

A cosmic background image featuring a deep space scene with various galaxies and stars. Two prominent galaxies are visible in the upper half, one on the left and one on the right, both showing spiral structures and bright central cores. The background is filled with numerous smaller stars and distant galaxies, creating a sense of vastness and depth.

PHYS 141/241

Lecture 02: Newtonian/Lagrangian/Hamiltonian Mechanics

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Classical equations of motion

- Several formulations
 - Newtonian
 - Lagrangian
 - Hamiltonian
- Advantages of non-Newtonian formulations
 - More general, no need for “fictitious” forces
 - Better suited for multiparticle systems
 - Better handling of constraints
 - Can be derived from more basic postulates
- Assume conservative forces:
 $\mathbf{F} = -\nabla U$ (gradient of a scalar potential energy) .

Newtonian formulation

- Cartesian spatial coordinates $\mathbf{r}_i = (x_i, y_i, z_i)$ are primary variables

- N 2nd-order ordinary differential equations: $m_i \frac{d^2 \mathbf{r}_i}{dt^2} \equiv m_i \ddot{\mathbf{r}}_i = \mathbf{F}_i$

- Example: 2D motion in a central force field

$$m\ddot{x} = \mathbf{F} \cdot \hat{x} = -f(r)\hat{\mathbf{r}} \cdot \hat{x} = -xf(r)$$

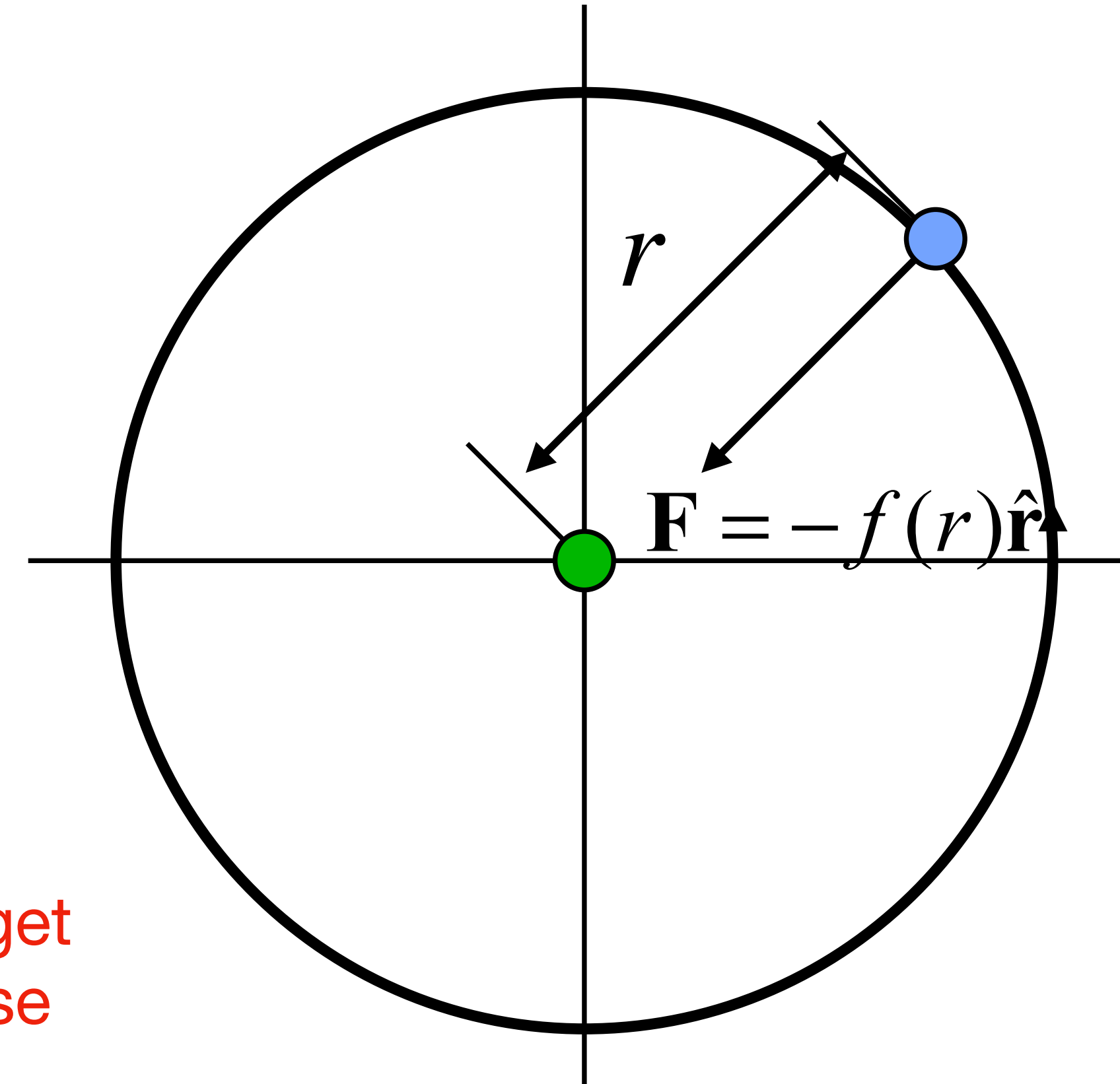
$$m\ddot{y} = \mathbf{F} \cdot \hat{y} = -f(r)\hat{\mathbf{r}} \cdot \hat{y} = -yf(r)$$

