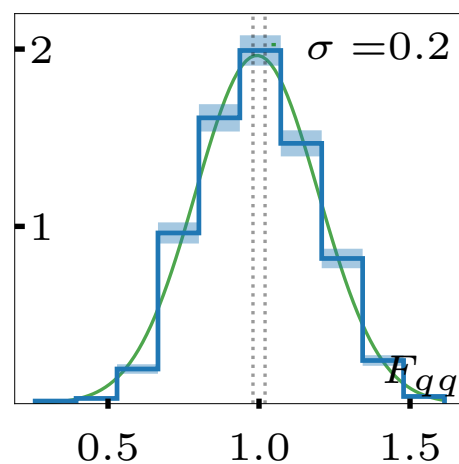
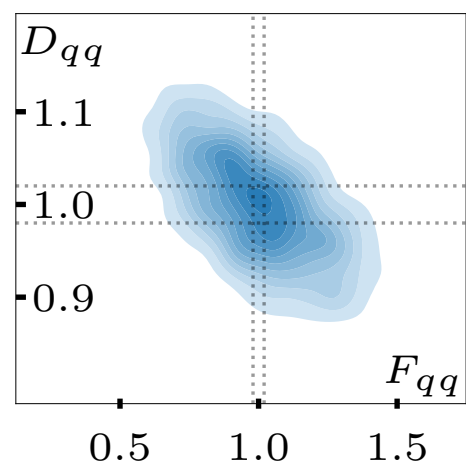
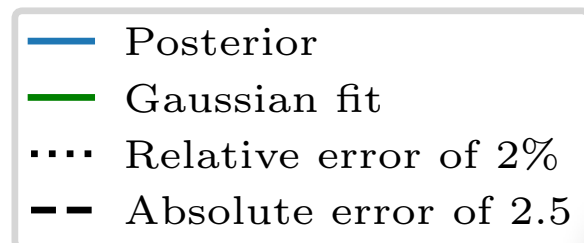
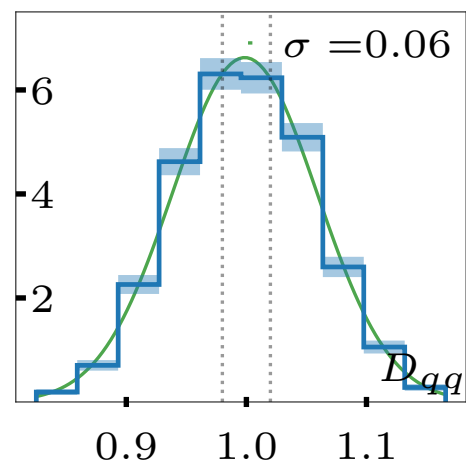
Example 1: Inferring fragmentation functions



See also 1804.09720 (“JUNIPR”) and 2012.06582 (GAN-based)

# Inferring Parton/particle-level Dynamics

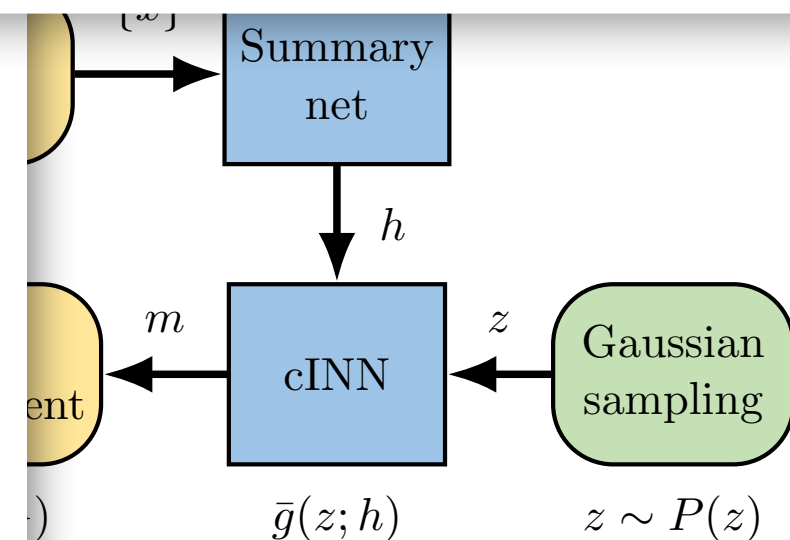
directly for inference?  
(accelerating simulation)



$$P_{qq}(z, y) = C_F \left[ D_{qq} \frac{2z(1-y)}{1-z(1-y)} + F_{qq}(1-z) + C_{qq}yz(1-z) \right]$$

$$P_{gq}(z, y) = T_R \left[ F_{qq} (z^2 + (1-z)^2) + C_{gq}yz(1-z) \right]$$

$$P_{gg}(z, y) = 2C_A \left[ D_{gg} \left( \frac{z(1-y)}{1-z(1-y)} + \frac{(1-z)(1-y)}{1-(1-z)(1-y)} \right) + F_{gg}z(1-z) + C_{gg}yz(1-z) \right]$$



