# PHYS 142/242 Lecture 23: Preview of PHYS 139/239

Javier Duarte – March 6, 2024





## **Course Evaluations!**

- https://academicaffairs.ucsd.edu/Modules/Evals?e11230304
- Available until Wednesday, March 27 at 8:00 AM

• If we reach above 80% submitted evaluations, extra credit on final projects

### **Final Presentations**

- 1 person per group fill out the form ASAP: <u>https://forms.gle/</u> RVSJ2RNAoAL83n286
- Discuss amongst yourselves then fill it out by Friday lecture
- Limited to 15+5 minutes
- Example presentations from previous years: <u>https://docs.google.com/</u> presentation/d/1jklBqTkAAXtWCO88Lm4ulzRwCAi8nwUU/edit? <u>usp=sharing&ouid=117156458781940750638&rtpof=true&sd=true</u>
- Presentations will take place Tuesday Friday in Lab/Lecture
- forms.gle/9w36yFgoB8PqwZPE9

All groups are expected to be present to give feedback via form: <u>https://</u>

### **Final Presentations Schedule**

Tuesday Lab	Wednesday Lecture	Thursday Lab	Friday Lecture
Double Well Potential	Double Well Potential	Drell-Yan Production	2D Ising Model with
with MCMC (B)	with MCMC (A)	with VEGAS (C)	MCMC (B)
Drell-Yan Production	Drell-Yan Production	2D Ising Model with	Drell-Yan Production
with VEGAS (A)	with VEGAS (B)	MCMC (C)	with VEGAS (D)
2D Ising Model with MCMC (A)		Double Well Potential with MCMC (C)	

# **Final Report**

- Template: <u>https://www.overleaf.com/</u> read/drwctbrvmzfs#705aa6
- Due Friday of Finals Week

### Manuscript Title: Subtitle

Ann Author<sup>\*</sup> and Second Author<sup>†</sup> University of California San Diego

(Group: Double Well Potential with MCMC (A)) (Dated: March 6, 2024)

An article usually includes an abstract, a concise summary of the work covered at length in the main body of the article.

### I. INTRODUCTION

Introduce the problem you are solving. Discuss the physics behind the project and introduce the computational methods you will use. Also, mention and cite any papers you use [1].

Describe the main objective of your project.

### **II. METHODS**

Describe methods, e.g. Markov chain Monte Carlo with Metropolis-Hastings algorithm. Here's an example of an equation for the path integral

$$K(x_b, t_b; x_a, t_a) = \int \mathcal{D}x(t) \exp\left[\frac{i}{\hbar} \int_{t_a}^{t_b} L(x(t)) dt\right] \quad (1)$$

Make sure you define all variables in any equations you write!

Describe and discuss any parameters you choose for your computational method, e.g. burn-in steps, etc.

Provide the link to your software in GitHub repository [2].

### III. RESULTS

Report the results of your simulations. Add figures showing your results, as in Fig. 1.

Discuss significance of results. In particular answer questions posed in the assignment, e.g. explain the connection to statistical mechanics.

### IV. CONCLUSION

Brief summary of the project and results. Describe any lessons learned or possible future work.

### V. CONTRIBUTIONS

Briefly describe contributions from each team member.



FIG. 1. Describe your figure in full.

### ACKNOWLEDGMENTS

Add any acknowledgments (optional).

- [1] S. Mittal, M. J. Westbroek, P. R. King, and D. D. Vvedensky, Path integral Monte Carlo method for the quantum
- anharmonic oscillator, Eur. J. Phys. 41, 055401 (2020). [2] J. M. Duarte, UCSD PHYS 142 GitHub (2024).

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### **Self & Peer Evaluations**

- Form: <u>https://forms.gle/</u> P9C7E9jYrmn4hEHZ7
- Due Friday of Finals Week

### Self & Peer Evaluation for PHYS 141 Midterm/Final Group Project

Please assess the work of you and your colleagues by using the following criteria. We will consider your feedback in assigning the grade for the project. Please try to be as honest and fair

as possible in your assessment.

5 = Excellent work; was crucial component to group's success

- 4 = Very strong work; contributed significantly to group
- 3 = Sufficient effort; contributed adequately to group
- 2 = Insufficient effort; met minimal standards of group
- 1 = Little or weak effort; was detrimental to group

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\* Indicates required question

Email \*

Your email

UCSD PID \*

Your answer

 $\odot$ 

# PHYS 143/243: Machine Learning in Physics

- Study of machine learning methods applied to physics
- Currently offered as a 139/239 Special Topics
- Overview:
  - Supervised learning
  - (Boosted) decision trees tabular data
  - (Deep) neural networks tabular data
  - Convolutional neural networks image-like data
  - Graph neural networks graph-like data and point clouds
  - Unsupervised learning
  - (Variational) autoencoders for anomaly detection
  - Model compression
  - Special topics via guest lectures (TBD)
    - Equivariant models
    - Generative models
    - Reinforcement learning
    - Explainability  $\bullet$
    - Uncertainty
  - Final team projects



# What is machine learning?

Science and art of learning automatically from data and experience



- Large overlap with data mining:
  - ML focuses on algorithms, DM on discovering patterns



- Example 1: Predict stellar radius given stellar mass



- Example 2: Classify images of neutrino interactions

# • Learn a function $f: X \to Y$ from an input space X (observations) to an output $P_{OP Publishing}$



- signal



• Learn a function  $f: X \to Y$  from an input space X (observations) to an output space Y (targets), using a set of labeled examples  $(x_1, y_1), (x_2, y_2), \dots, (x_N, y_N)$ .





arXiv:2101.08578

• Learn a function  $f: X \to Y$  from an input space X (observations) to an output space Y (targets), using a set of labeled examples  $(x_1, y_1), (x_2, y_2), \dots, (x_N, y_N)$ .

• Example 4: Estimate particle momentum, charge, type, etc. from detector hits

## Linear models: workhorse of machine learning

- Linear models on top of good features can yield excellent results
- More complex model classes (e.g., as their basic building block



Neural network: linear model after inputs are mapped to features through a nonlinear transformation

 $f(x | w_1, w_2) = w_2^{\mathsf{T}} \sigma(w_1^{\mathsf{T}} x)$ 



### Neural networks

input features weights non-S<sup>ums</sup> linearities  $x_1$  $x_2$  $x_M$ 

**network**: sequence of parallelized weighted sums and non-linearities  $\sigma($  $= \sigma( \cdots \sigma($ ••• ) 2nd weights 1st weights input output





### Symmetries

### Invariance

 $f(\rho_g(x)) = f(x)$ 



### **Equivariance** $f(\rho_g(x)) = \rho'_g(f(x))$



### **Translational invariance**

- For the purpose of classifying galaxy morphologies (e.g. spiral), the answer shouldn't depend on the absolute location of the pixels
- For simplicity, imagine there are 4 possible locations the galaxy might show up (top left, top right, bottom left, and bottom right)



### **Convolutional neural networks**

What if the same fully-connected neural network is applied to each corner?



## Graph neural networks



- Convolutional: sender node features are multiplied with a constant
- receiver over the sender
- sender and receiver

### arXiv:104.13478

Attentional: multiplier is implicitly computed via an attention mechanism of the

Message-passing: vector-based messages are computed based on both the