- Transform to polar coordinates

$$mr^2\dot{\theta} = \ell \quad \text{Constant angular momentum}$$

$$m\ddot{r} = -f(r) + \boxed{\frac{\ell^2}{mr^3}} \quad \text{"Fictitious" force}$$

Tedious to get
general case



Lagrangian formulation

- Independent of coordinate system

- Define the Lagrangian $L(q, \dot{q}) \equiv K(q, \dot{q}) - U(q)$ and action $S = \int_{t_1}^{t_2} L(q, \dot{q}) dt$

- Equations of motion arise from the principle of least action (PLA) $\delta S = 0$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_j} \right) - \frac{\partial L}{\partial q_j} = 0, \quad j = 1, \dots, N \quad (N \text{ 2nd-order differential equations})$$

- Central-force example: $L = K - U = \frac{1}{2}(m\dot{r}^2 + r^2\dot{\theta}^2) - U(r)$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}} \right) = \frac{\partial L}{\partial r} \Rightarrow m\ddot{r} = mr\dot{\theta}^2 - f(r) \quad \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\theta}} \right) = \frac{\partial L}{\partial \theta} \Rightarrow \frac{d}{dt}(mr^2\dot{\theta}) = 0$$

$$\mathbf{F}_r = -\nabla_r U = -f(r)\hat{\mathbf{r}} \quad 4$$

Hamiltonian formulation

- Appropriate for application to statistical and quantum mechanics
- Newtonian and Lagrangian viewpoints take the q_i as the fundamental variables
 - N -variable configuration space
 - $\dot{q}_i$ appears only as convenient shorthand for dq/dt
 - Formulas are 2nd-order differential equations
- Hamiltonian formulation seeks to work with 1st-order differential equations
 - $2N$ variables: phase space density of the N -body system
 - Treat the coordinate and its time derivative as independent variables

Hamiltonian formulation

- Mathematically, Lagrangian treats q and $\dot{q}$ as distinct
 - Identify generalized momentum as $p_j = \frac{\partial L}{\partial \dot{q}_j}$
- Example:
 - Lagrangian $L = K - U = \frac{1}{2}m\dot{q}^2 - U(q)$ and momentum $p = \frac{\partial L}{\partial \dot{q}} = m\dot{q}$
 - Lagrangian equations of motion: $\frac{\partial p_j}{\partial t} = \frac{\partial L}{\partial q_j}$
- We would like a formulation in which p is an independent variable
 - p_i is the derivative of the Lagrangian with respect to $\dot{q}_i$ and we're looking to replace $\dot{q}_i$ with p_i

Hamiltonian formulation

- Using a *Legendre transform*, we can define the Hamiltonian

$$\begin{aligned} H(q_j, p_j) &= - \left(L(q_j, \dot{q}_j) - \sum_i p_i \dot{q}_i \right) \\ &= -K(q_j, \dot{q}_j) + U(q_j) + \sum_i \frac{\partial K}{\partial \dot{q}_i} \dot{q}_i \\ &= -\frac{1}{2} \sum_i m_i \dot{q}_i^2 + U(q_j) + \frac{1}{2} \sum_i (2m_i \dot{q}_i) \dot{q}_i \\ &= +\frac{1}{2} \sum_i m_i \dot{q}_i^2 + U(q_j) \\ &= K + U \end{aligned}$$