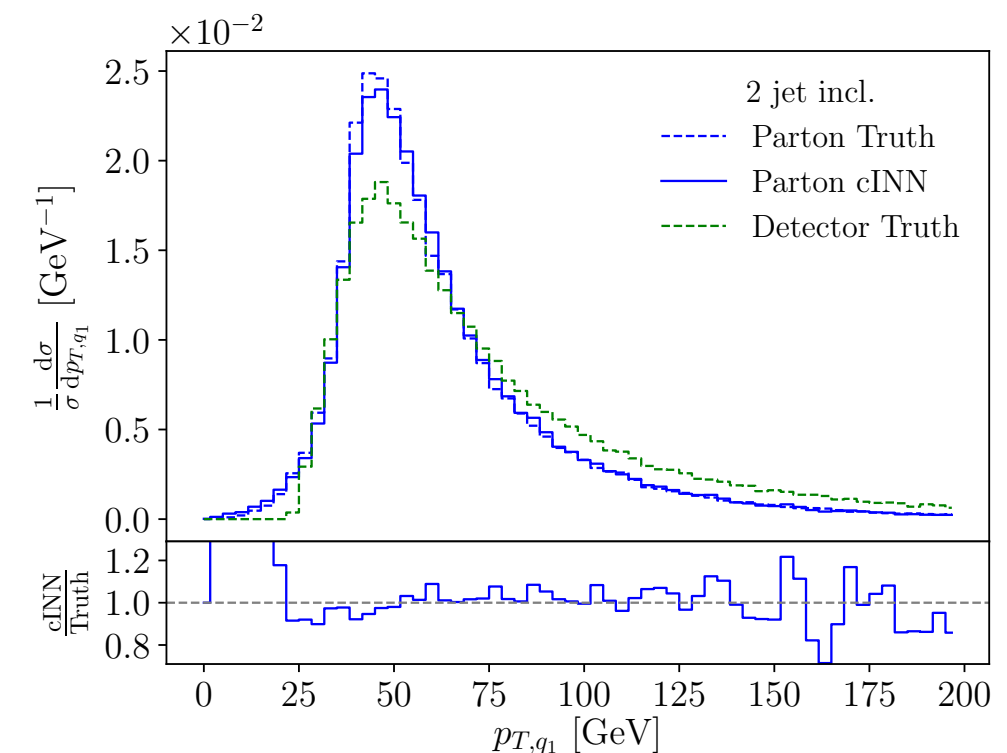
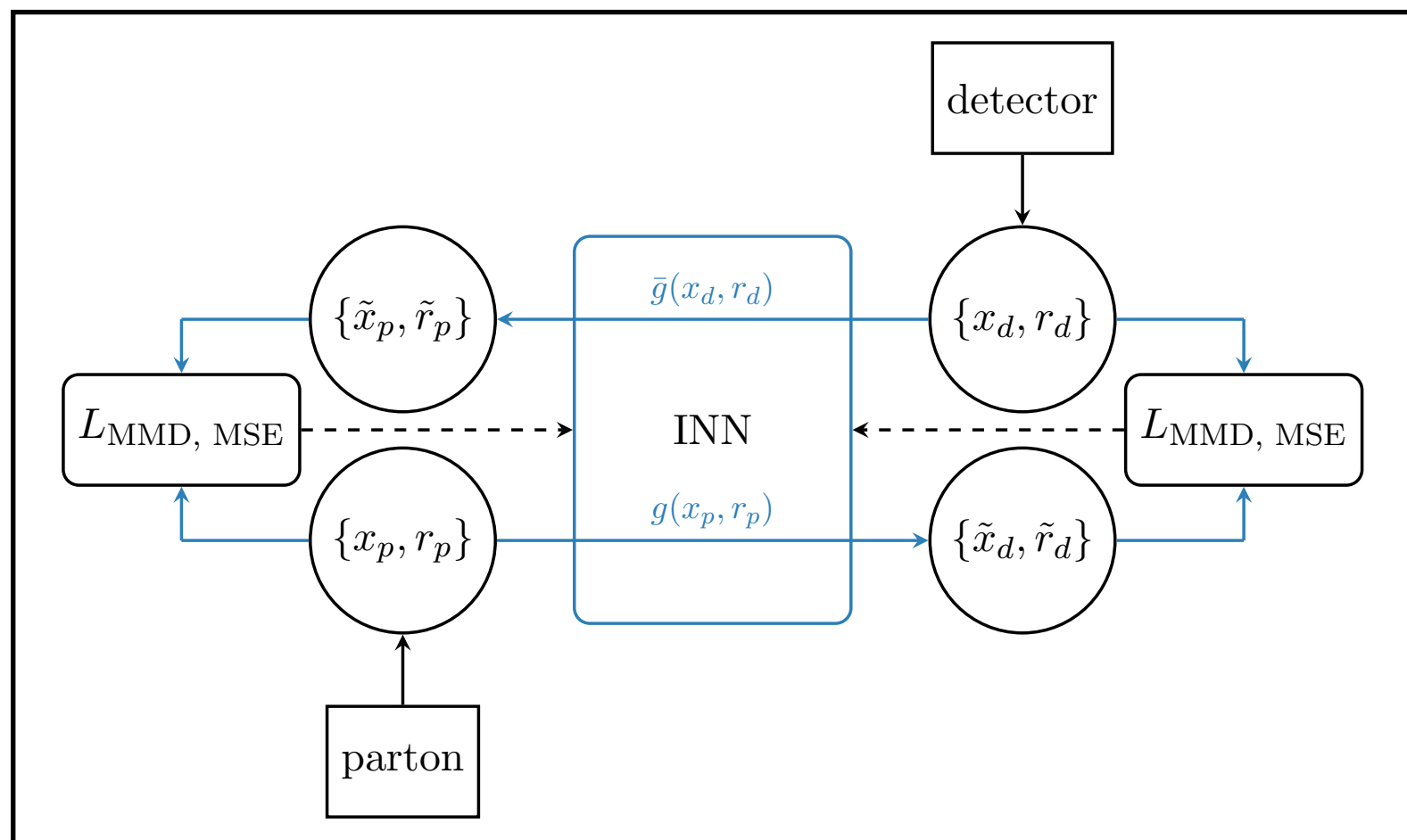
See also 1804.09720 (“JUNIPR”) and 2012.06582 (GAN-based)



# Infering Parton/particle-level Dynamics

Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)

## Example 2: Unfolding

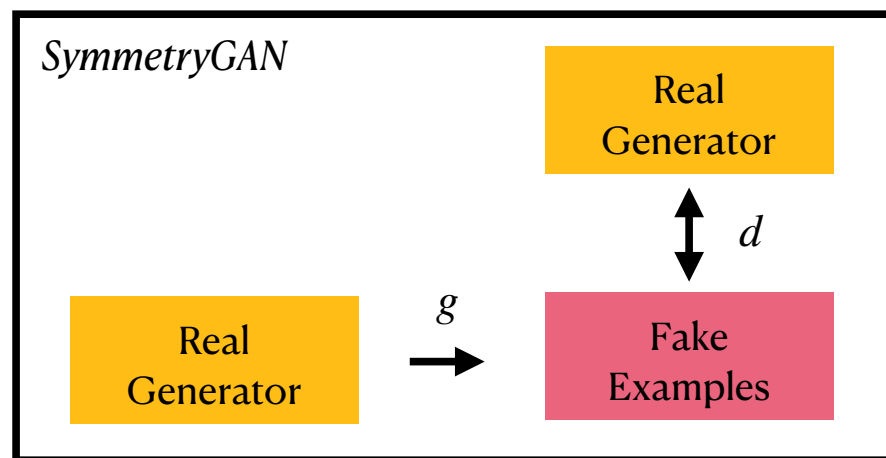
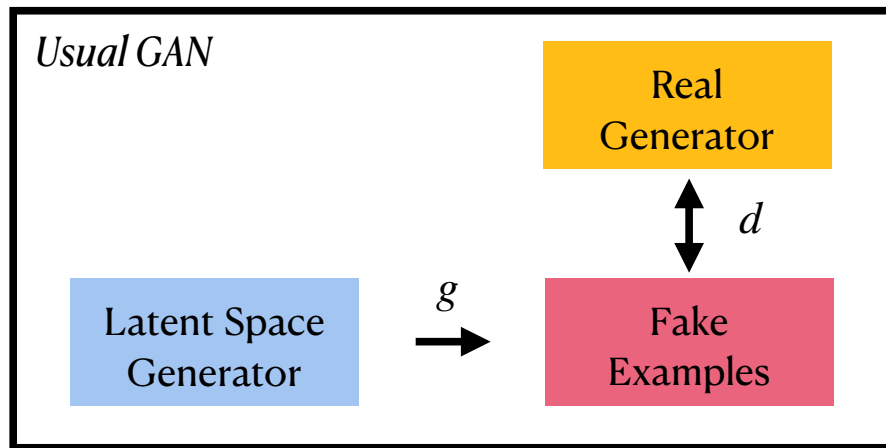


See also 1911.09107 (“OmniFold”) and 2101.08944 (“OTUS”)

Bonus: What else can we do?