Hamiltonian formulation

- Hamiltonian's equations of motion
 - Rewrite Lagrangian equations in terms of momentum

$$\frac{dp}{dt} = \dot{p} = \frac{\partial L}{\partial q} \quad \text{Lagrange's equation of motion}$$

$$p = \frac{\partial L}{\partial \dot{q}} \quad \text{Definition of momentum}$$

Differential change in L :

$$\begin{aligned} dL &= \frac{\partial L}{\partial q} dq + \frac{\partial L}{\partial \dot{q}} d\dot{q} \\ &= \dot{p} dq + p d\dot{q} \end{aligned}$$

$$\left\{ \begin{aligned} \dot{q} &= + \frac{\partial H}{\partial p} \\ \dot{p} &= - \frac{\partial H}{\partial q} \end{aligned} \right.$$

Hamilton's equation of motion

Legendre transform:

$$\begin{aligned} H &= - (L - p\dot{q}) \\ dH &= - \dot{p} dq + \dot{q} dp \end{aligned}$$

Conservation of energy: $\frac{\partial H}{dt} = - \dot{p} \frac{dq}{dt} + \dot{q} \frac{dp}{dt} = 0$

Hamiltonian formulation

Lagrange's equations

- Particle motion in a central force field

$$H = K + U$$
$$= \frac{p_r^2}{2m} + \frac{p_\theta^2}{2mr^2} + U(r)$$

Hamilton's equations

$$\dot{q} = + \frac{\partial H}{\partial q}$$

$$p = + \frac{\partial H}{\partial \dot{q}}$$

$$(1) \frac{dr}{dt} = \frac{p_r}{m}$$

$$(2) \frac{d\theta}{dt} = \frac{p_\theta}{mr^2}$$

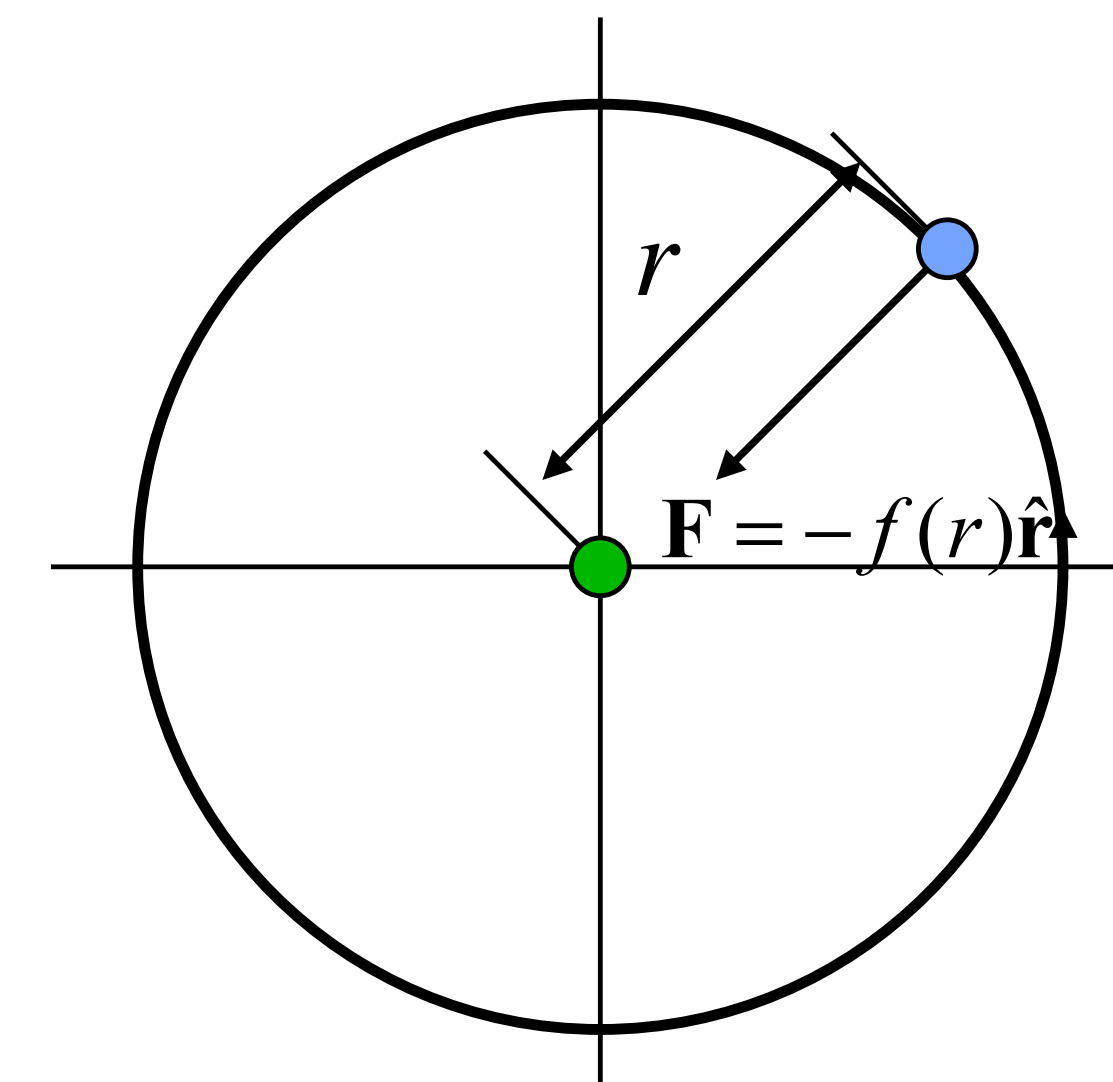
$$(3) \frac{dp_r}{dt} = \frac{p_\theta}{mr^3} - f(r)$$