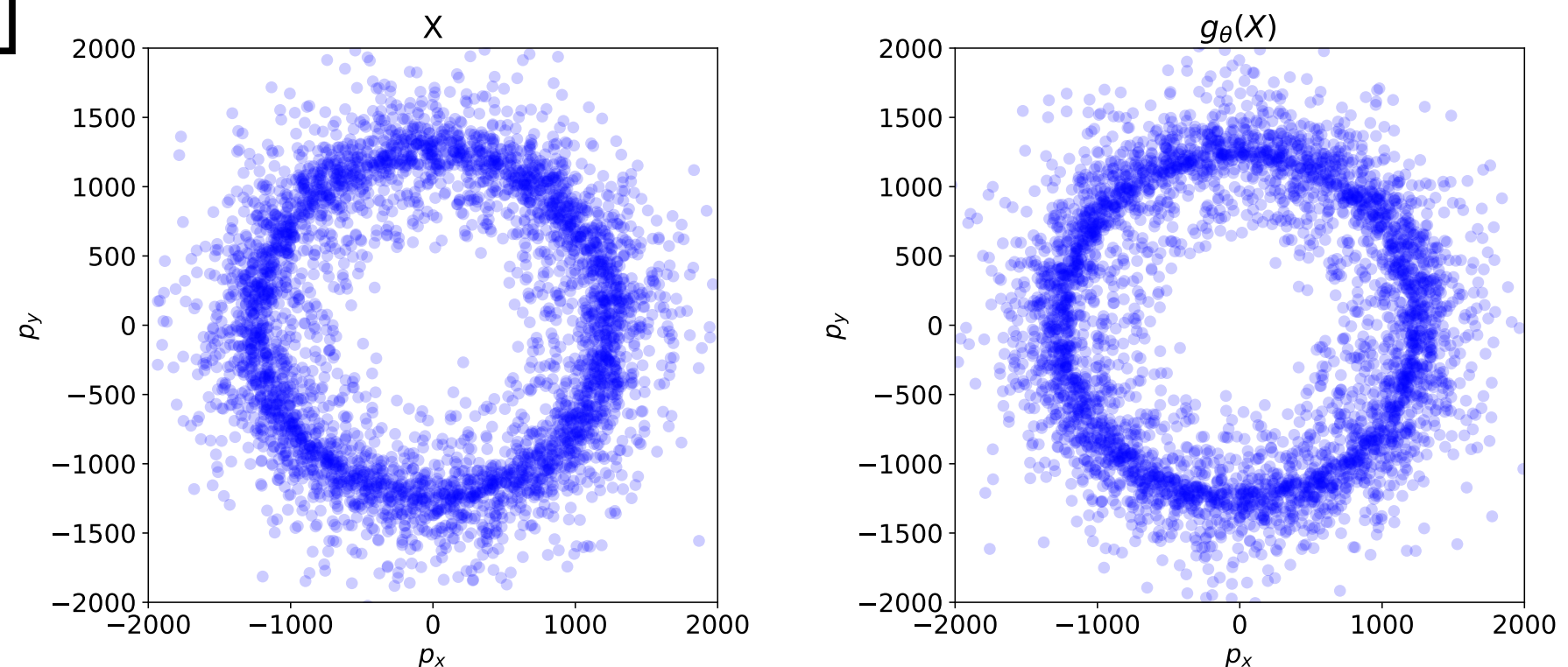


# ex. GANs: Symmetry Discovery



The framework of generative models is quite flexible and we can do more than generate events.