$$(4) \frac{dp_\theta}{dt} = 0$$

$$(1) m\ddot{r} = mr\dot{\theta}^2 - f(r)$$

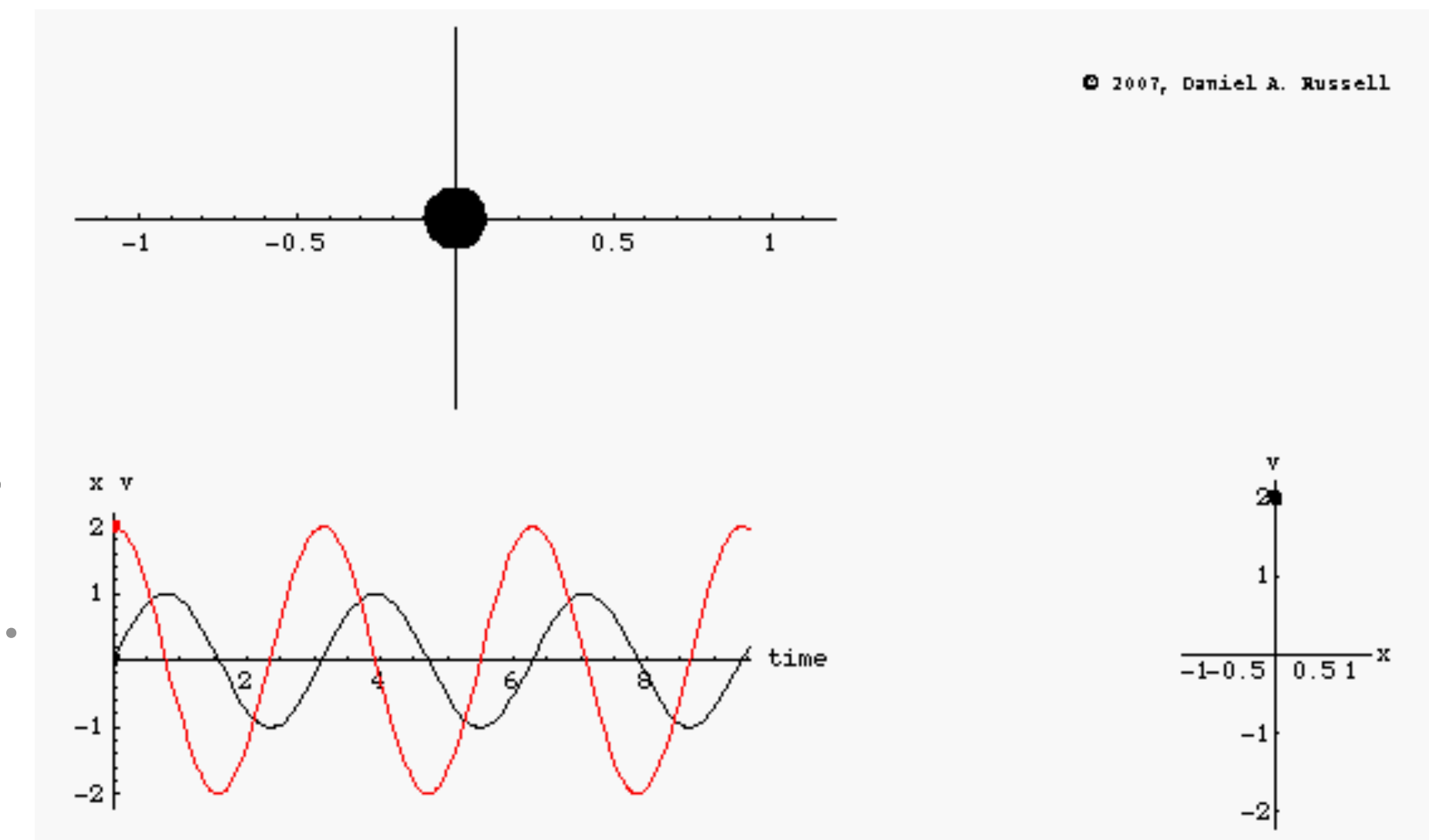
$$(2) \frac{d}{dt}(mr^2\dot{\theta}) = 0$$

- Equations equivalent, but theoretical basis is better



Phase space

- Complete picture of phase space
 - Full specification of micro state of the system is given by the values of all positions and all momentum of all bodies
$$G = (p_j, r_j)$$
 - View all positions and momenta as independent coordinates
 - Connection between them comes through equations of motion
- Motion through phase space
 - Helpful to think of dynamics as “simple” movement through phase space
 - Facilitates connection to quantum mechanics
 - Basis for theoretical treatment of dynamics



Integration algorithms

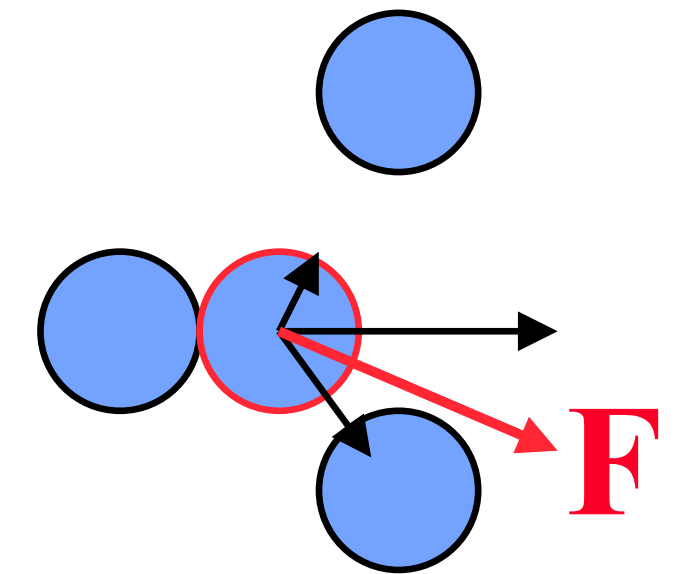
- Equations of motion in cartesian coordinates

$$\frac{d\mathbf{r}_j}{dt} = \frac{\mathbf{p}_j}{m}$$

$$\frac{d\mathbf{p}_j}{dt} = \mathbf{F}_j = \sum_{i=1, i \neq j}^N \mathbf{F}_{ij} \quad (\text{pairwise additive forces})$$

- Desirable features of an integrator
 - Minimal need to compute forces (expensive)
 - Good ability for large time steps
 - Good accuracy
 - Conserves energy and momentum
 - Time-reversible
 - Phase space area-preserving (symplectic)