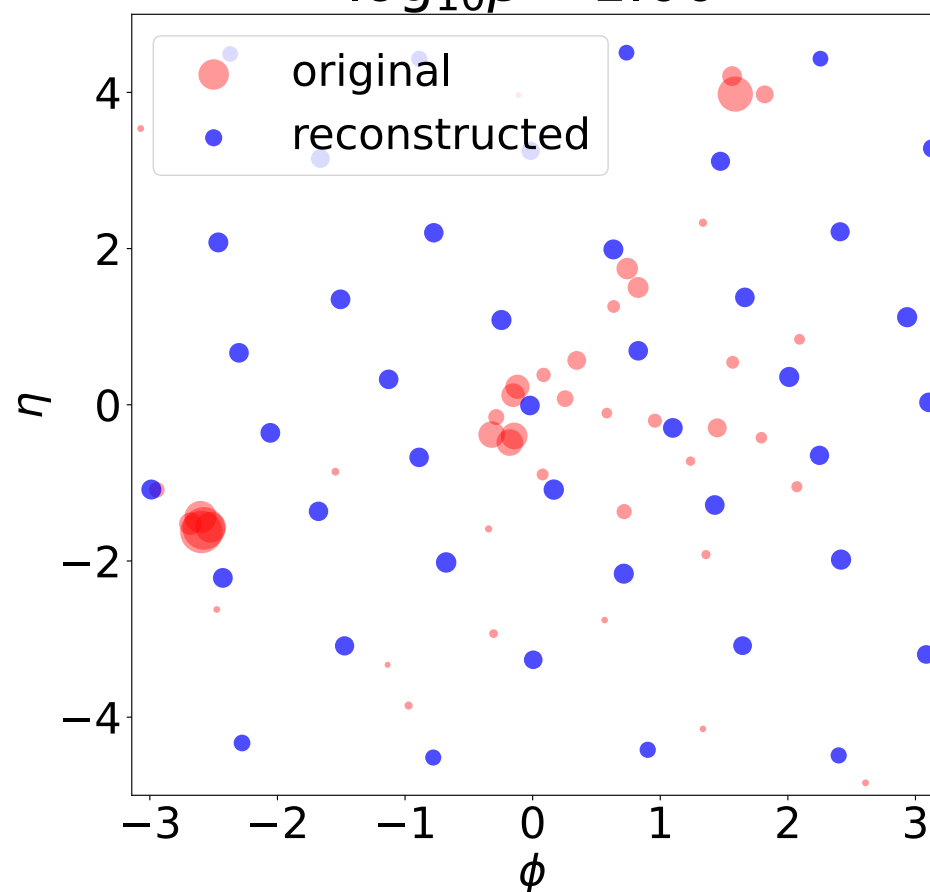
For example, can **discover symmetries** in data!