More on these later



Time integration methods

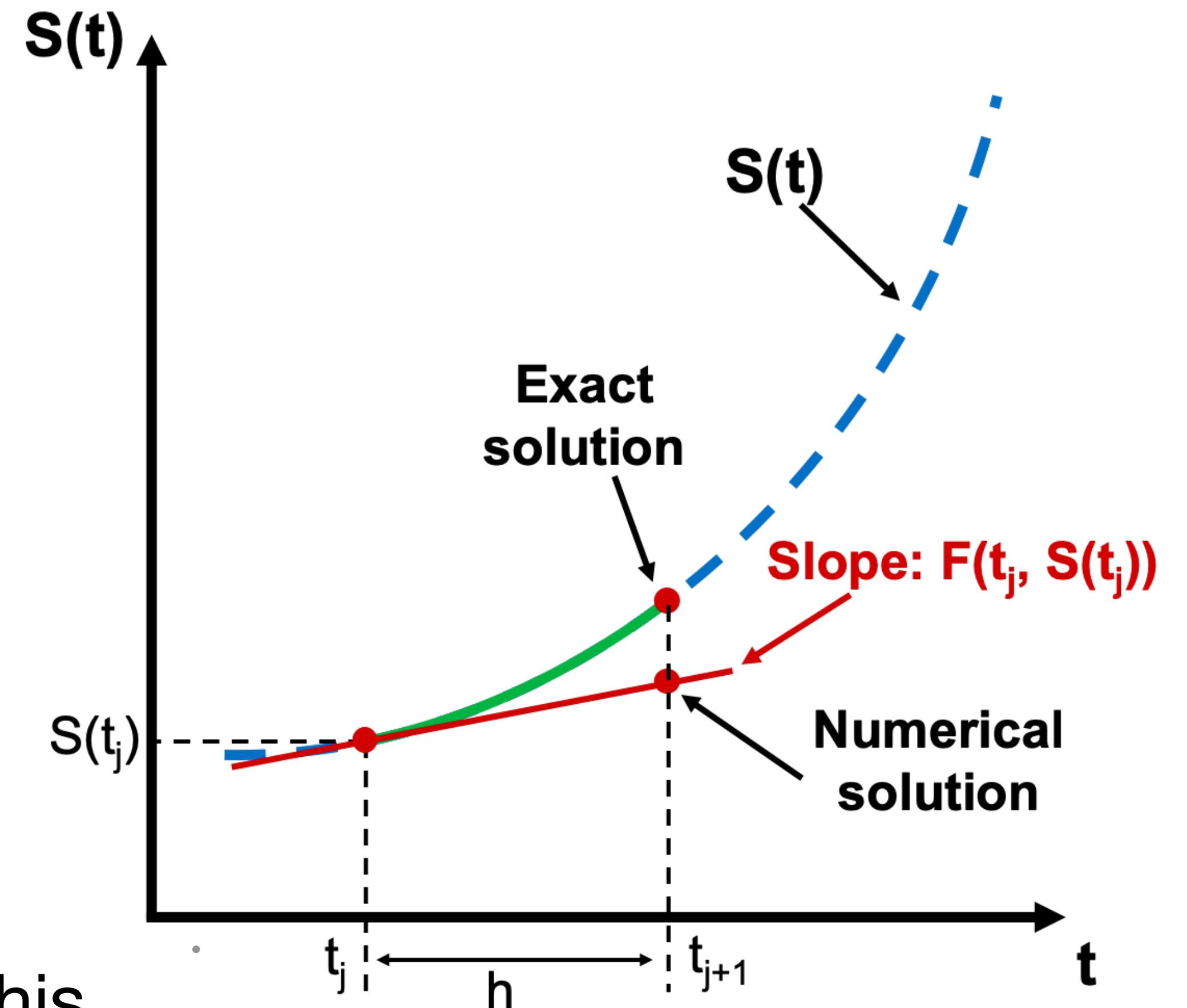
- Let's integrate a first-order ordinary differential equation (ODE):