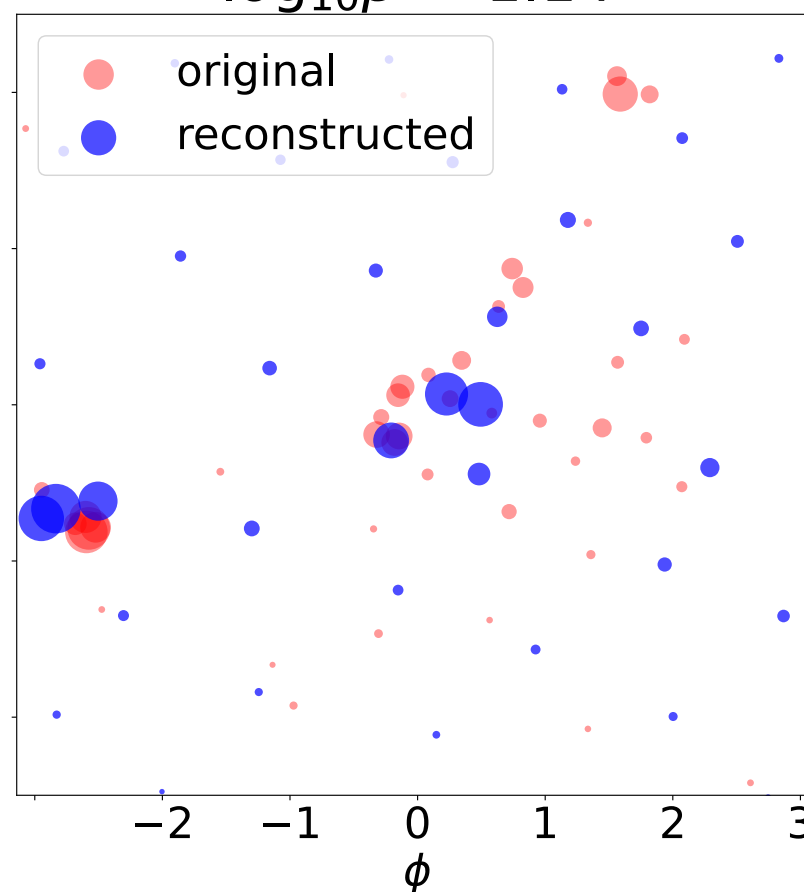


# ex. VAEs: Compression

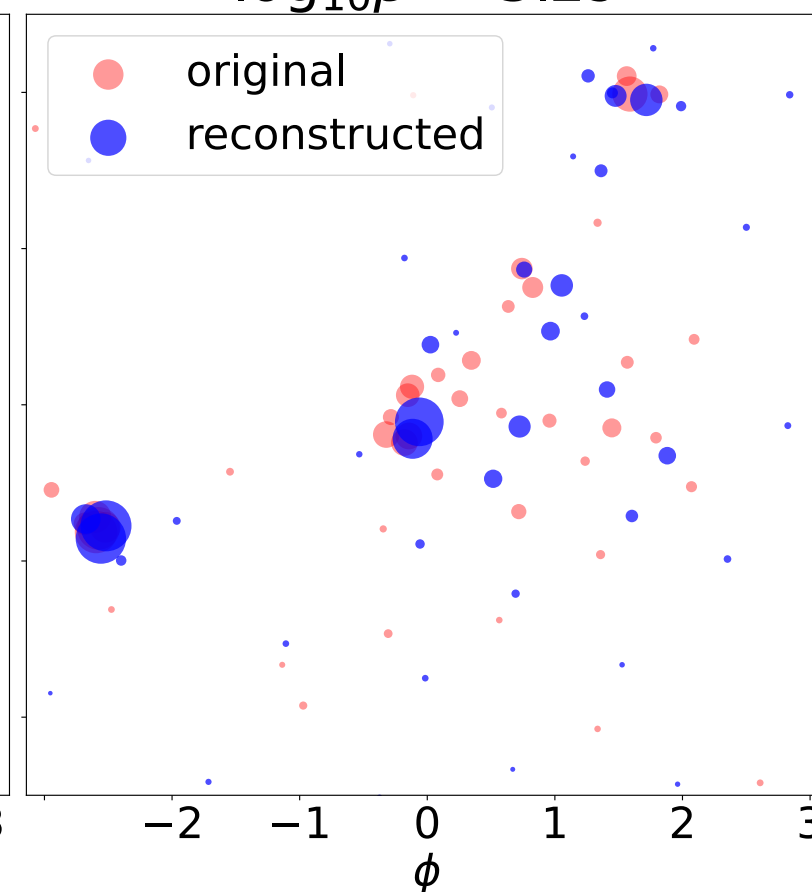
$\log_{10}\beta = 1.00$



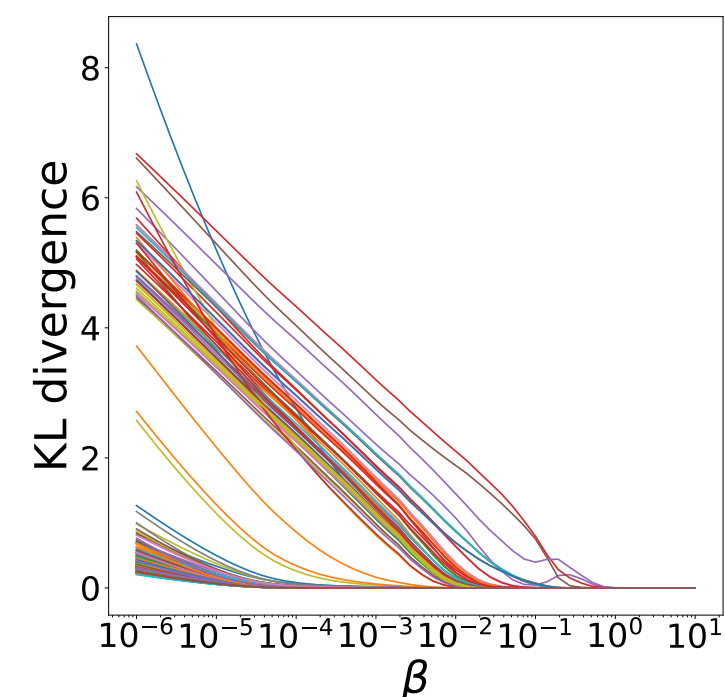
$\log_{10}\beta = -1.14$