$$\frac{dS}{dt} = \dot{S}(t) = F(t, S(t))$$

- Example: exponential function

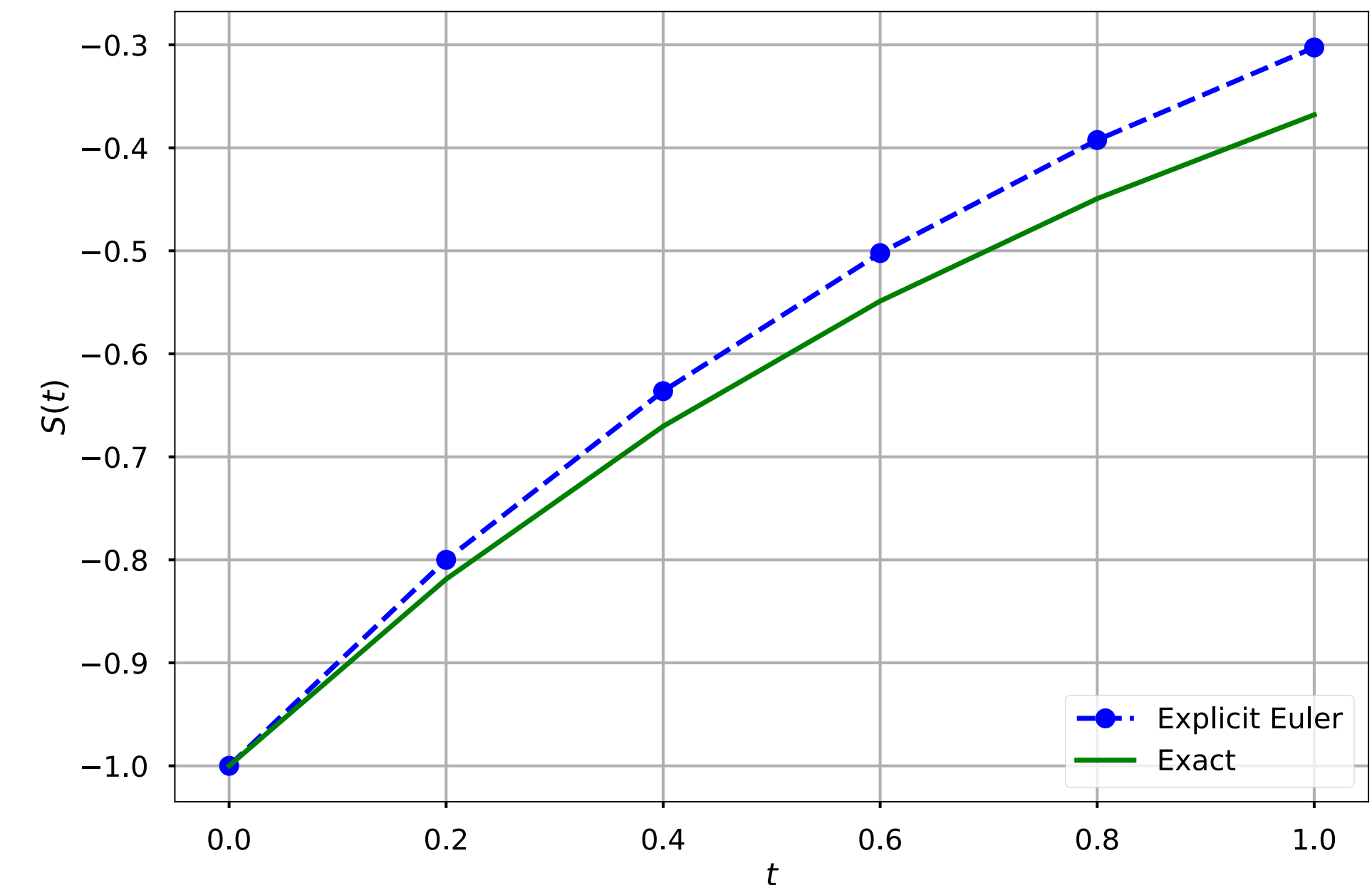
$$\dot{y} = e^{-t}$$

- A numerical approximation to the ODE is a set of values: $\{S_0, S_1, S_2, \dots\}$ and $\{t_0, t_1, t_2, \dots\}$
- There are many different ways of obtaining this



Euler method

- Explicit Euler method: $S_{n+1} = S_n + F(t_n, y_n)\Delta t$
 - Simplest of all
 - Right-hand side depends on things already known: **explicit** method
 - The error in a single step is $\mathcal{O}(\Delta t^2)$, but for N steps needed for a finite time interval, the total error scales as $\mathcal{O}(\Delta t)$!
 - Only first-order accurate, not advised to use!
- Implicit Euler method: $S_{n+1} = S_n + F(t_{n+1}, y_{n+1})\Delta t$
 - Excellent stability properties
 - Suitable for stiff ODE
 - Requires implicit solver for y_{n+1} (i.e. more computations)



Predictor-corrector methods

- Predictor-corrector methods of solving initial value problems improve the approximation accuracy by querying the function several times at different locations (predictions), and then using a weighted average of the results (corrections) to update the state
- Two formulas: the predictor and corrector
 - The **predictor** is an explicit formula and first estimates the solution at t , i.e. we can use Euler method or some other methods to finish this step.
 - Using the found $S(t_{n+1})$, the **corrector** can calculate a new, more accurate solution