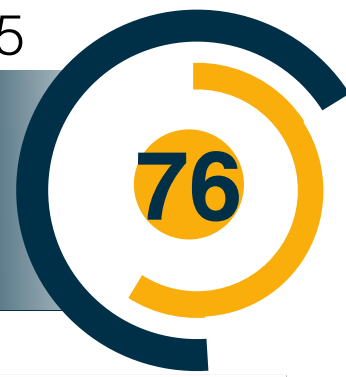


$\log_{10}\beta = -3.29$



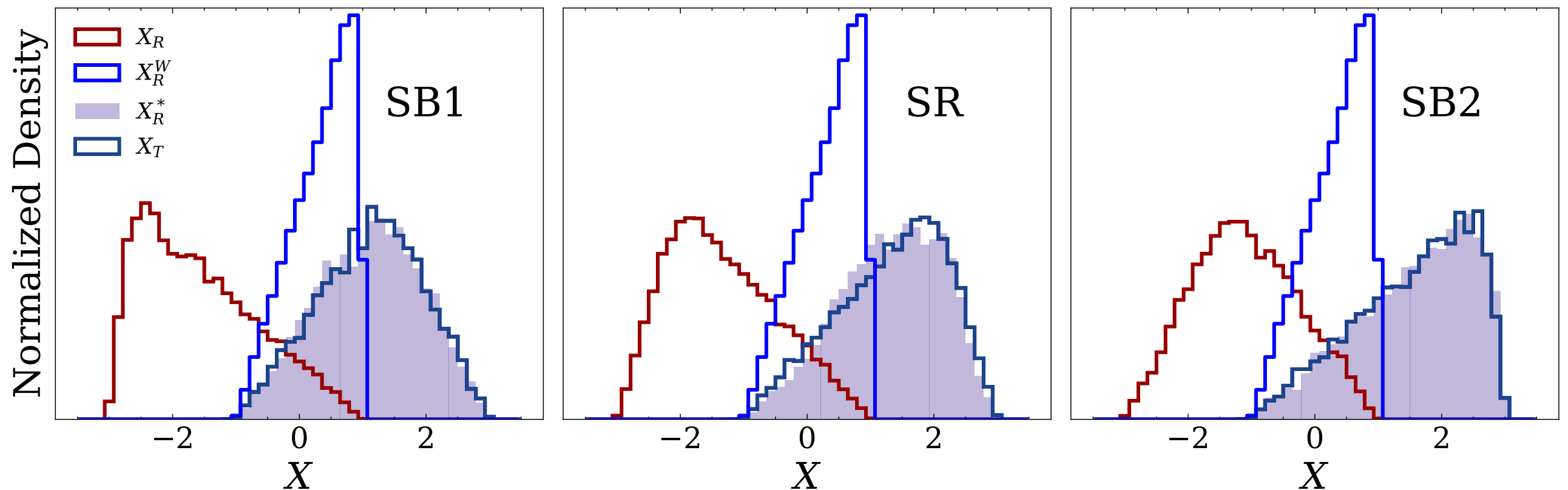
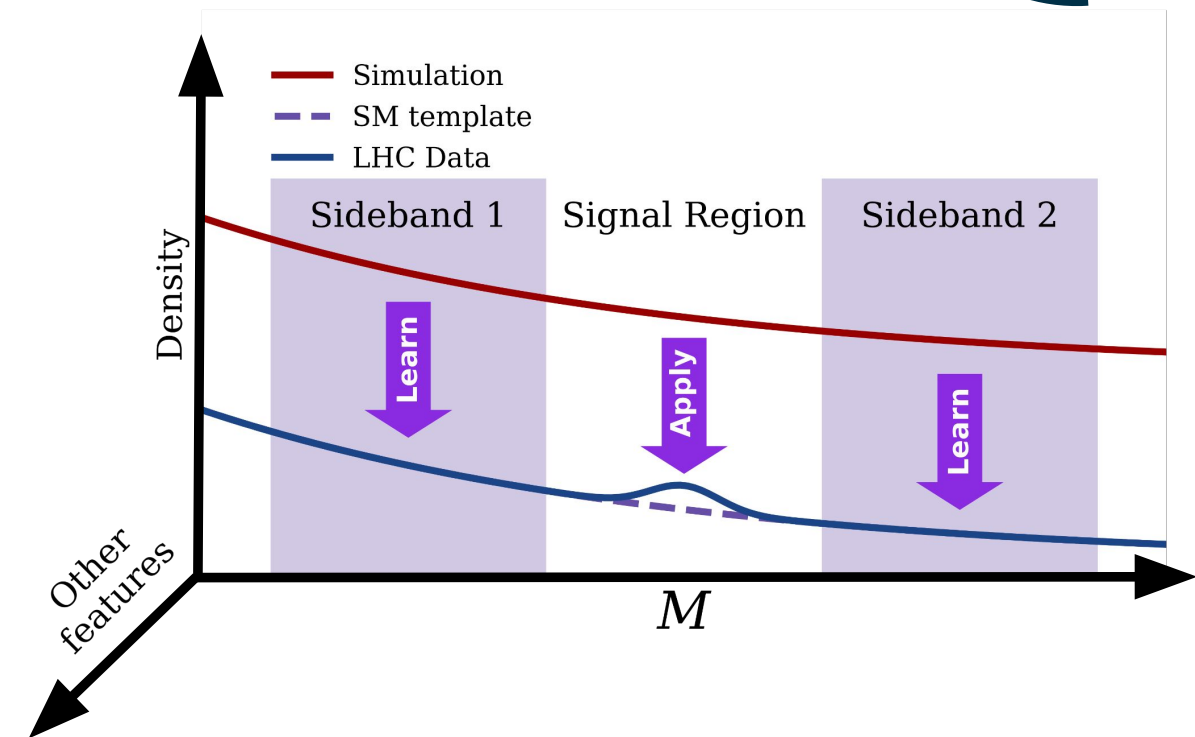
Automatically vary the  
compression level by varying  $\beta$





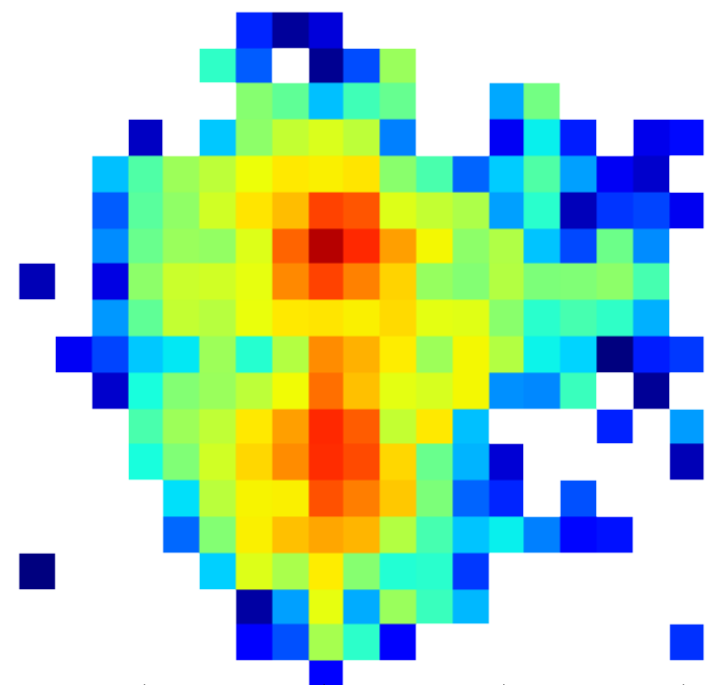
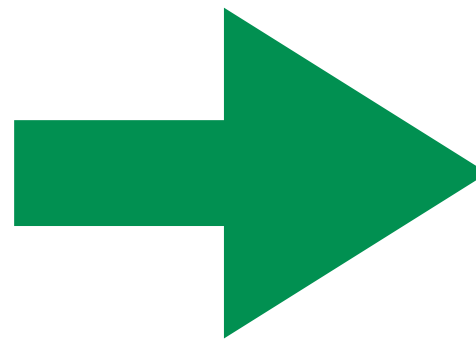
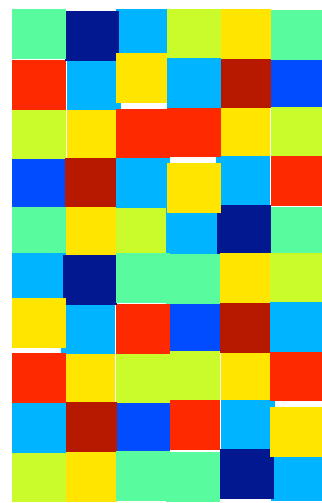
# ex. NFs: Data Morphing

“Move” instead of “Reweight”  
with normalizing flows





1. Why Generative Models?
2. Warm up: mixture density networks
3. Next: brief overview of GANs, VAEs, NFs
4. Statistical amplification
5. Applications
6. Bonus



## Generative models hold great promise for many areas of physics research

*This is an area of rapid development at the intersection of theory & experiment and I'm excited for where it will take us!*

*Examples today were not exhaustive ...  
Prof. Duarte has also been a pioneer in this area!*

This is a link to a recent Berkeley workshop dedicated to generative models →

**Deep Generative Models for Fundamental Physics**

March 17, 2021

**Organizing Committee**

- Ellianna Abrahms, Department of Astronomy, UC Berkeley
- Vanessa Boehm, Department of Physics, UC Berkeley
- Aishik Ghosh, UC Irvine / Physics Division, Berkeley Lab
- Yue Shi Lai, Nuclear Science Division, Berkeley Lab
- Mustafa Mustafa, NERSC, Berkeley Lab
- Ben Nachman, Physics Division, Berkeley Lab
- Giuseppe Puglisi, Space Science Laboratory, UC Berkeley

Questions?