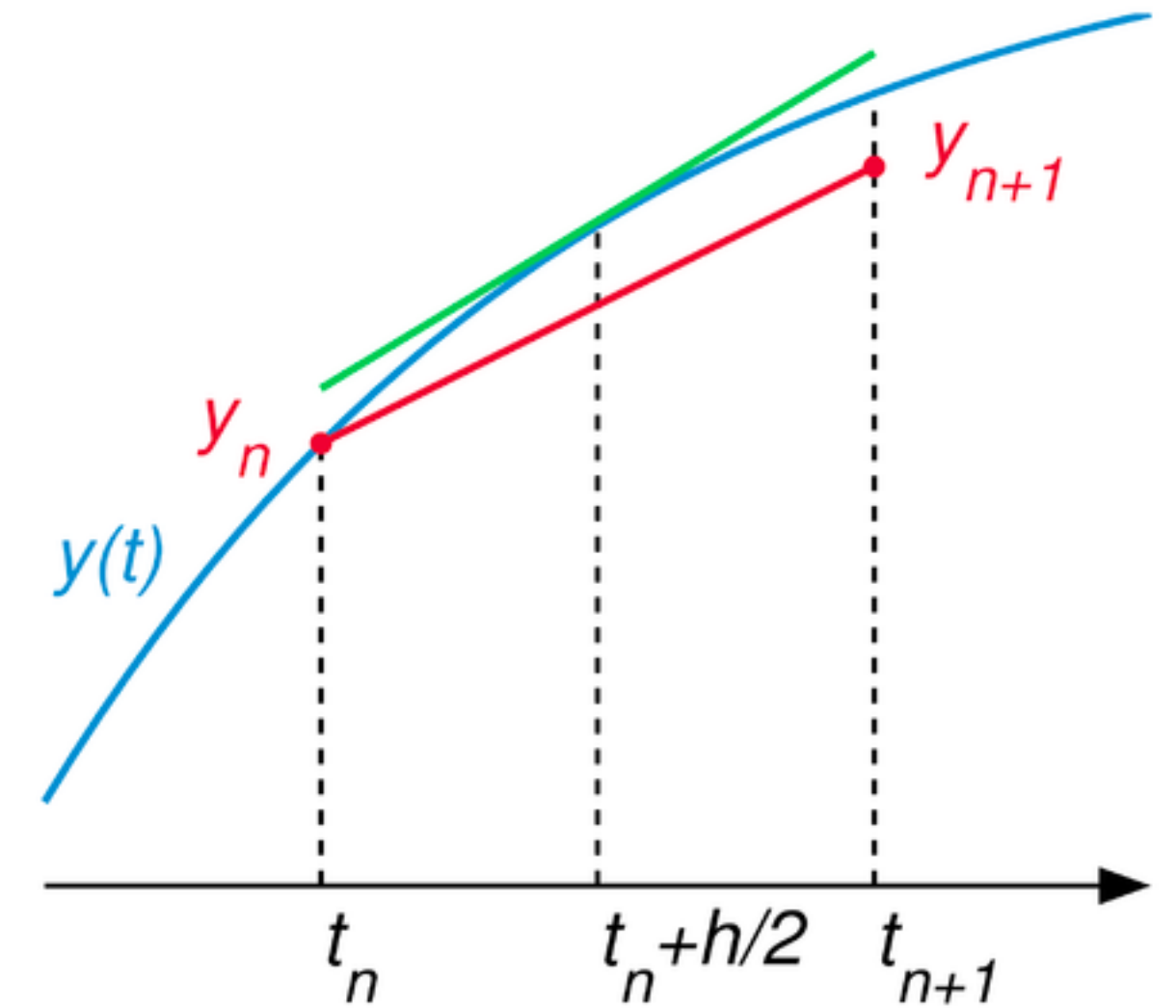
Midpoint

- Implicit midpoint method:

$$S_{n+1} = S_n + F \left(\frac{t_n + t_{n+1}}{2}, \frac{S_n + S_{n+1}}{2} \right) \Delta t$$

- 2nd-order accurate
- Time symmetric and **symplectic**
- But still implicit
- Explicit midpoint method

$$S_{n+1} = S_n + F \left(t_n + \Delta t/2, S_n + F(t_n, S_n)\Delta t/2 \right) \Delta t$$



Runge-Kutta Motivation

- Runge-Kutta (RK) methods are one of the most widely used methods for solving ODEs
- Euler method uses the first two terms in Taylor series to approximate the numerical integration

$$S(t_{n+1}) = S(t_n + \Delta t) = S(t_n) + \dot{S}(t_n)\Delta t$$

- We can improve the accuracy of the numerical integration if we keep more terms!

$$S(t_{n+1}) = S(t_n + \Delta t) = S(t_n) + \dot{S}(t_n)\Delta t + \frac{1}{2!}\ddot{S}(t_n)\Delta t^2 + \dots + \frac{1}{m!}\frac{d^m S}{dt^m}(t_n)\Delta t^m$$

- In order to get this more accurate solution, we need to derive expressions for the higher order derivatives

Runge-Kutta

- Runge-Kutta methods:
 - Whole class of integration methods

2nd-order accurate $\mathcal{O}(\Delta t^2)$

$$k_1 = F(t_n, S_n)$$

$$k_2 = F(t_n + \Delta t, S_n + k_1 \Delta t)$$

$$S_{n+1} = S_n + \left(\frac{k_1 + k_2}{2} \right) \Delta t$$

4th-order accurate $\mathcal{O}(\Delta t^4)$

$$k_1 = F(t_n, S_n)$$

$$k_2 = F(t_n + \Delta t/2, S_n + k_1 \Delta t/2)$$

$$k_3 = F(t_n + \Delta t/2, S_n + k_2 \Delta t/2)$$

$$k_4 = F(t_n + \Delta t, S_n + k_3 \Delta t/2)$$

$$S_{n+1} = S_n + \left(\frac{k_1}{6} + \frac{k_2}{3} + \frac{k_3}{3} + \frac{k_4}{6} \right) \Delta t$$